Magnetic dipole excitations in nuclei: elementary modes of nucleonic motion

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The nucleus is one of the most multi-faceted many-body systems in the universe. It exhibits a multitude of responses depending on the way one 'probes' it. With increasing technical advancements of beams at the various accelerators and of detection systems the nucleus has, over and over again, surprised us by expressing always new ways of 'organized' structures and layers of complexity. Nuclear magnetism is one of those fascinating faces of the atomic nucleus we discuss in the present review. We shall not just limit ourselves to presenting the by now very large data set that has been obtained in the last two decades using various probes, electromagnetic and hadronic alike and that presents ample evidence for a low-lying orbital scissors mode around 3 MeV, albeit fragmented over an energy interval of the order of 1.5 MeV, and higher-lying spin-flip strength in the energy region 5-9 MeV in deformed nuclei, nor to the presently discovered evidence for low-lying proton-neutron isovector quadrupole excitations in spherical nuclei. To the contrary, we put the experimental evidence in the perspectives of understanding the atomic nucleus and its various structures of well-organized modes of motion and thus enlarge our discussion to more general fermion and bosonic many-body systems.

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I. INTRODUCTION

A. General remarks

Nucleons moving inside the atomic nucleus naturally generate orbital and spin magnetism. In certain mass regions - in particular for nuclei between closed shells the orbital magnetism can give rise to cooperative effects between the many nucleons in the nucleus. Collective modes might result from the out-of-phase motion of protons and neutrons and of those, magnetic dipole modes at fairly low energies ($0\hbar\omega$ excitations) excited with electromagnetic probes are one of the most pronounced ones. Besides these, at higher excitation energies, cooperative effects may even lead to collective spin-flip modes, as well as to even higher-lying genuine collective dipole modes $(2\hbar\omega)$ excitations which so far have not even been seen directly in experiments. In the present article, we start from a succinct discussion of the early attempts in order to describe nucleonic out-of-phase motion leading to magnetic collective excitations (Sec. II) before entering into a discussion of magnetic dipole excitations in heavy nuclei (Sec. III). The experimental evidence that has been accumulated over the years concerning the observation of a so-called scissors mode in which the neutrons and protons in a deformed nucleus perform small angle vibrations in a scissors-like motion with respect to each other, using both electromagnetic and hadronic probes, is listed. Theoretical concepts concerning the description of the low-lying orbital scissors strength in eveneven nuclei are presented. Collective (geometric and algebraic) and microscopic (shell-model and Quasi-Particle Random Phase Approximation (QRPA) studies) models are discussed and they are related to one another in order to better understand both the complementarity and the specific model effects. The difficult problem of addressing the observed fragmentation of orbital magnetic strength will be looked at also in the light of collective and microscopic approaches.

The aspects related to the experimental evidence and the derivation of a theoretical description of the concentration of spin-flip strength at higher excitation energies is presented in Sec. III as well, using mainly QRPA and shell-model calculations. In Sec. IV magnetic dipole excitations in heavy odd-mass nuclei are discussed. There the problem of missing strength in the measured spectra is looked at in some detail and the need for a better theoretical description of the fragmented transition strength is pointed out. Section V deals with experimental examples from magnetic dipole excitations in light and medium-heavy nuclei and their theoretical treatment in terms of the shell model and the QRPA. In Sec. VI, we discuss the magnetic dipole isovector transitions in vibrational nuclei, illuminating the intimate connection with the scissors mode typical to rotational nuclei. In Sec. VII, we bring the former discussion within a broader context of general many-body systems e.g. deformed metallic clusters, quantum dots and scissors motion in trapped Bose-Einstein condensates. In the final Sec. VIII, conclusions and an outlook are given concerning the issue of magnetic dipole excitations within the broader context of past and future nuclear physics research. Relevant literature on the subject that appeared until the end of 2009 has been considered in this review.

B. Magnetic dipole response in atomic nuclei: a qualitative overview

Even on rather general grounds, one can make a fingerprint figure for the magnetic dipole response in heavy rare-earth deformed nuclei on one side and more spherical and light nuclei on the other side. We illustrate the salient features of this response in two figures that shall be referred to quite often in the present review.

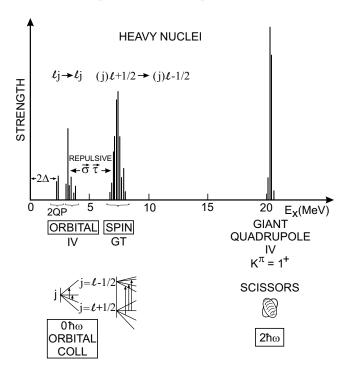


FIG. 1 Schematic illustration of the magnetic dipole strength distribution in even-even heavy deformed nuclei and its model character (Richter, 1990).

In even-even strongly-deformed (rare-earth and actinide) nuclei, particularly due to the lifting of the spher-

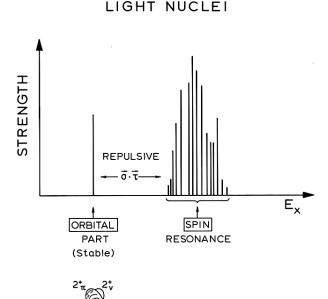
ical symmetry and the associated degeneracy in the various m-components for a spherical orbital with angular momentum j, one can separate four different energies regions (see Fig. 1):

(i) At the excitation energy of about twice the pairing gap $(2\Delta \simeq 2-2.5 \, \mathrm{MeV})$, two quasi-particle $(2\mathrm{qp}) \, J^\pi = 1^+ \, \mathrm{excitations}$ show up with a very specific shell-model structure, and thus, if these stay rather pure, can be detected in electromagnetic decay and selective transfer reactions. (ii) At the excitation energy around 3 MeV, one observes a concentration of orbital magnetic dipole strength, built up from various 2qp configurations, into a weakly collective $0\hbar\omega$ mode, called the scissors mode. Here, a number of proton and neutron 2qp configurations $(lj) \to (lj)$ contribute in a more or less coherent way, depending on external quantities like nuclear deformation and the position of the Fermi level (number of protons Z and the number of neutrons N) in the Nilsson deformed single-particle spectrum.

(iii) In the excitation energy interval of 6-8 MeV, one starts observing the spin Gamow-Teller giant resonance (with a bound part, depending on the precise location in energy, and a resonance part) resulting from particle-hole (p-h) excitations across the major closed shells. In particular, shell-model transitions of the type $(j=l+\frac{1}{2}) o (j=l-\frac{1}{2})$ play a major role. In the rare-earth region, depending on the precise proton and neutron number, the $1g_{9/2} \rightarrow 1g_{7/2}$, $1h_{11/2} \rightarrow 1h_{9/2}$, and $1i_{13/2} \rightarrow 1i_{11/2}$ transitions, respectively, contribute most to the 1⁺ spin mode. Moreover, the residual $\vec{\sigma} \cdot \vec{\sigma} \vec{\tau} \cdot \vec{\tau}$ repulsive part of the effective nucleon-nucleon interaction concentrates the spin strength from the lower-lying 2qp states around 2-4 MeV into this state. The final result is a large concentration of spin M1 strength. (iv) Still higher in excitation energy, near 20 MeV, the $K^{\pi}=1^{+}$ component of the isovector giant quadrupole resonance should eventually show up. This particular mode, built in a microscopic way from a coherent superposition of $2\hbar\omega$ configurations, has originally been studied in macroscopic collective models and this state would correspond to the 'real scissors mode' in strongly deformed, rotational nuclei. Unfortunately, due to the very high excitation energy and due to the fact that its transition strength vanishes at the photon point because the M1operator contains no radial dependence, such a state in the continuum is very difficult to observe unambiguously.

The excitation mechanisms described in (i)-(iv) should thus lead to the typical fingerprint pattern of the magnetic dipole response in deformed rare-earth nuclei, in the actinides and partly also in the medium-heavy deformed nuclei in the fp-shell region as illustrated schematically in Fig. 1. In the following discussion, we concentrate on these various patterns, and give a thorough but succinct discussion of the experimental facts validating this former, more 'idealized' picture, emerging from the basic underlying shell structure in deformed nuclei in combination with the essential multipole components of the residual two-body interactions: the pairing component,

the repulsive $\vec{\sigma} \cdot \vec{\sigma} \vec{\tau} \cdot \vec{\tau}$ component, and the long-range quadrupole force.



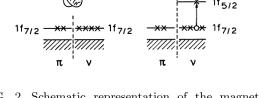


FIG. 2 Schematic representation of the magnetic dipole strength distribution in even-even light nuclei and its interpretation in terms of the shell model. The subscripts π and ν denotes the proton and neutron contribution, respectively (Richter, 1990).

In the spherical, lighter fp-shell nuclei, but also in the region of lighter rare-earth nuclei (nuclei with neutron number N near to 82, and proton number Z between 50 to 66), a somewhat different and slightly more simple structure emerges (see Fig. 2):

- (i) On the lower excitation energy side, one observes a rather stable (in energy) predominantly orbital 1^+ excitation which is produced through coupling the lowest proton 2_{π}^+ excitation with the lowest neutron 2_{ν}^+ excitation. Depending on the precise shell structure, one can find also some higher-lying $4_{\pi}^+ \otimes 4_{\nu}^+$, $6_{\pi}^+ \otimes 6_{\nu}^+$ fragments, in general, with rapidly decreasing M1 excitation strength. This feature is particularly evident in the light nuclei with protons and neutrons filling the spherical $1f_{7/2}$ shell-model orbital.
- (ii) Removed from this lower state by several MeV and near to 10 MeV of excitation energy for the nuclei in the mass A=50 region, one finds the spin Gamow-Teller giant resonance. Here, the repulsive spin-isospin force component is responsible for a coherent state made of 1p-1h excitations, mainly. For nuclei around mass A=50, this will predominantly be a $1f_{7/2} \rightarrow 1f_{5/2}$ excitation, whereas in the somewhat heavier mass A=90

region, one will encounter $1g_{9/2} \rightarrow 1g_{7/2}$ transitions.

The particular magnetic dipole response shown schematically in Fig. 2 forms the salient features amply discussed in the present article when highlighting $J^{\pi} = 1^{+}$ states and the associated M1 strength for light and mediumheavy even-even nuclei. Besides the extensive exploration of the magnetic dipole response in deformed nuclei and also in the region of light to medium-heavy nuclei, uncovering the presence of a scissors mode (Sec. I.B), during the last decades, much progress has been made lately in the study of proton-neutron 2⁺ excited states in vibrational and transitional nuclei, corresponding with mixed-symmetry wave functions in the proton and neutron building blocks. Here, small amplitude quadrupole oscillations (phonons) dominate the low-energy nuclear structure properties. Coupling the proton 2^+_{π} and neutron 2^+_{ν} phonons can result in multiphonon states where the phonons move in phase (characterized by symmetric wave functions with shorthand notation S), but also with out-of-phase motion (with mixed-symmetry wave functions and shorthand notation MS). This is illustrated in Fig. 3. Recent review articles (Kneissl et al., 2006; Pietralla et al., 2008) extensively discuss the regions in which such 2^+ states have been observed as well as present the experimental techniques needed to characterize unambiguously those isovector excitations (photon scattering, electron scattering, Coulomb excitation, β decay, inelastic neutron scattering and light-ion fusion reactions as the major probes).

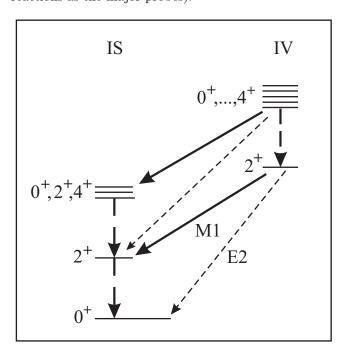


FIG. 3 Schematic representation of the M1 and associated E2 transitions of states with symmetric (l.h.s.) and mixed-symmetric (r.h.s.) wave functions in vibrational nuclei near closed shells.

In view of the isoscalar and isovector components of the magnetic dipole operator (see also Secs. III.B.1 and VI for its precise structure and a more detailed discussion), the experimental identification of these isovector excitations, shown schematically in Fig. 3, is characterized by

- (i) strong M1 transitions (matrix element of $\simeq 1\mu_N$) between the MS and S states (thick full lines),
- (ii) weakly-collective E2 transitions (a few % of the large isoscalar $B(E2; 2_1^+ \to 0_1^+)$ transition probability, typical for vibrational and transitional nuclei) between the MS and S states (thin dashed lines), and
- (iii) strongly collective E2 transitions in between states with MS character, with a strength typical for the collective E2 transitions between symmetric states (thick dashed lines).

These particular fingerprints properties result from a systematic exploration of isovector excitations in vibrational and transitional regions during the last decade (Pietralla et al., 2008), and clear-cut evidence for the observation of the lowest 2^+_{ms} (both of a one-phonon and two-phonon nature) as well as two-phonon 3^+_{ms} and 1^+_{ms} isovector excitations exists now for the mass regions with $Z\sim 40, N\sim 50$ and $Z\sim 50, 50 < N < 82$, rare-earth nuclei in the region $54 < Z \le 82, 72 < N \le 82$ (i.e., Xe,Ba,Ce,Nd and Sm nuclei), as well as nuclei in the $A\sim 60$ mass region, and heavy nuclei near $^{208}{\rm Pb}$.

C. Scissors modes in other many-body systems

In the introductory part to this review, we have pointed out that the magnetic dipole response in strongly deformed atomic nuclei, is characterized by a clearly separated orbital scissors mode (small-angle vibration of neutrons versus protons) at the lower energies and at the higher energy side by a resonance-like structure comprised of proton and neutron spin-flip excitations. In nuclei with vibrational and transitional character, near to closed shells, mixed-symmetry excitations result from the isovector coupling of the lowest one-phonon proton and neutron 2_1^+ excitations. The essential ingredient in both cases is the presence of two distinguishable components, proton and neutron fluids, in the atomic nucleus.

It turns out that in other many-body systems very similar rotational oscillatory scissors motion has been discussed and, in certain cases, also been observed experimentally. The presence of a two-component or twofluid quantum system is essential in this respect. In Sec. VII, we address the main results obtained in the study of (i) metallic clusters (Lipparini and Stringari, 1989a; Nesterenko et al., 1999), (ii) elliptically deformed quantum dots (Alhassid, 2000; Serra et al., 1999), and (iii) the oscillatory behavior induced by the rotation of the atomic cloud in a deformed trap and the corresponding superfluid effects caused by Bose-Einstein condensation (Guéry-Odelin and Stringari, 1999; Maragò et al., 2001, 2000), as well as their connections with the study of magnetic dipole excitations in atomic nuclei. The underlying common mechanism as well as some typical illustrations will be presented too.

II. EARLY THEORETICAL SUGGESTIONS FOR THE MAGNETIC SCISSORS MODE AND ITS EXPERIMENTAL DISCOVERY

A. Overview: a piece of history

Low-energy collective modes in atomic nuclei, for both spherical and deformed nuclei, displaying nuclear density oscillations or more permanently deformed structures of the density, have been rather well described within the Bohr-Mottelson model (Bohr and Mottelson, 1975). In those excitations, both the proton density $\rho_{\pi}(\vec{r})$ and the neutron density $\rho_{\nu}(\vec{r})$ exhibit variations that act in phase, i.e. isoscalar collective modes are obtained. It was rather soon realized that, besides these symmetric collective modes, non-symmetric density variations might show up too. The latter excitations are expected to occur at much higher excitation energies though because of the symmetry energy term, coupling the proton and neutron density oscillations preferentially in a symmetric way. The giant electric-dipole mode - which in even-even nuclei excites negative parity states - is the best known and welldocumented example for such isovector excitations, the mode in which the centre-of-mass for the charge and mass distributions do not coincide but perform an out-of-phase motion around the equilibrium value.

Non-symmetric collective modes were considered quite early by Greiner (1965, 1966) and Faessler (1966), independently, at the end of the 60's for spherical nuclei and this on the basis of isovector quadrupole collective excitations. Somewhat later but again, about at the same time and independently, Hilton (1976), Suzuki and Rowe (1977), and Lo Iudice and Palumbo (1978, 1979) suggested an extension of the Bohr-Mottelson description of rotational motion. The latter treated the nucleus in terms of a geometrical two-rotor model (TRM), in which a collective magnetic dipole (M1) mode could be formed by a rotational oscillation of the proton versus the neutron deformed density distribution (or fluids giving a rotational flow that is strongly excited as an orbital magnetic excitation). The name 'scissors mode' was suggested much later after the experimental discovery of this mode originating in the peculiar nature of its geometrical picture (Bohle et al., 1984b; Richter, 1983).

At that time, these non-symmetric excitations, in particular the scissors mode in deformed systems, were a mere theoretical suggestion built on the strong fundament in describing nuclear collective motion (both for spherical and deformed nuclei). A determining factor in order to estimate the excitation energy and the strength was of course the knowledge of the symmetry energy connected to the non-symmetric motion. In the early calculations, the full symmetry energy term, known from a liquid-drop model treatment of global nuclear structure properties, was considered giving rise to energies and B(M1) values much too large. It was only when the ex-

citation energy of the scissors mode (having both a mass parameter and a restoring force strength) was adjusted to the observed experimental low-lying B(E2) values including also pairing correlation in the protons and neutrons participating in the scissors motion that excitation energies of about 3-5 MeV (De Franceschi et~al., 1983, 1984; Lipparini and Stringari, 1983; Suzuki and Rowe, 1977) but still fairly high B(M1) values of $9-18~\mu_N^2$ were obtained (De Franceschi et~al., 1983, 1984; Lipparini and Stringari, 1983).

Strong support for the above idea came from a different way of treating collective modes of motion in the nuclear many-body problem. Working in an algebraic framework and using the concepts of symmetry, Arima and Iachello (1975a,b) formulated a model in which the interacting fermion problem is replaced by an interacting boson problem, only considering s (L=0) and d (L=2) boson degrees of freedom (interacting boson model, IBM). It was known from standard shell-model two-body interactions that the 0⁺ coupled pair state and, to a much lesser extent, the 2⁺ coupled pair dominate the binding of two-nucleon systems. By studying the symmetries of such an interacting boson model with the U(6) symmetry (the symmetry describing an interacting system of s and d bosons) and at the same time incorporating the proton and neutron degrees of freedom (IBM-2), a class of states with non-symmetric spatial and also non-symmetric charge structure (mixed-symmetry states) showed up naturally. By fitting the parameters in the IBM-2 model to known fermionic properties, a B(M1) value for a scissors like mode of about $2-3 \mu_N^2$ has been predicted (Iachello, 1981).

B. Experimental discovery

Both the predictions of Lo Iudice and Palumbo (1978, 1979) within the TRM and of Iachello (1981) and Dieperink (1983) within the IBM-2 formed the essential basis for high-resolution inelastic electron scattering experiments at the Darmstadt Electron Linear Accelerator (DALINAC) to search for this mode. And indeed after great experimental efforts it has first been detected in the strongly deformed nucleus ¹⁵⁶Gd with essentially all the properties that were predicted (Richter, 1983). Figure 4 from the work of Bohle et al. (1984b) displays two of the original spectra measured at a backward angle of $\theta = 165^{\circ}$ where magnetic excitations are expected to show up strongly. The spectrum, taken at the low bombarding energy $E_0 = 30 \text{ MeV}$ (corresponding to a low momentum transfer), reveals a rich fine structure of excited low-spin states with known experimental levels up to about 2.5 MeV of excitation energy. The only strong transition is to a state at $E_x = 3.075$ MeV. This state is almost absent in the $E_0 = 50$ MeV spectrum (i.e. at a higher momentum transfer), in which, however, a well known collective $J^{\pi}=3^{-}$ state at $E_{x}=1.852~\mathrm{MeV}$ is strongest. The state at $E_x = 3.075$ MeV dominates all

measured spectra at low momentum transfer, has a form factor behaviour consistent with a $J^{\pi}=1^+$ assignment and a transition strength B(M1) \uparrow of about 1.5 μ_N^2 .

Immediately after its discovery in ¹⁵⁶Gd it has been shown that the newly found *M*1 mode of course is not unique to this particular nucleus but is a general property of heavy deformed nuclei. This has been proven by inelastic electron scattering measurements at the DALINAC on the deformed nuclei ¹⁵⁴Sm, ¹⁵⁸Gd, ¹⁶⁴Dy, ¹⁶⁸Er and ¹⁷⁴Yb (Bohle *et al.*, 1984a), where also first evidence for the fragmentation of the scissors mode strength has been presented in the nucleus in which it has been discovered, ¹⁵⁶Gd. Furthermore, the new mode has also immediately been verified in nuclear resonance fluorescence experiments with real photons (Berg, 1984; Berg *et al.*, 1984a,b).

Since then, the experimental evidence for such scissorslike excitations in strongly deformed nuclei but also in vibrational, transitional and gamma-soft nuclei has become compellingly large. Moreover, the scissors mode has been explored besides the purely electromagnetic probes (electrons, photons) also with hadronic probes (proton scattering and low-energy neutron-induced compound reactions) all over the nuclear mass table. A rather complete summary of the experimental data from electron and photon scattering can be found in Enders et al. (2005). Several articles covering this work both from the experimental and the theoretical point of view appeared in the past (Berg and Kneissl, 1987; Heyde, 1989; Kneissl et al., 1996; Lo Iudice, 1997, 2000; Richter, 1985, 1990, 1991, 1993, 1994, 1995; Zawischa, 1998). In the next section (Sec. III.A.1), we will discuss this experimental evidence for magnetic excitation of states of mixed-symmetric character (spatially and in the proton-neutron charge coordinates) and present all of the salient features in depth

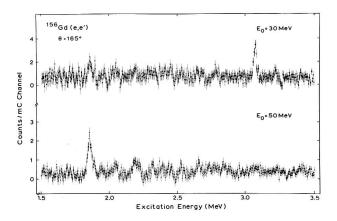


FIG. 4 Backward angle $^{156}{\rm Gd}(e,e')$ spectra indicating the almost uniform excitation of many-low spin states in the measured energy region except for a strongly excited $J^{\pi}=1^+$ state seen in the 30 MeV spectrum at $E_x=3.075$ MeV and a known $J^{\pi}=3^-$ state at $E_x=1.825$ MeV in the 50 MeV spectrum. The excitation of the state at $E_x=3.075$ MeV is due to the scissors mode (Bohle et~al., 1984b).

although not exhaustively. We put the accent mainly on the overall and systematic features rather than on each individual case, separately. Thereby, we are able to focus on the essential properties of this mode of motion.

III. MAGNETIC DIPOLE EXCITATIONS IN HEAVY NUCLEI

A. Low-energy scissors mode

1. Experimental evidence

a. Overview The early systematics of the scissors mode in heavy deformed nuclei known mainly from highresolution electron and photoexcitation experiments at the DALINAC and the S-DALINAC in Darmstadt and the DYNAMITRON accelerator in Stuttgart has been described by Richter (1995) and Kneissl et al. (1996). The majority of nuclei studied are even-even ones but the scissors mode has also been detected in a number of heavy odd-mass nuclei (Sec. IV). Very often the transition strength of the scissors mode is fragmented. In order to detect weak transitions, highly efficient gammaray detector systems with proper background suppression have to be used in photon scattering (also called nuclear resonance fluorescence, NRF) experiments as comprehensively discussed by Kneissl et al. (1996). In presentday nuclear resonance fluorescence experiments reduced transition strengths of $B(M1) \uparrow \approx 0.01 \mu_N^2$ for magnetic dipole transitions at an energy of about 3 MeV can be detected. The scissors mode has been studied in vibrational and rotational nuclei, in chains of isotopic nuclei. Advances in gamma spectroscopy - e.g. by the use of EUROBALL detector modules (von Neumann-Cosel, 1997) - have extended our knowledge of the scissors mode to γ -soft nuclei, ^{194,196}Pt (von Brentano et al., 1996; Linnemann et al., 2003) and the chain of Xe isotopes (von Garrel et al., 2006). For example, in the rareearth region a total of 42 isotopes ranging from Nd to Pt have been studied providing detailed experimental information on the systematics of the scissors mode. Furthermore, it has been verified in the actinide mass region (Heil et al., 1988; Margraf et al., 1990).

A comparison of different probes for the case of $^{156}\mathrm{Gd}$ is made in Fig. 5 (Richter, 1990). A combined analysis of the (e,e') and (γ,γ') experiments (Bohle et al., 1986) revealed five more weakly excited $J^{\pi}=1^+$ states in the vicinity of the strongest one, bringing the total M1 strength up to about $2.4~\mu_N^2$. The strongest state might be viewed acting as a doorway for the others. A high-resolution inelastic proton scattering spectrum (Wesselborg et al., 1986) is also shown in Fig. 5 and the absence of any of the states seen in the (e,e') and (γ,γ') reactions is already a strong hint that the scissors mode is excited through the orbital part of the magnetic dipole operator.

The salient features of the scissors mode in heavy deformed nuclei unraveled in high-resolution electron, pho-

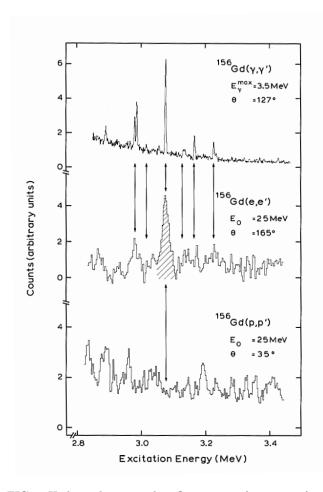


FIG. 5 High-resolution nuclear fluorescence (upper part) and inelastic electron scattering (middle part) spectra showing a strongly excited $J^{\pi}=1^+$ state (hatched) and several weaker 1^+ states, all marked by arrows. These states are conspicuously absent in the inelastic proton scattering (lower part) spectrum (Richter, 1990).

ton and proton scattering experiments (Richter, 1995) are the following:

(i) Its mean excitation energy scales roughly as $66\delta A^{-1/3}$ MeV with δ being the deformation parameter.¹ This puts the center of gravity of the orbital M1 strength distribution in rare earth nuclei at $E_x \approx 3$ MeV. (ii) The total transition strength from the ground state into the $J^{\pi}=1^+$ states is $\sum B(M1) \uparrow \approx 3 \ \mu_N^2$ and the maximum strength that is carried in the transition to an individual state is roughly 1.5 μ_N^2 .

(iii) In the nuclear transition current the orbital part dominates over the spin part and one has typically $B_l(M1)/B_{\sigma}(M1) \approx 10.$

(iv) The summed experimental transition strength up to $E_x \approx 4$ MeV varies quadratically with the quadrupole ground state deformation parameter.

Before these observations are compared with various nuclear model predictions in Sec. III.A.2 we concentrate briefly on a few more experimental characteristics of the mode.

b. Form factor One of the first strong hints that indeed $J^{\pi} = 1^{+}$ states were excited from the $J^{\pi} = 0^{+}$ ground state came from the measurements of form factors in inelastic electron scattering at low momentum transfers (Bohle et al., 1984a,b). In all cases a shape characteristic for a magnetic dipole form factor has been found. As an example, one such form factor (Bohle et al., 1987a) is shown in Fig. 6 for the transition into a scissors mode state in 164 Dy at $E_x = 3.11$ MeV. It is compared to a form factor calculated (Scholtz et al., 1989) in distorted wave Born approximation (DWBA) with a QRPA transition density to a state at $E_x = 3.128$ MeV, i.e. very close to the experimental state. The full curve is the total form factor (orbital plus spin) and the dashed one its orbital part alone. It is evident that the experimental form factor is well described at and above its first maximum by the model, considering the sizable uncertainty due to the smallness of the form factor values at momentum transfers larger than 1 fm^{-1} . Similarly to the QRPA model, the Interacting Boson Model-2 (IBM-2)- to be discussed later on - describes the data as well and points in particular also to the orbital nature of the transition.

c. Photon polarization and parity assignments Although the excitation of the scissors mode states by inelastic electron scattering and the measurement of the respective form factors has already provided a clear indication of the magnetic nature of the transition, it can be established beyond any doubt by using polarized photons either in the entrance or in the exit channel in NRF experiments (see Kneissl et al. (1996) for a detailed discussion). In the former method of $(\vec{\gamma}, \gamma')$, which has been successfully used at the photon scattering facilities at Gent (Govaert et al., 1994) and Rossendorf (Schwengner et al., 2007), the linearly polarized bremsstrahlung causes a positive azimuthal asymmetry for pure electric and a negative one for pure magnetic dipole excitations of the detected photon in the exit channel. In the latter method of $(\gamma, \vec{\gamma}')$, which has been employed vigorously in recent years at Stuttgart (Margraf et al., 1995), the parity assignments for the excited states in the nucleus result from the measurement of the linear polarization of the scattered photons with a Compton polarimeter. In a pioneering experiment (Kasten et al., 1989) in the field of the scissors mode this technique has been used to measure directly (and modelindependently) the parity of three strongly excited states

¹ In this article, the quadrupole deformation parameter used is mainly denoted by δ . Slightly different parameterizations $(\beta_2, \ \epsilon_2)$ exist, which can all be related with each other (Hasse and Myers, 1988; Löbner *et al.*, 1970). To lowest order $\epsilon_2 \approx \delta \approx \frac{3}{4} \sqrt{\frac{5}{\pi}} \beta_2$.

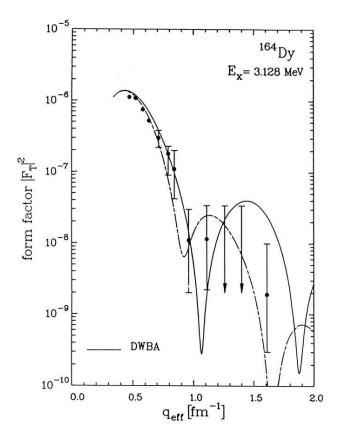


FIG. 6 Comparison of the experimental form factor for the transition into one of the $J^{\pi}=1^+$ scissors mode states in 164 Dy with form factors predicted in QRPA (Scholtz *et al.*, 1989). The dashed curve denotes the orbital part only (Richter, 1990).

near 3 MeV in the deformed nucleus 150 Nd. They were shown to be of positive parity, i.e. $J^{\pi} = 1^{+}$ states excited through the scissors mode (Fig. 7).

However, for transitions into states at excitation energies above 4 MeV, the application of this technique will be difficult, simply because the already small analyzing power of the Compton polarimeter at low energies becomes even smaller at higher energies. Here the use of polarized photons in the entrance channel of the reaction and the subsequent measurement of the intensity distribution of the scattered photon is preferable for the parity determination of nuclear dipole transitions. The analyzing power of this process is 100 % and independent of the energy of the scattered photon. In a conceptually simple experiment, Pietralla et al. (2001b) recently measured unambiguously the parity of a number of dipole excitations in ¹³⁸Ba between $E_x = 5.5 - 6.5$ MeV using photons from the Duke/OK-4 Storage Ring Free Electron Laser backscattered from relativistic electrons, now called the High Intensity Gamma-Ray Source (HI γ S) facility (Weller et al., 2009). The produced photon beam has been intense (10⁷ photons/s) and nearly monochromatic ($\Delta E_{\gamma}/E_{\gamma} \approx 3.8 \%$). This very efficient technique indeed will have a future in the study of elementary dipole

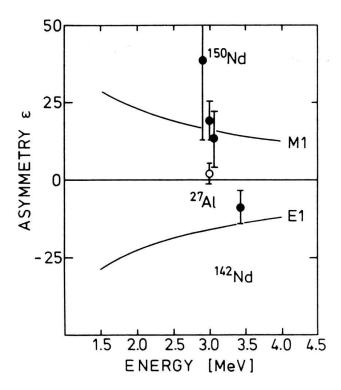


FIG. 7 Experimental asymmetries $\epsilon = (N_\perp - N_\parallel)/(N_\perp + N_\parallel)$ of Compton-scattered photons determined by a five-detector polarimeter. The experimental data are compared with the calculated polarization sensitivity of the setup given by the solid lines. Three strongly excited M1 transitions into states around 3 MeV excitation energy in the deformed nucleus $^{150}\mathrm{Nd}$ have been identified. For comparison an E1 transition into a state at 3.4 MeV in the spherical nucleus $^{142}\mathrm{Nd}$ is also shown. (Reprinted with permission from Kasten et~al. (1989). ©(2009) Am. Phys. Soc.)

transitions below threshold for particle emission.

d. Branching ratios of spin-one states in deformed nuclei and the K quantum number In NRF measurements, spin-one levels of both parities are selectively exited which decav either to the 0⁺ ground state or to low-lying excited states. Figure 8 displays two parts of a NRF spectrum taken at a bremsstrahlung endpoint energy slightly above 4.6 MeV for the strongly deformed ($\delta = 0.27$) nucleus ¹⁵⁴Sm. A number of transitions clustering around 3 MeV, i.e. the location of the states excited by the scissors mode, are observed (Ziegler et al., 1993). Multipolarities of individual transitions were ascertained by simultaneous two-point angular distribution measurements at 90° and 127° with respect to the direction of the incoming bremsstrahlung beam. Those data sufficed to clearly distinguish between quadrupole and dipole transitions. In the case of ¹⁵⁴Sm it has been possible to determine the nature of the dipole transitions (M1 or E1) by supplementing the present (γ, γ') data with (e,e') form factor measurements (Ziegler et al., 1993). However, there is

still another property attached to the J=1 levels excited from the $J^{\pi}=0^+$ ground state, i.e. its K quantum number. The pairs of the lines connected by brackets in the spectrum of Fig. 8 correspond to the decay of the J=1 levels into the ground state or the first excited $J^{\pi}=2^+$ state of $^{154}\mathrm{Sm}$. From the observed branching ratio one obtains information about the K quantum number of the J=1 state as demonstrated in Fig. 8. The

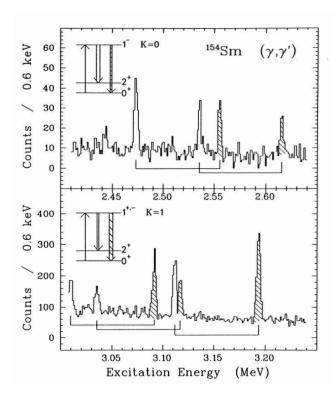


FIG. 8 Two parts of a NRF spectrum of 154 Sm. The ratio of the areas of peaks linked by brackets, corresponding to the ground state and the 2^+ transitions, allows the assignment of the K quantum number to the excited spin-1 state. In the upper part two examples for $J^{\pi}; K = 1^-; 0$, and in the lower part two examples for $J^{\pi}; K = 1^{+,-}; 1$ assignments are shown (Richter, 1991).

branching ratios for the deexcitation of levels in deformed nuclei to various states of a rotational band are governed by the so-called Alaga rules (Alaga and Mottelson, 1955) which yield for the ratio of transition strengths $B(1 \rightarrow 2)/B(1 \rightarrow 0) = 0.5$ for $\Delta K = 1$ and 2 for $\Delta K = 0$ transitions, respectively.

The examination of the decay of about $200 \ J=1$ levels in a number of rare earth nuclei (Zilges et al., 1990) indeed showed maxima at experimental branching ratios of 0.5 (as expected for states with K=1) and of 2.0 (as expected for states with K=0), respectively. Figure 9 presents an update including additional data which appeared since the original publication (Enders et al., 2005). The number of K=1 states is at least twice the number of K=0 (note that there are no positive parity states with K=0). Thus, the majority of the branching

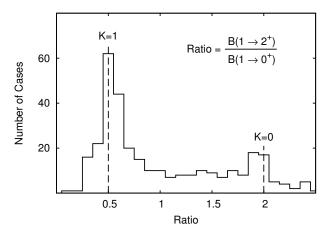


FIG. 9 Frequency distribution of experimental branching ratios for about 320 transitions observed in NRF in deformed rare-earth and actinide nuclei. One notices two distinct maxima at 0.5 for K=1 and at 2.0 for K=0 spin-one levels (adapted from Zilges *et al.* (1990)).

ratios are in good agreement with the Alaga rules. Furthermore, recent parity measurements of strong dipole transitions in 172,174 Yb confirm the E1/M1 assignments based on the K quantum numbers (Savran *et al.*, 2005).

But what is the reason for a number of cases with branching ratios in between the limits set by the Alaga rules? Some of them are known to result from rather strong E1 transitions of $J^{\pi}=1^{-}$ states with a strength hitherto still unexplained (Zilges et al., 1996). The deviations, however, may also be taken as evidence for possible K mixing. In fact K mixing matrix elements have been calculated from the measured energies, branching ratios and absolute transition strengths and are about 50 keV (von Brentano et al., 1994).

e. Evidence for quasi-particle excitations at low energy The high-resolution (e, e') data, obtained by Bohle et al. (1984a) as well as NRF results (Wesselborg et al., 1988) on ¹⁶⁴Dy have shown, in addition to the strongly excited $J^{\pi} = 1^{+}$ states around $E_{x} \approx 3$ MeV, a second group of $J^{\pi} = 1^{+}$ states about 0.5 MeV lower in excitation energy which carry a much weaker strength . Similar results have been obtained in the $^{168}\text{Er}(e,e')$ spectrum proving that the occurrence of a lower-lying weakly excited group of $J^{\pi} = 1^{+}$ states seems to be a general phenomenon in heavy deformed nuclei. On the basis of (e, e') form factor measurements, albeit difficult for the lower group of states because of the smallness of the experimental cross sections and the corresponding large uncertainty, it has been speculated upon that the two groups of states might be of very different structure (Richter, 1990).

That this is indeed the case has been proven independently in a study of the proton pick-up reaction $^{165}{\rm Ho}$ (\$\alpha\$,t) $^{164}{\rm Dy}$ (Freeman et~al., 1989). As can be seen from the top part of Fig. 10 the $J^{\pi}=3^+$ to 8^+ members

of a $K^{\pi} = 1^{+}$ rotational band have been identified in this single-particle transfer reaction. The reconstruction of the (not populated) 1⁺ band head energy resulted in $E_x = 2.543(13)$ MeV which can be safely identified with the $J^{\pi}=1^{+}$ state energy of $E_{x}=2.539~\mathrm{MeV}$ observed in (e, e') scattering (Bohle et al., 1987a) and the NRF experiments (Wesselborg et al., 1988). This 1⁺ state thus corresponds most likely to a two-quasiproton configuration (bottom part of Fig. 10) of the form $\frac{7}{2}$ [523] $\otimes \frac{5}{2}$ [523], and such a configuration is of course not consistent with a collective magnetic dipole transition from the ground state into it, as has been pointed out earlier by Hamamoto and Åberg (1984) and later also by Otsuka (1990). Furthermore, no rotational states built upon the $J^{\pi} = 1^+$ scissors mode states around $E_x \approx 3.1$ MeV are seen in the spectrum supporting the collective interpretation of the latter.

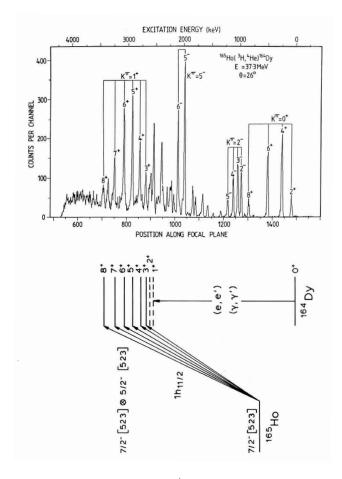


FIG. 10 Members of a $K^{\pi}=1^+$ rotational band observed in the single-particle transfer reaction $^{165}\mathrm{Ho}(^3\mathrm{H},^4\mathrm{He})^{164}\mathrm{Dy}$. The (not populated) 1^+ band head lies at an excitation energy of 2.543(13) MeV which can be identified with the 1^+ state energy of 2.539 MeV, observed in nuclear resonance fluorescence and inelastic electron scattering experiments (Richter, 1990).

f. The scissors mode and deformation So far deformation has already been mentioned alongside the discussion of the experimental data on the scissors mode. The certainly most important observation made since the discovery of the mode itself has been, that the measured orbital magnetic dipole strength increases linearly with the square of the deformation parameter δ . This is shown in Fig. 11, where the summed M1 strength is plotted vs. δ^2 , for a chain of even-even Sm isotopes (Ziegler et al., 1990). This striking result has later been verified in corresponding experiments on a series of eveneven Nd isotopes (Margraf et al.). Those observations have been anticipated in a systematic theoretical study (De Coster and Heyde, 1989) of M1 strength in nuclei of the rare-earth region within the Nilsson model. Immediately after the experimental results displayed in Fig. 11, it was further realized that the orbital M1strength in a given nucleus is also proportional to the strength of the E2 transition to the lowest $J^{\pi} = 2^{+}$ state (Rangacharyulu et al., 1991). The empirical relation

$$\sum_{f} B_{f}(M1) \uparrow \sim B(E2; 0_{1}^{+} \to 2_{1}^{+}) \sim \delta^{2}, \quad (1)$$

thus represents - as expressed by Lo Iudice (2000) - the most spectacular manifestation of the scissors nature of the low-lying magnetic dipole transitions and their collectivity. Since the neutron-proton interaction is mainly responsible for the nuclear quadrupole deformation the experimental observation of the strong M1/E2 correlation is of great interest for a test of nuclear models of deformation as will be shown below.

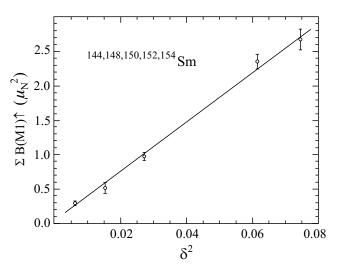


FIG. 11 The summed orbital M1 strengths observed in eveneven Sm isotopes and plotted vs. the square of the deformation parameter δ (Ziegler et~al., 1990). See also Fig. 24 for the systematics of rare-earth nuclei including the present data.

g. Summary In the subsections a-f above, experimental evidence has been presented for the existence of

strong magnetic dipole transitions of orbital character into states at an excitation energy of $E_x \approx 3 \text{ MeV}$ in eveneven heavy deformed nuclei. The total orbital strength into these states amounts in strongly deformed nuclei to about $\sum B(M1) \approx 3 \,\mu_N^2$. It hardly moves at all with excitation energy, i.e. it remains low-lying, is scissors-like and weakly collective, but strong on the single-particle scale. Its observability is a strong effect as a consequence of the fact that the nuclear particle-hole force has swept the competing stronger spin-flip strength up to higher excitation energy (see Fig. 1 and Sec. III.B). Furthermore the existence of weakly excited two-quasiparticle $J^{\pi} = 1^{+}$ excitations at about twice the pairing gap, i.e. below the states associated with the scissors mode, has been shown. Thus, returning to the schematic picture of Fig. 1 above, the salient experimental features of magnetic dipole excitations below about 4 MeV of excitation energy in heavy deformed nuclei have been demonstrated in this subsection.

The weakly collective scissors mode excitation has become an ideal test of models - especially microscopic models - of nuclear vibrations. Most models are usually calibrated to reproduce properties of strongly collective excitations (e.g. of $J^{\pi}=2^+$ or 3^- states, giant resonances, ...). Weakly-collective phenomena, however, force the models to make genuine predictions, and the fact that the transitions in question are strong on the single-particle scale makes it impossible to dismiss failures as a mere detail, especially in the light of the overwhelming experimental evidence for them in many nuclei (Kneissl *et al.*, 1996; Richter, 1995). This should be kept in mind in the assessment of the wide variety of nuclear models which the scissors mode has inspired after its discovery about two and a half decades ago.

2. Theoretical description: from collective to microscopic models

a. Geometric collective models As a special case of the generalized isovector Bohr-Mottelson model, a two-rotor model considering both proton (π) and neutron (ν) degrees of freedom was worked out in detail by Lo Iudice and Palumbo (1978, 1979). Describing these two systems as axially symmetric rigid rotors that are able to perform rotational oscillations around a common axis orthogonal to their symmetry axes (Fig. 12, l.h s.), the following Hamiltonian was set up

$$H = \frac{(\hat{I}_{\pi} + \hat{I}_{\nu})^2}{2J_{intr}} + \frac{(\hat{I}_{\pi} - \hat{I}_{\nu})^2}{2J_{intr}} + \frac{1}{2}C\theta^2,$$
 (2)

in which the restoring force constant C is related to the symmetry energy constant in the semi-empirical mass formula. Using known properties of deformed nuclei: the moment of inertia J_{intr} of the ground-state band, the symmetry energy and the $B(E2; 0_1^+ \rightarrow 2_1^+)$ transition strength, one can determine both the excitation energy

 $E_{sc} = \hbar \sqrt{C/J_{intr}}$ and the B(M1) transition strength

$$B(M1; 0^+ \to 1^+) = \frac{3}{16\pi} \frac{J_{intr}}{\hbar^2} E_{sc} (g_p - g_n)^2 \mu_N^2.$$
 (3)

Here g_p and g_n denote the orbital gyromagnetic factors associated with the rotation of the deformed proton and neutron systems, respectively. It turned out that the early calculations of both the excitation energy and the M1 strength gave too large values compared with the first experimental observations of 1⁺ scissors excitations in deformed nuclei (the B(M1) value exceeds experimental values for the strongest 1^+ states by a factor of almost 7). However, it cannot be emphasized enough how important the seminal work of Lo Iudice and Palumbo (1978, 1979) has been in the experimental search which finally led to the discovery of the scissors mode in ¹⁵⁶Gd (Bohle et al., 1984b). Moreover, the confrontation of the conceptually simple two-rotor model (TRM) with the wealth of experimental data having accumulated rapidly in the late eighties and early nineties of the last century has led to a steady improvement of the model (Lo Iudice, 1997, 2000) which is still the one allowing a good first insight in the dynamics which causes the scissors mode to show up at all in deformed nuclei.

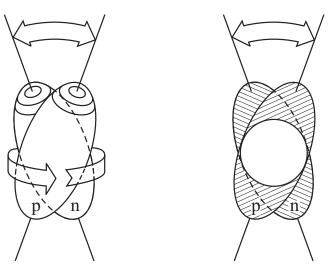


FIG. 12 Pictorial representation of the scissors mode in a proton-neutron two-fluid model (left part) and using a presentation where an inert core exists and only a small part of the proton and neutron fluids take part in the scissors motion (right part). The proton-neutron rotational oscillation is the basis of the two-rotor model which is the rotational analogon of the semiclassical model of the electric dipole giant resonance.

A geometrical model for strongly deformed nuclei (assuming axial symmetry) with separate proton and neutron deformations has been formulated by Rohozinski and Greiner (1985). There, the scissors mode was explained as a relative vibration of the proton and neutron collective surface. Moreover, rotational bands are obtained on top of $K^{\pi} = 0^+$, 1^+ and 2^+ vibrational excitations.

Through the experimental observations, we now know the restoring force constant C in Eq. (2) much better which determines largely the excitation energy E_{sc} of the mode. Furthermore, introducing pairing correlations amongst the participating protons and neutrons and thus effectively reducing the number of protons and neutrons participating in the scissors motion (Fig. 12, r.h.s.) resulted in reduced M1 strengths, but still too big by factors 4-5 as compared to experiment. In this purely collective approach, all M1 strength is concentrated in a single state whereas the data are much more fragmented.

As a consequence of the too high excitation energy and the too large M1 strength for the 1^+ scissors excitation, Faessler and Nojarov (1986) concentrated on a more detailed study of the restoring force for this isovector mode (Faessler *et al.*, 1986). Considering the fact that the symmetry energy coefficient exhibits a strong $\rho^{2/3}$ density dependence, the isovector symmetry energy becomes (Faessler and Nojarov, 1987; Nojarov *et al.*, 1986)

$$E_{sym}^{iv} = D \int (\rho_p - \rho_n)^2 / (\rho_p + \rho_n)^{1/3} d\vec{r},$$
 (4)

with $D=91.6~{\rm MeVfm^2}$. This symmetry energy has been calculated microscopically, using proton and neutron densities ρ_p, ρ_n derived from a deformed Woods-Saxon potential, also including pairing. Identifying the isovector symmetry energy with the scissors potential energy $1/2C\theta^2$, a much improved restoring force constant C could be derived. This reduces the scissors excitation energy by a factor of about 2 but still a too large B(M1) value is obtained. Isovector motion of protons and neutrons has also been discussed (Faessler, 1966; Faessler and Nojarov, 1987) for systems performing harmonic small-amplitude vibrations around a spherical equilibrium shape, using an extended Bohr-Mottelson model. The implications will be studied extensively in Sec. VI.

The collective two-fluid model, both for rotational as well as for vibrational excitations, has been studied and refined over the years in great detail. The results have been summarized in review articles by Lo Iudice (1997, 2000) and partly in an article by Zawischa (1998). We refer the reader to these for interesting details.

b. Algebraic collective models When discussing the nuclear structure aspects of an interacting fermion system, it is striking that for the low-energy collective modes to develop, the nucleon-nucleon correlations acting in the L=0 (paired state) and also in the L=2 configuration are particularly important. It has been shown that the L=0 correlations amongst identical nucleons lead to a generalized seniority classification while the addition of the L=2 pair component gives rise to the possibility to develop strong collective excitations when both proton and neutron valence particles are present. Considering those two-nucleon pair configurations, it is possible to formulate a model in which these pairs are now

treated as genuine bosons: the L=0 pair is mapped onto the s boson and the L=2 pair onto the d boson (Arima and Iachello, 1975a,b). This system of interacting bosons (IBM concept) has been studied in great detail, in particular emphasizing the group structure (which is the group U(6)) and its reductions, by Arima and Iachello in a series of papers (Iachello and Arima, 1987).

In making the model more closely-looking to the shell model with protons and neutrons interacting, the charge degree of freedom was introduced in order to distinguish between s and d proton and neutron bosons, doubling the space of independent degrees of freedom. The fact that for bosons the total wave function needs to be symmetric under the interchange of any two bosons, it is still possible to construct wave functions that have mixed-symmetry character in both the spatial and the charge part, separately. Using the group-theoretical formulation, the product irrep. (irreducible representations) of $U_{\pi}(6) \otimes U_{\nu}(6)$ contains, besides the one-row, also two-row irrep., or even more explicitly,

$$[N_{\pi}] \otimes [N_{\nu}] = [N_{\pi} + N_{\nu}, 0] \oplus [N_{\pi} + N_{\nu} - 1, 1] \oplus \dots$$
 (5)

The physics of these mixed-symmetry U(6) irrep. becomes most clear when studying the energy eigenvalues for a total IBM-2 Hamiltonian which contains, besides the pure proton and neutron parts, also the coupling term between the proton-neutron combined parts. This IBM-2 Hamiltonian can be written as

$$H = \epsilon_{d_{\pi}} \hat{n}_{d_{\pi}} + \epsilon_{d_{\nu}} \hat{n}_{d_{\nu}} + \kappa (\hat{Q}_{\pi} + \hat{Q}_{\nu}) \cdot (\hat{Q}_{\pi} + \hat{Q}_{\nu}) + \lambda \hat{M}_{\pi\nu}.$$
(6)

We can show the essential results easily in the case of only two bosons (Fig. 13) where one s_{π} and d_{π} and one s_{ν} and d_{ν} boson are considered. The symmetric coupling [2,0] corresponds to the 0, 1 and 2 quadrupole phonon structure of the well-known symmetric quadrupole vibrator; the [1,1] irrep. gives rise to the non-symmetric 1^+ , 2^+ and 3^+ levels. Here too, the energy separation between the 0^+ and the 2_{ms}^+ states is related to the collective symmetry energy in the interacting boson model, an energy which is governed by the strength λ of the Majorana term $M_{\pi\nu}$ in the IBM-2 Hamiltonian of Eq. (6), see Scholten et al. (1985b). This is very much the same physics underlying the splitting of the various isospin T components resulting from combining protons and neutrons in fermion space, as discussed in the introduction (Fig. 3). An equivalent two-valued variable (called F-spin) has thus been introduced (Arima et al., 1977; Iachello, 1984) to characterize the charge (or spatial) part of the boson wave function. If the proton and neutron bosons are characterized with the help of their F-spin quantum number $F = 1/2, F_z = +1/2$ and $F = 1/2, F_z = -1/2$, respectively, a system of N_{π} proton and N_{ν} neutron bosons can be classified according to its total F-spin. The totally symmetric orbital (sd boson space) states have maximal F-spin, i.e. $F_{max} = 1/2(N_{\pi} + N_{\nu})$ while the mixedsymmetric states are labeled by decreasing F-spin values down to $F_{min}=1/2|N_{\pi}-N_{\nu}|$. The class of mixed-symmetry states with $F=F_{max}-1$ are the lowest-lying

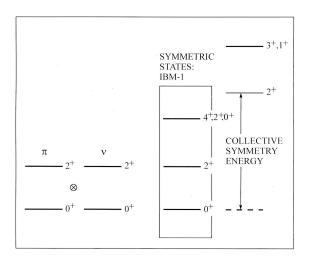


FIG. 13 Schematic representation of a coupled protonneutron system with boson numbers $N_{\pi} = 1$, $N_{\nu} = 1$. The symmetric states (in the box) and the antisymmetric states are drawn at the right-hand side of the figure (adapted from (Heyde, 1989)).

and can be excited from the ground state of an even-even nucleus by a $\Delta F = 1$ transition.

The usual M1 operator in fermion space

$$T^{F}(M1) = \sqrt{\frac{3}{4\pi}} \sum_{i} [g_{l}(i)\hat{l}_{i} + g_{s}(i)\hat{s}_{i}]\mu_{N}, \qquad (7)$$

has its image in boson space

$$T^{B}(M1) = \sqrt{\frac{3}{4\pi}} (g_{\pi} \hat{L}_{\pi} + g_{\nu} \hat{L}_{\nu}) \mu_{N}, \tag{8}$$

with g_l , g_s being the fermion orbital and spin g-factors and g_{π} , g_{ν} the respective proton and neutron boson g factors. The quantities \hat{L}_{π} and \hat{L}_{ν} are the corresponding orbital angular momentum operators of the proton and neutron boson system.

In IBM-2 studies, concentrating on deformed nuclei, one is using the SU(3) reduction of the U(6) group structure and, for those nuclei, it was shown that the lowest-lying states of the family of mixed-symmetry character were characterized by the $F_{max}-1, J^{\pi}=1^+$ quantum numbers (Iachello, 1981, 1984). These findings corroborate results obtained from a totally different starting point, viz. the TRM. The IBM-2 $J^{\pi}=1^+$ states are also called scissors states although there is no immediate reference in the algebraic formulation to specific coordinate forms and thus also not of shapes and shape dynamics. Using a coherent-state formalism, Dieperink (1983) was able to show the correspondence explicitly and, moreover, found indeed that only the valence nucleons are contributing to the strength of the scissors mode thus

leading in a natural way to a much lower B(M1) strength compared to the early TRM calculations.

In studying the M1 excitation properties within the IBM-2, because of the specific difference in magnetization properties for proton and neutron bosons, it was clear that M1 transitions could appear naturally now, in contrast to the former IBM-1. Using mapping from fermion magnetic properties onto boson ones, it was possible to also determine the analogous boson g_{π} and g_{ν} factors (Allaart et al., 1988; Sambataro and Dieperink, 1981; Sambataro et al., 1984). This item has been a topic of much discussion because the mapping calculations all seem to come up more or less with values $g_\pi \simeq 1~\mu_N$ and $g_\pi \simeq 0~\mu_N$ but empirical fits in various mass regions have indicated quite important deviations (Kuyucak and Stuchbery, 1995; Mizusaki et al., 1991; Wolf et al., 1987). For the pure SU(3) limit, though, an analytically closed form could be derived for the transition strength

$$B(M1) = \frac{3}{4\pi} \frac{8N_{\pi}N_{\nu}}{2N - 1} (g_{\pi} - g_{\nu})^2 \ \mu_N^2, \tag{9}$$

an expression when applied to 156 Gd and using the above boson factors g_{π}, g_{ν} gives the result of $B(M1) \simeq 2.8 \ \mu_N^2$, quite close to the experimentally observed value. Thus the realistic estimate of Iachello (1981), in the IBM-2, for the transition strength has been equally important in the search for the scissors mode (Bohle *et al.*, 1984b) as the estimate of Lo Iudice and Palumbo (1978, 1979) in the TRM for the excitation energy.

The subject of mixed-symmetry states appearing as a new class of states in the IBM-2 has been investigated afterwards in much detail. There has been an investigation of the various limiting cases that appear if dynamical symmetries hold (Van Isacker $et\ al.$, 1986) but also rather extensive numerical studies have been carried out (Scholten $et\ al.$, 1985b). As an illustration, we compare in Fig. 14 the experimental M1 transition strengths in Sm isotopes with the results of the pure SU(3) limit and of the numerical IBM-2 calculations (Scholten $et\ al.$, 1985b).

c. Microscopic descriptions On the opposite side from the collective model concepts, the nuclear shell-model allows for possibilities to describe nuclear coherent motion from first principles using a Hartree-Fock basis and a self-consistent procedure in order to determine both global and local nuclear structure properties. The study of scissors motion starting from a microscopic shell-model basis can be separated into two parts. For light and medium-heavy nuclei, regular shell-model calculations have been performed and also M1 excitation properties been studied. Once entering the region of heavy and deformed nuclei, the model space to be considered becomes prohibitively large and approximations to the shell model have been used, mainly the QRPA.

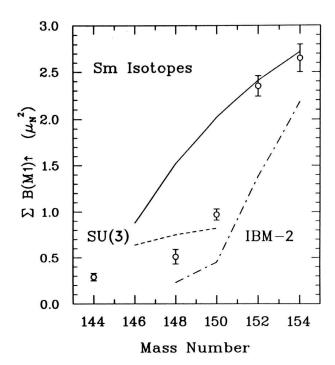


FIG. 14 Orbital M1 strength versus mass number of the Sm isotopes. The results from full IBM-2 calculations (dash-dotted line) of Scholten *et al.* (1985b) are shown, together with results in the SU(3) limit using a Z=50 subshell closure (solid line) and a Z=64 subshell closure (dashed line) (Richter, 1991).

Magnetic dipole excitations have, by now, been measured over a large part of the nuclear mass table (Sec. III.A.1). In contrast to most of the purely collective models, the low-lying M1 strength is spread out over an energy interval in the region of 2.5-4 MeV, depending on the specific nucleus and thus depending on its proximity to the closed shells.

Large-scale shell-model studies would be an ideal way to probe the presence of concentration and fragmentation of M1 strength but this has not been possible for heavy nuclei until recently. Within the context of a Monte-Carlo shell-model approach, worked out by Otsuka and collaborators (Honma et al., 1995, 1996; Mizusaki et al., 1996, 1999; Otsuka and Mizusaki, 1999; Otsuka et al., 1998; Shimizu et al., 2001; Utsuno et al., 1999), largescale shell-model studies have been performed in order to study the transition from spherical to deformed shapes with an application to the Ba isotopes. Starting from a given Hamiltonian and for a given single-particle energy spectrum that remains fixed through the Ba isotopes, such microscopic calculations beyond mean-field approaches have given first evidence, that shape changes indeed do occur due to a change of the number of interacting protons and neutrons. The M1 strength of the scissors mode serves as a good measure of the deformation of the ground state as has been discussed in Sec. III.A.1 in particular in the context of Fig. 11 and Eq. (1). Shimizu et al. (2001) have calculated the M1sum rule for the ground state with the orbital g-factors having the free nucleon values. The spin contributions have been omitted for reasons of simplicity, at present. In Fig. 15, the B(M1) sum rule values are plotted versus the corresponding $B(E2; 0_1^+ \to 2_1^+)$ values. One notices a nearly perfect linearity between these two quantities first observed experimentally (Ziegler et al., 1990). As pointed out in the previous subsection, starting from symmetries within the IBM and in the nuclear shell model, such a relation also exists. The Monte-Carlo shell-model calculation, however, presents a first microscopic underpinning of the connection between M1 and E2 properties in nuclei and shows a validity that does not rely any longer on particular symmetries of the nuclear many-body system, i.e. it holds for the whole isotopic series in the Ba nuclei.

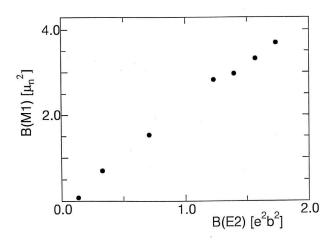


FIG. 15 Summed orbital B(M1) strength from the ground state calculated in a Monte-Carlo shell-model approach versus the corresponding $B(E2;0^+_1 \to 2^+_1)$ values in the Ba isotopes with mass numbers A=138-150. (Reprinted with permission from Shimizu *et al.* (2001). ©(2009) Am. Phys. Soc.)

Besides the nuclear shell-model explicitly, the QRPA or the Quasi-Particle Tamm-Dancoff Approximation (QDTA), where no ground-state correlations are considered, present an alternative to study the properties and the internal orbital and spin character of the magnetic dipole transitions involving the specific 1⁺ states under study. The QRPA and QDTA models in itself will not be discussed in the present review and we refer the reader to the literature for a concise discussion (Eisenberg and Greiner, 1987; Ring and Schuck, 1980; Rowe, 1970; Soloviev, 1992). In the early calculations when applying the QRPA to study the scissors mode excitations, most often rather schematic forces have been used: quadrupole protonneutron forces and spin-spin and spin-isospin separable interactions, including pairing (both monopole and quadrupole multipoles) in order to study the role of these

components (Bes and Broglia, 1984; Hamamoto, 1971; Iwasaki and Hara, 1984; Kurasawa and Suzuki, 1984). Quite recently (Balbutsev and Schuck, 2004, 2007), a separable quadrupole-quadrupole residual interaction has been used within a time-dependent Hartree-Fock (TDHF) framework and its Wigner transform. Here, the aim was to derive a set of equations describing different multipole moments, in particular the scissors and isovector giant quadrupole resonance and their coupling. Pairing, however, is not included in the formalism. The results obtained are closely related to the earlier work of Lipparini and Stringari (1989b).

The pairing force is particularly important in generating the necessary correlations that relate the overall summed orbital M1 strength to the nuclear quadrupole deformation characterizing a given nucleus. Moreover, the spin-isospin parts are determining factors in placing the spin-flip part of the M1 response (Fig. 1) at the excitation energy between 5 and 10 MeV. These studies were instrumental in finding out the first and foremost important physics elements that are at the origin of coherent orbital magnetic excitations with a scissors character. It has been pointed out (Heyde and De Coster, 1993; Ikeda and Shimano, 1993) that a correct treatment of the full Coriolis term and considering the coupling of 2qp excitations to a rotational core, induces very specific correlations that concentrate the independent M1components into a single peak.

The realistic application of the QRPA to the study of the scissors mode exciting $J^{\pi} = 1^{+}$ states has met with some initial problems because the overall rotational motion of an intrinsically deformed nucleus carries the same angular momentum as the scissors mode states themselves. Very much like in case of the electric dipole mode leading to the $J^{\pi} = 1^{-}$ states, one has to separate the spurious rotational motion of the whole nucleus from the relevant intrinsic 1⁺ excitations. This was the origin of the fact that various QRPA calculations gave rise to rather large differences in the 1⁺ excitation energy and, more importantly, in the M1 response. Removing this spurious rotational motion, however, using various techniques such as (i) constructing a basis orthogonal to the spurious rotational motion, (ii) adding a specific symmetry restoration term to the Hamiltonian, (iii) using the Pyatov prescription (Baznat and Pyatov, 1975) of replacing the quadrupole field in the Hamiltonian in such a way that rotational invariance is imposed, the results on $J^{\pi} = 1^+$ energies and on M1 strengths have all been converging with similar conclusions.

So, the various QRPA studies are now very close with respect to a number of key issues characterizing the observation of a magnetic scissors-type of excitation (Beuschel et al., 2000; Nojarov and Faessler, 1990; Soloviev et al., 1997a,b,c; Zawischa et al., 1990; Zawischa and Speth, 1990). Examples are shown in Figs. 16-18. Work using variants of QRPA starting from a deformed Woods-Saxon potential, HFB calculations, and schematic forces has also

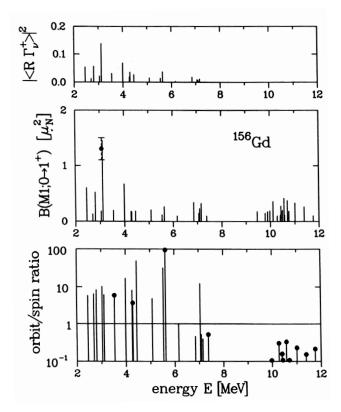


FIG. 16 Calculated QRPA magnetic dipole strength distribution compared to the experimental data point of the strongest transition observed in ^{156}Gd (middle part) and orbit-to-spin ratios of the M1 transition matrix elements (bottom part). The magnitude of these ratios is shown and negative ratios are indicated by a dot on the top of the bar. The average overlap of the calculated 1^+ states with the scissors mode state (denoted R) is displayed in the upper part. (From Nojarov and Faessler (1990), reprinted with permission from Springer Service + Business Media.)

been carried out (Hamamoto and Ronström, 1987; Hilton, 1995; Hilton *et al.*, 1998; Lo Iudice, 1996a,b; Sugawara-Tanabe and Arima, 1989).

- (i) A concentration of low-lying M1 strength is found close to the energy of 3 MeV, whose overlap with the scissors mode is as big as $\simeq 40$ % if summed up to 4 MeV(see Fig. 16, upper part). The low-energy M1 strength is of dominant orbital character with orbit/spin ratios of the order of 10 or larger (cf. Fig. 16, bottom part as an example). Furthermore, all calculations predict the existence of a higher-lying scissors part. This issue, however, is not closed at all since various authors come to largely different conclusions concerning its mean energy and, more importantly, its fragmentation.
- (ii) There is common agreement on the very importance of the pairing correlations in establishing a strong relationship between the summed M1 strength and nuclear deformation. These at first quite different observables have a deeper connection which is born out from the microscopic QRPA studies too as will be pointed out

below in some more detail.

(iii) An important spin-flip M1 mode is observed in the strongly deformed rare-earth nuclei in the energy region 5-10 MeV. There is, also here, an open debate on the specific way in which the spin-flip strength is distributed in energy. A two-bump picture shows up but in some calculations the bumps are mainly isoscalar and isovector in character, whereas other calculations comply with a rather definite separation between proton and neutron spin-flip modes. The solution here should come from selective reaction studies (see Sec. III.B for a deeper discussion on this issue).

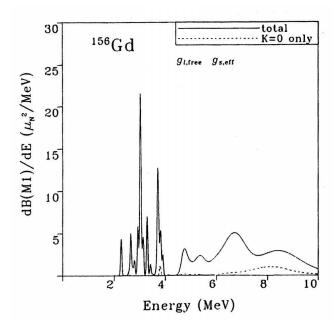


FIG. 17 The complete magnetic dipole response of ¹⁵⁶Gd up to 10 MeV excitation energy, calculated using the QRPA approach. The individual levels have been folded with Gaussian functions. (Reprinted with permission from Zawischa et al. (1990); Zawischa and Speth (1990). ©(2009) Am. Phys. Soc.)

We cannot provide a complete discussion of the multitude of QRPA and QTDA studies in both the rareearth and actinide (and even light) nuclei. We refer to Lo Iudice (1997) for a detailed but still succinct presentation of the major results. We note, however, that very recently relativistic QRPA calculations within a self-consistent relativistic mean-field (RMF) framework have been carried out for axially deformed nuclei (Peña Arteaga and Ring, 2008). Spin, orbital and total M1 strengths were derived for 160 Gd and 20 Ne, with clear evidence for a scissors mode in 160 Gd (Peña Arteaga and Ring, 2007).

d. Relationship between collective and microscopic models Starting from collective models, the proton and neutron degrees of freedom form the essential ingredients to gen-

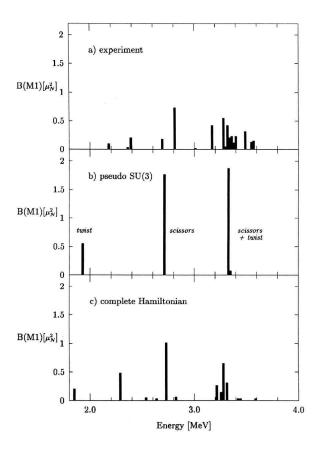


FIG. 18 M1 transition strengths in 160 Gd. (a) Experimental data, (b) pure pseudo-SU(3) scheme, and (c) complete pseudo-SU(3) calculation. The twist mode results when considering triaxial proton and neutron distributions which allow additional rotation around their principal axes. It can be combined with the scissors mode into a scissors plus twist mode. (Reprinted with permission from Beuschel *et al.* (2000). \odot (2009) Am. Phys. Soc.)

erate mixed-symmetry charge (and spatial) wave function of the $J^{\pi}=1^+$ states. This was the case in the two-rotor geometrical model and also within the protonneutron interacting boson model (IBM-2).

Because the building blocks constituting the bosons are nucleon pairs, the IBM-2 approach is rooted closely in the nuclear shell-model (see e.g. the studies in the light $1f_{7/2}$ nuclei by Liu and Zamick (1987a,b,c); McCullen et al. (1964); Zamick (1986a,b)). In the QRPA, on the other hand, the building blocks are highly correlated particle-hole (or 2 qp) excitations that make up for a microscopic description to the collective phonon modes in the nucleus. It has been pointed out explicitly (Lo Iudice, 1997) that in schematic models, the connection between the microscopic QRPA and the collective two-fluid approaches (TRM) can be shown in detail. An approximate relation for the scissors total summed M1 strength is then

$$B(M1) \uparrow \simeq \frac{3}{16\pi} \frac{J_{intr}}{\hbar^2} E_{sc} \ \mu_N^2, \tag{10}$$

if all the 1^+ states could be approximated by a single centroid energy E_{sc} like in Eq. (3).

In all of the more detailed microscopic calculations, be it large-scale shell-model studies or QRPA calculations, the immediate connection is not so obvious anymore. Still, the observed spreading (or fragmentation) of the population of 1⁺ states is rather well reproduced. The QRPA calculations mainly start from different single-particle structures (Nilsson or Woods-Saxon deformed mean-field, self-consistent energy spectra,...) and also have been using differing two-body interactions and so, a large sensitivity of the 1⁺ energy and strength distribution in realistic cases is expected.

There exists an interesting avenue in trying to give a microscopic support of collective magnetic dipole excitations starting from the SU(3) for light nuclei and pseudo-SU(3) for the heavy deformed nuclei. This model goes back to the seminal work of Elliott (1958, 1963) indicating that collective model aspects are inherent in the proton-neutron shell model structure for light sd shell nuclei (Elliott, 1985, 1990). Calculations within this approximation have been carried out in sd- and fp-shell nuclei for the orbital M1 properties and compared to the experimental values (Chaves and Poves, 1986; Poves et al., 1989; Retamosa et al., 1990). The extension to heavier nuclei was hampered for some time as SU(3) is largely broken because of the strong spin-orbit splitting. A suggestion has, however, been made in order to reorganize the level structure conform with a pseudo SU(3) scheme, explicitly incorporating proton and neutron degrees of freedom. In former studies of this type, one had to rely on very simple Hamiltonians; this has been overcome and one can now handle general one- and two-body Hamiltonians. So this model, deeply rooted in the shellmodel, serves as an interesting bridge to connect the underlying microscopic structure to the collective building blocks. Results for Sm, Gd and Dy nuclei have been obtained (Beuschel et al., 1998, 2000; Rompf et al., 1998) and even a reasonable fragmentation of M1 strength is achieved as illustrated in Fig. 18 for ¹⁶⁰Gd.

3. Fragmentation of orbital dipole strength and sum rules

a. Fragmentation of the orbital strength An important element in discussing relationships, similarities but also complementary aspects of the collective model approaches and shell-model or QRPA studies is the amount of fragmentation resulting from these various models. As discussed in Sec. III.A.2, in collective models, the M1 scissors strength is concentrated in a single or very few strongly excited 1^+ states (Scholten et al., 1985b) while in the microscopic models, strength needs to become concentrated in fewer strong states compared with the unperturbed spectrum of 1^+ states.

What are the major issues here? We present the above elements of fragmenting collective M1 strength into a background of microscopic configurations (2qp,4qp,..)

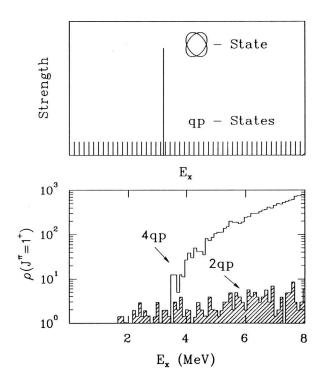


FIG. 19 Upper part: schematic picture of the scissors mode embedded into the background of dense qp states (no coupling). Lower part: level density for 2qp and 4qp 1^+ states for Gd isotopes calculated using the Nilsson model (adapted from (Heyde, 1989)).

much in the same way as fragmentation is generally described in nuclear reaction theory (Fig. 19, upper part). There might remain some structure in the fragmented strength due to interactions with states that are intermediate in complexity between the strongly collective states on one side and the regular shell-model configurations on the other side. Such states could be due to hexadecapole configurations (in certain regions of the nuclear mass table this degree of freedom in the upper part of the rareearth mass region, particularly, can be important), triaxial shape configurations,.. that first split the M1 scissors strength in the manner of a doorway-state before it gets fragmented into the microscopic background of 1⁺ states (Scholten et al., 1985b). These strength function phenomena have been discussed e.g. in the Appendix D2 of Bohr and Mottelson (1969). If the average coupling strength $\langle V \rangle$ between a single collective state and the background configurations is larger than the distance between the discrete levels $D = 1/\rho$ with ρ denoting the level density, then a Breit-Wigner damping of collective strength over the microscopic background results, given by a width of $\Gamma = 2\pi \langle V \rangle^2 \rho$. The strength function, i.e. the probability of finding a simple, collective state in a unit energy interval of the spectrum, can subsequently be derived. If, on the other hand, the level density becomes too low for the above conditions to be valid, one has to resort to diagonalizing the coupled system of collective and shell-model configurations. A model where this has been worked out explicitly for the scissors 1^+ state, as obtained in the IBM-2, coupling to the underlying background of 1^+ 2qp and 4qp configurations was described by Heyde et al. (1996). The background structure, as calculated using a Nilsson model for the Gd nuclei, is shown in the lower part of Fig. 19. The damping and fragmentation of the scissors mode into this background has been derived using a constant coupling matrix element between the states in the two model spaces (Fig. 20, upper part). For the fragmentation down to the 1% level, a coupling matrix element of about 50 keV is obtained between the scissors mode and the 2qp states. The experimental M1 strength plotted in the lower part of Fig. 20 indicates a similar fragmentation.

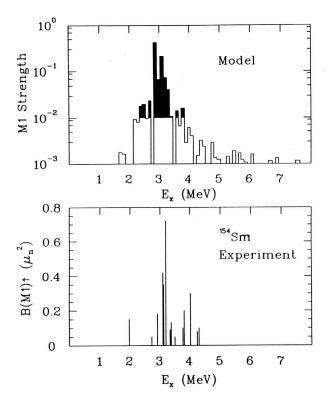


FIG. 20 Upper part: scissors M1 strength damped into the background of 2qp and 4qp states shown in lower part of Fig. 19, averaged over intervals of 100 keV. The black part indicates transitions with a relative strength larger than 1% (adapted from Heyde (1989)). Lower part: experimentally measured orbital M1 strength distribution (Ziegler *et al.*, 1993).

Starting from the opposite side, i.e., a shell-model or QRPA approach, the diagonalization within the unperturbed basis of 2qp, 4qp,... states will eventually result in the building up of collective correlations that are reminiscent of a scissors magnetic dipole mode. Examples have been shown in Figs. 16-18 in the heavy deformed rare-earth nuclei. The discussion in Sec. III.A.2 has precisely concentrated on this issue with references

to a number of calculations of that type. One can make use of very similar arguments as those used at the time by Brown and Bolsterli (1959) in the discussion of how individual 1p-1h 1⁻ configurations, interacting through a zero-range force, could built a single collective, isovector electric dipole mode. In contrast to the study of this giant electric dipole resonant state in which the collective state is mainly built from $1\hbar\omega$ excitations, the orbital magnetic dipole strength is mainly of $0\hbar\omega$ nature. Richter and Knüpfer (1980) have shown that the separable characteristics of the two-body matrix elements, which is essential for the schematic Brown-Bolsterli model, does not hold any more for the magnetic multipole excitations, i.e. the particle-hole matrix elements do not scale with the M1 transition amplitudes. As a result, rather small energy shifts show up compared to the unperturbed energy spectrum. Therefore, there remains a concentration of mainly orbital M1strength of $0\hbar\omega$ origin in the energy interval 2.5-4MeV. An example of the fragmentation process, starting from the pseudo SU(3) description, but treating a more general Hamiltonian, has been discussed by the group of Draayer and a particularly interesting case is ¹⁹⁶Pt (Beuschel et al., 2000) which is a prime example of a γ soft nucleus (von Brentano et al., 1996).

b. Level spacing distribution of scissors mode states As we have noticed in the foregoing subsection, the complexity of the nuclear many-body problem is clearly manifest in the fragmentation of the experimental transition strength, which is distributed over several levels of the same spin and parity. This complexity has led Wigner, more than forty years ago, to the introduction of Random Matrix Theory (RMT), reviewed in detail by Guhr et al. (1998) and Weidenmüller and Mitchell (2009). This statistical approach models spectral fluctuation properties: if the levels are correlated, one expects a linear repulsion between them and Wigner-Dyson statistics for the nearest neighbor spacing distribution (NNSD). However, if correlations are absent, there is no level repulsion and the NNSD is of Poisson-type, i.e. an exponential distribution. The validity of this ansatz has been confirmed in various data analyses and has been summarized in the two review articles cited.

In heavy nuclei, the picture emerges that high-lying single-particle states containing many complex configurations show Wigner-Dyson statistics (Haq et~al., 1982), whereas low-lying collective states of simple structure lack correlations and yield a Poisson distribution (Garrett et~al., 1997; Shriner Jr. et~al., 1991). This in turn allows to use RMT to conclude from spectral statistics if excitations are mainly of single-particle or of collective character. This idea has also been applied to the states which belong to the scissors mode (Enders et~al., 2000). As has been pointed out in Sec. III.A.1 above, an unprecedented data set is now available covering doubly even nuclei in the N=82-126 major shell. By com-

bining the data sets from 13 heavy deformed nuclei, a data ensemble has been constructed with a total number of 152 states in the excitation energy window of about $2.5 < E_x < 4.0$ MeV (Fig. 1).

After unfolding the experimental spectra, i.e. removing the energy dependence of the average level spacing, the NNSD, the cumulative NNSD, the number variance Σ^2 and the spectral rigidity Δ_3 (Bohigas, 1991; Guhr et al., 1998) were extracted from the data ensemble. The results are shown in Fig. 21. All evaluated statistical measures agree very well with the Poissonian behavior of uncorrelated levels. Although the individual level sequences are rather short (of order ten only), the functions Σ^2 and Δ_3 clearly show the lack of long-range correlations. It has also been shown by Enders et al. (2000) that the influence of missing levels due to the experimental conditions is negligible. Consequently the remarkable conclusion from the statistics of the level spacings distribution may be drawn that the scissors-mode states all have the same structure and are excited collectively by the same mode.

c. Sum rules and relation to other observables Even though the magnetic dipole strength appears rather fragmented in the energy region 2.5 to 4.0 MeV, experimental methods have been set up in order to distinguish those 1^+ states that carry mainly orbital M1 strength (see also the discussion by Kneissl et al. (1996)). The summed experimental strengths show remarkable correlations to collective observables of the low-energy spectrum.

Firstly, as pointed out already above, the summed M1strength correlates with the square of the equilibrium quadrupole deformation value for series of isotopes like the Nd and Sm nuclei (Fig. 11, taken from Ziegler et al. (1990)). Secondly, the summed M1 strength also correlates linearly with the $B(E2; 0_1^+ \rightarrow 2_1^+)$ value for most nuclei in this mass region (Richter, 1995). Furthermore, both the summed M1 strength and the particular B(E2) value scale in the same way and saturate (Rangacharyulu et al., 1991) when plotted, not as a function of proton(neutron) number (Fig. 22), but using the variable $P = N_p N_n / (N_p + N_n)$ with $\frac{N_p}{2}, \frac{N_n}{2}$, the number of proton and neutron pairs, respectively, counted from the nearest closed shells (the number of bosons in the IBM-2 model), introduced by Casten et al. (1987). Thirdly, it has been shown that the summed M1 strength even scales linearly with the isotopic shift for those nuclei (Nd, Sm, Dy) where both sets of data are available (Heyde et al., 1993). These three important observations all point towards close interconnections between the M1, E2 and the E0 electromagnetic properties for the rareearth region.

Before going into some more detail discussing what theoretical models predict for the behavior of the summed M1 strength, the deformation dependence of the $B(E2;0_1^+ \rightarrow 2_1^+)$ strength and the radius are rather obvious from a collective geometrical approach

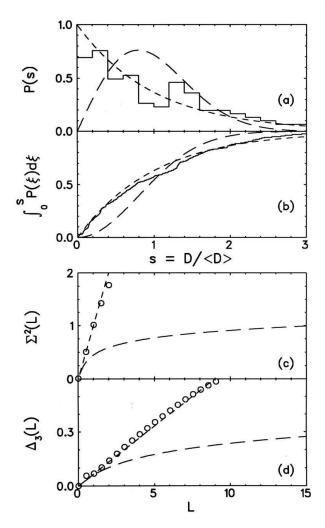


FIG. 21 Level spacing distribution of scissors mode states in heavy deformed nuclei (152 states from 13 nuclei). (a) nearest-neighbor spacing distribution, (b) cumulative nearest-neighbor distribution, (c) number variance $\Sigma^2(L)$, and (d) spectral rigidity $\Delta_3(L)$, where L denotes the length of the sequence. The histograms and open circles display the data. Poissonian behavior and expectations from the Gaussian Orthogonal Ensemble (Wigner) are shown as short and long-dashed lines, respectively (Enders et al., 2000).

(Bohr and Mottelson, 1975). The connection of the latter quantities to the summed magnetic dipole strength, however, was much less expected to appear, in particular the quadratic dependence on deformation and the same saturation behavior when passing through the rare-earth region.

Starting from the generalized Bohr-Mottelson model, Lo Iudice and Richter (1993) have worked out a sum rule that holds very generally and is essentially model independent and parameter free. The resulting expression

$$B(M1) \uparrow \simeq 0.0042 E_{sc} A^{\frac{5}{3}} \delta^2 (g_p - g_n)^2 \mu_N^2,$$
 (11)

directly contains the dependence on deformation as well as on the gyromagnetic factors associated with the col-

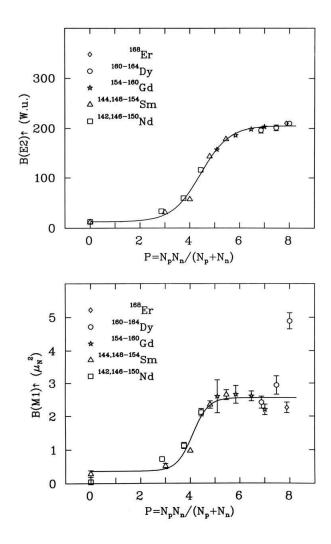


FIG. 22 Upper part: E2 transition strengths of the 2_1^+ states in the even-even rare-earth nuclei indicated in the figure versus P. Lower part: The same for the summed M1 strength in the energy range $E_x = 2.5-4$ MeV. The solid lines correspond to fits explained in Rangacharyulu $et\ al.\ (1991)$.

lective motion of the deformed proton and neutron systems. In the region covering transitional and strongly deformed rare-earth nuclei, the ratio of the experimental over the theoretical summed M1 strength is close to one over a large mass span, with a drop off towards the heavier masses. A specific example in which the sum rule is compared to a number of other models and the data for the Sm nuclei is given in Fig. 23 (Garrido et al., 1991; Hamamoto and Magnusson, 1991; Heyde and De Coster, 1991; Hilton et al., 1993; Lo Iudice and Richter, 1993). While all of them roughly reproduce the quadratic dependence, the theoretical model results exhibit rather different slopes, in some cases with serious deviations from the data for the more deformed A = 152, 154 nuclei. Lo Iudice (1997) has also shown that the TRM and IBM-2 sum rules are very closely related and all can be

brought back to the basic sum rule structure as deriving from the QRPA study of collective magnetic dipole excitations.

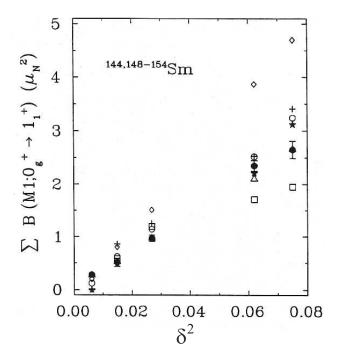


FIG. 23 Orbital M1 strength versus the square of deformation in the Sm isotopes. The experimental points (black circles with error bars) are compared to six theoretical predictions from Hamamoto and Magnusson (1991) (circles), Garrido et al. (1991) (crosses), Hilton et al. (diamonds), Heyde and De Coster (1993)(squares), Lo Iudice and Richter (1993) (triangles), and Sarriguren et al. (1994) (stars), respectively (Richter, 1995). Note that the experimental points do overlap with the triangles in a number of cases, making the distinction difficult.

After the above experimental observations were well established, various theoretical ideas in deriving closed expressions for this summed M1 strength, aiming at establishing at the same time a relation to the E2 and the E0 nuclear properties, have been explored. Here, we discuss implications of both non-energy weighted and energy-weighted sum rules, using collective model as well as shell-model approaches.

Ginocchio (1991) proposed a non-energy weighted M1 sum rule within the IBM for an N-boson system

$$\sum B(M1) = \frac{9}{4\pi} (g_{\pi} - g_{\nu})^2 \frac{P}{N-1} \langle 0^+ | \hat{n}_d | 0^+ \rangle, (12)$$

an expression which connects the summed M1 strength with the expectation value of the number of d bosons in the nuclear ground state. This latter quantity is also a measure of deformation because the average number of d bosons in the ground-state $\langle 0^+ | \hat{n}_d | 0^+ \rangle / N$ can be expressed by a deformation parameter β_{IBM} . A relation was derived (von Neumann-Cosel et al., 1995) between

this specific IBM quantity and the corresponding geometrical definition of deformation like the Bohr-Mottelson parameter β_2 , as

$$\beta_{IBM} = \frac{3\lambda}{2\sqrt{\pi}} \frac{Z}{Z_{val}} \beta_2. \tag{13}$$

Here, Z_{val} describes the number of protons in the valence shell and λ - not to be confused with Majorana strength parameter in Eq. (6) - is a measure of how much of the E2sum rule is exhausted by the transition to the 2_1^+ state. The original sum rule of Ginocchio has thereby been extended and applied to the full range of nuclei spanning the Nd to W region providing an excellent description of the quadratic quadrupole deformation dependence. This is illustrated in Fig. 24.

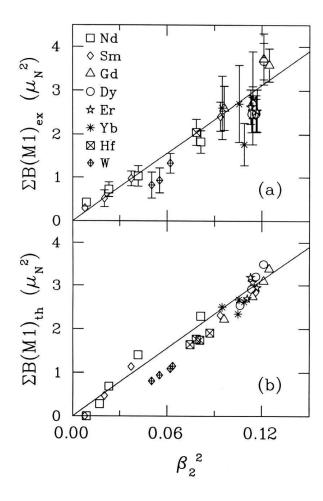


FIG. 24 Experimental M1 scissors mode strengths as a function of the deformation parameter β_2^2 . (a) The straight line is a least-squares fit assuming intercept zero. (b) Prediction of an IBM-2 sum rule for the scissors mode strength of all eveneven stable nuclei from Nd to W (von Neumann-Cosel $et\ al.$, 1995).

Making use of the IBM-2, also other sum rules have been derived (energy-weighted M1 sum rule and even

sum rules for other multipoles). The energy-weighted sum rule indicates a direct relation between summed M1 strength and the $B(E2;0^+_1 \rightarrow 2^+_1)$ value when the assumption is made that most E2 strength remains in the first excited 2^+ state (which is a rather good approximation for transitional and definitely so for deformed nuclei) and gives rise to the following relations

$$\left\langle 0_{1}^{+} \left| [[\hat{H}, \hat{T}(M1)], \hat{T}(M1)] \right| 0_{1}^{+} \right\rangle$$

$$= \frac{2}{\sqrt{3}} \sum B(M1) \left\{ E_{x}(1^{+}) - \lambda N \right\},$$
(14)

or,

$$\sum B(M1)\left\{E_x\left(1^+\right) - \lambda N\right\} = c\sum B(E2) \qquad (15)$$

when using F-spin symmetry in evaluating the quadrupole expectation value. Here, λ denotes the strength of the Majorana term in Eq. (6). This expression indeed relates the energy-weighted M1 sum rule with the non-energy weighted E2 sum rule. The quantity c is introduced to match dimensions of the left- and right-hand side of this equation. This relation is discussed and illustrated more explicitly in Heyde $et\ al.\ (1992)$ in which it is shown under what approximations the above relation reduces to the non-energy weighted M1 sum rule. There, it is also shown that the effect of the Majorana term can be incorporated to a large extent.

It is important though to study analogous sum rules for the magnetic dipole strength but now starting from a shell-model formulation of the problem. Using protons and neutrons, explicitly, Zamick and Zheng (1992) were able to derive an energy-weighted magnetic sum rule which was refined subsequently (Moya de Guerra and Zamick, 1993) with the more general result

$$\sum B(M1) E_x(1^+) = \frac{9\chi}{16\pi} [B(E2)_{IS} - B(E2)_{IV}], (16)$$

which is indeed very close to the form of the IBM-2 result in Eq. (15), except for the additional isovector (IV) contribution. Here, χ denotes the strength of the quadrupole-quadrupole interaction and the B(E2)values are in units of e²fm⁴. For many of the transitional and definitely for the strongly deformed nuclei, the second term is very small and thus one recovers the IBM-2 result exactly. For a number of cases, though, Zamick and Zheng (1992) explicitly show the need for the isovector term in order to have the correct physics in connecting M1 and E2 electromagnetic properties. It is important to note that, similar to the collective IBM-2 formulation and the TRM treatment, the residual twobody forces may contain besides the strong quadrupole forces, pairing interactions amongst identical nucleons. The sum rule is not affected by the addition of the latter term.

Moreover, one can relate the non-energy weighted M1 sum rule to the nuclear monopole properties

(Heyde *et al.*, 1993). Starting from the IBM expression for the monopole operator

$$T(E0) = \gamma_0 \hat{n}_s + \beta_0 \hat{n}_d = \gamma_0 \hat{N} + \beta_0' \hat{n}_d, \tag{17}$$

with \hat{n}_s , \hat{n}_d , \hat{N} representing the s, d and total boson number operators, respectively, and $\beta'_0 \equiv \beta_0 - \gamma_0$, one derives the mean-square radius as

$$\langle r^2 \rangle = \gamma_0 N + \beta_0' \langle \hat{n}_d \rangle. \tag{18}$$

Thus, the M1 sum rule of Eq. (12) can be recast in the form

$$\sum B(M1) = \frac{9}{4\pi} (g_{\pi} - g_{\nu})^2 \frac{1}{\beta_0'} \frac{P}{N-1} \left[\langle r^2 \rangle - \gamma_0 N \right] (19)$$

This latter expression (see Fig. 25 for a comparison with data) shows the connection between summed M1 strength and nuclear radial properties (the latter taken from the compilation of Otten (1989)). Similar connections have also been suggested by Iachello (1981) and by Otsuka (1992), separately because of the connection between the nuclear radial variation $\Delta \langle r^2 \rangle$ for a liquid drop and the variation in quadrupole deformation $\Delta \langle \beta_2^2 \rangle$.

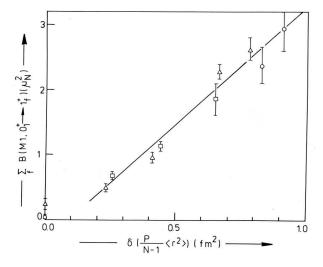


FIG. 25 Relation between the experimental summed M1 strength of all 1⁺ states below 4 MeV for $^{142,146-150}$ Nd (squares), $^{144,148-154}$ Sm (triangles), $^{160-162}$ Dy (circles) and the variation of the quantity $\delta(\frac{P}{N-1}\langle r^2\rangle)$ related to the isotope shift (Heyde et~al., 1993).

d. Deformation dependence and saturation Starting from a microscopic approach (deformed mean-field with residual pairing and quadrupole interactions and even using Hartree-Fock calculations using Skyrme effective forces (Garrido et al., 1991; Lo Iudice, 1998; Smith et al., 1995)), the very same dependence between the summed M1 strength and nuclear deformation results, as discussed before. Remarkably, it is the presence of pairing forces that modifies the dependence of the summed

M1 strength from linear to quadratic. This was shown to be the case by Hamamoto and Magnusson (1991) in a Nilsson model study, and also by Lo Iudice and Richter (1993) using an RPA study of magnetic scissors motion.

The fact that the summed M1 strength shows this striking collective behavior immediately leads to saturation because the equilibrium quadrupole deformation (in passing through the rare-earth region from A = 140towards the mass A = 180 region) stabilizes at a value of $\delta \simeq 0.25$ in the region $A \geq 160$ of strongly deformed even-even nuclei. The precise origin of this saturation stems from the specific single-particle structure in the deformed mean field and from the balancing effects of shell- and pairing corrections to the liquid-drop energy (Heyde et al., 1992). The saturation arises after a steep increase in deformation, which is reflected in a steep rise in both the summed M1 strength and in the $B(E2; 0_1^+ \rightarrow$ 2_1^+) value when entering the region of deformation, starting from closed-shell nuclei (see Fig. 22). This strong correlation between the summed M1 strength and the ground-state equilibrium quadrupole deformation has been discussed by De Coster and Heyde (1989).

The dominant role of pairing to obtain the correct deformation dependence for the magnetic summed strength is not straightforward but comes in indirectly. In a quadrupole deformed potential, the strength of M1 transitions between Nilsson orbitals, characterized by Ω (the projection of j on the symmetry axis) and illustrated by the arrows in the top part of Fig. 26, can be expressed as

$$B(M1) = \frac{3}{4\pi} (u_1 v_2 - u_2 v_1)^2 |\langle \Omega_1 | g_l \hat{l}_+ + g_s \hat{s}_+ | \Omega_2 \rangle|^2.$$
(20)

The occupation probabilities v_i^2 (with $u_i^2 = 1 - v_i^2$) of the Nilsson orbitals Ω_i are schematically drawn in the bottom part of Fig. 26 for both small and large quadrupole deformations. The doubly-hatched lines indicate the position of the Fermi level (top part) and for an occupation number 0.5 (lower part). For small deformation the pairing factor $(u_1v_2 - u_2v_1)^2$ quenches the M1 strength and vanishes for zero deformation. With increasing deformation, the Nilsson single-particle orbitals Ω_i originating from a single j shell-model orbital are more spread out and the corresponding occupation probabilities become different resulting in a rather large pairing factor. Then, as deformation does not change much for increasing mass numbers, the pairing factors remain roughly constant, causing saturation before they start to decrease again towards the end of the shell. Using the approximation that the energy of the Nilsson orbitals vary linearly with deformation for not too large values of deformation, it can even be shown that the pairing factor $(u_1v_2 - u_2v_1)^2$ becomes proportional to δ^2 . This result comes very close to what Hamamoto and Magnusson (1991) obtained too.

A particularly interesting example of this deformation dependence has resulted from the study of magnetic dipole strength in superdeformed nuclei (Hamamoto and Nazarewicz, 1992, 1994). The summed B(M1) strength was found to be much larger than in

nuclei at normal deformation. This can be understood from the growing proton orbital contribution with increasing deformation and the fact that for the weak pairing present in the superdeformed configuration, the pairing factor in Eq. (20) becomes maximal. Applications for nuclei in the proton-rich deformed Kr-Zr nuclei show similar results (Nakatsukasa et al., 1994).

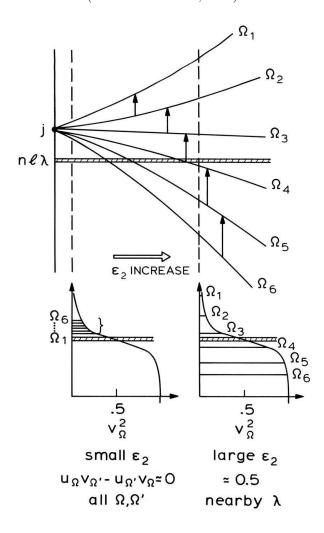


FIG. 26 Schematic representation of the effect of deformation on the orbital $(lj) \rightarrow (lj)$ transition strength. The Fermi level is denoted by the hatched lines (Richter, 1990).

Whereas QRPA calculations for strongly deformed nuclei give rise to the δ^2 dependence of the M1 strength and saturation for both the $B(E2;0^+_1 \to 2^+_1)$ and summed M1 strength due to the importance of pairing correlations amongst the interacting nucleons, the IBM does not exhibit this characteristic behavior. In both the SU(3) and in the O(6) limit of the IBM-2, the B(M1) transition into the scissors 1^+ mixed-symmetric state is proportional to the P factor (Casten et al., 1987) which implies an almost linear rise towards mid-shell before dropping off towards the end of the major shell (Scholten et al., 1985a). This sheds light on the way in which to count the boson number when highly deformed systems are being de-

scribed. Casten *et al.* (1988) have shown that one should consider in this region an effective boson number, following arguments by Otsuka *et al.* (1990), through which Pauli blocking is taken into account.

e. A comprehensive analysis Even though it was shown in the foregoing sections that one can obtain a qualitative understanding of the observed behavior of the summed magnetic dipole strength at low energies (deformation dependence, saturation, relation to other multipoles), a quantitative agreement between the large body of experimental data and theory has still been lacking. Part of the problem is related to the fact that most model approaches use still too idealized assumptions concerning the way in which nucleons behave inside the atomic nucleus: moments of inertia, gyromagnetic ratios, etc.. Therefore, a global study in heavy even-even nuclei was carried out in order to obtain as accurate a description of the scissors mode when one uses as input the realistic physical parameters in the calculations (Enders et al., 1999, 2005).

It makes use of the sum rule method as described by Lipparini and Stringari (1983, 1989b), LS in shorthand notation, and starts from the energy-weighted (S_{+1}) and the inverse energy-weighted (S_{-1}) sum-rule expressions

$$S_{+1} = E_{sc}B(M1)$$

$$= \frac{3}{20\pi}r_0^2 A^{\frac{5}{3}} \delta^2 E_D^2 \frac{m_N}{\hbar^2} (g_p - g_n)^2 \mu_N^2 \text{MeV}, (21)$$

and (cf. Eq. (10))

$$S_{-1} = \frac{B(M1)}{E_{sc}} = \frac{3}{16\pi} \frac{J_{sc}}{\hbar^2} (g_p - g_n)^2 \frac{\mu_N^2}{\text{MeV}},$$
 (22)

with $r_0=1.15$ fm, A the nuclear mass number, δ the nuclear deformation parameter, E_D the isovector giant electric dipole resonance (IVGDR) excitation energy, \mathbf{m}_N the nucleon mass and \mathbf{g}_p (\mathbf{g}_n) the g factors for protons (neutrons). With E_{sc} we denote the excitation energy of the scissors mode and J_{sc} describes the moment of inertia associated with the scissors mode vibrations, which are of isovector type. These expressions are rather general and express e.g. the fact that the scissors mode and the IVGDR are both of isovector nature and strongly related through the restoring force acting on the deformed proton and neutron bodies and also that the major contribution to S_{-1} comes from the low-lying scissors mode ($2\hbar\omega$ strength in Fig. 1) whereas the high-lying scissors mode ($2\hbar\omega$ strength in Fig. 1) mainly contributes to S_{+1} .

First we discuss S_{-1} because the low-lying scissors mode, after all, is very well studied by now in the heavy nuclei. To start with, we use the common relative g-factor values $g_{rel} = g_p - g_n = 2Z/A$ (Bohr and Mottelson, 1975), and so deduce the moment of inertia (taking the tacit assumption that the scissors M1 strength resides in the energy region 2.5-4 MeV). It is an interesting observation to see (Enders et al., 1999, 2005) that the ground-band and scissors motion moments

of inertia are very close to each other (except for a systematic deviation in heavier nuclei).

In a next step, we take these two moments of inertia to be equal², i.e. $J_{sc} = J_{gb}$, and evaluate the g factors starting from the sum-rule value of S_{-1} . There appears a striking agreement between these values with only small deviations (of the order of 10%). When recalculating the moment of inertia related to the scissors motion in Eqs. (21,22), using the experimentally deduced $g(2^+)$ values and comparing these moments with the ground-state moments of inertia, an almost perfect overlap between the two sets results. So, we can draw the conclusion that the g factors acting in the scissors mode are the same as the ones for the ground band.

Since all quantities in the sum rules (21,22) are fixed, we are now in a position to rederive the energies and the strengths for the scissors modes. However, when dealing with S_{+1} one has to take into account that contributions from the K=1 component of the isovector giant quadrupole resonance (IVGQR) will dominate, which have to be removed in order to compare with the experimental data. This is achieved with a procedure described by Lipparini and Stringari (1983), which leads to a correction factor $\xi = E_Q^2/(E_Q^2 + 2E_D^2)$ to Eq. (21) where E_Q denotes the centroid energy of the isoscalar giant quadrupole resonance (ISGQR).

When putting all low-lying (high-lying) M1 strength $\sum_{low} B(M1) = B_l \ (\sum_{high} B(M1) = B_h)$ into a single state with energy $E_l(E_h)$, one obtains the relations

$$S_{-1} \approx \frac{B_l}{E_l}, \quad S_{+1} \approx B_h \cdot E_h,$$
 (23)

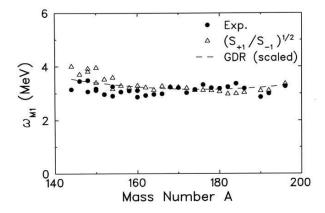
and using the expressions for S_{-1} and S_{+1} , we can derive an average energy

$$\bar{\omega} = \frac{2}{\sqrt{15}} \sqrt{\frac{m_N}{\hbar^2}} r_0 \sqrt{\frac{4NZ}{A^2}} A^{5/6} E_D \sqrt{E_{2+}\xi} \delta, \qquad (24)$$

where use was made of the relation between scissors mode and ground-state band moment of inertias established above and the latter is expressed through the energy of the first excited state of the rotational band E_{2+} by $J_{gb} = 3\hbar^2/E_{2+}$. The centroid energies of the IVGDR and ISGQR are taken from mass-dependent systematics (Harakeh and van der Woude, 2001).

The values resulting from Eq. (24) are drawn as triangles in Fig. 27 (upper part). Excellent agreement with the experimental data is obtained. When scaled by the ratio of the scissors-mode to the liquid-drop moment of inertia, the energy of the isovector giant dipole resonance (dashed line) also shows the proportionality to the energy of the scissors mode predicted in Eq. (24). We note that

the observed near constancy of ω could also be explained by Pietralla *et al.* (1998b) within a schematic RPA approach after inclusion of the deformation dependence of paring effects analogous to the discussion in the previous subsection.



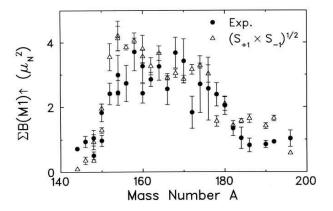


FIG. 27 Excitation energy (upper part) and transition strength systematics (lower part) from a sum-rule analysis of the scissors mode in rare-earth nuclei. Experimental values (solid circles) and parameter-free predictions (open triangles) are shown for the mean excitation energy (upper part) and the summed M1 strength (lower part). The deformation dependence of the moment of inertia leads to the proportionality of the excitation energies of the scissors mode and the IVGDR as indicated by the dashed line (Enders $et\ al.$, 2005).

In the same way, one obtains for the low-lying scissors strength

$$B_l = \frac{3}{\pi} \sqrt{\frac{3}{20}} r_0 \sqrt{\frac{4NZ}{A^2}} A^{5/6} E_D \sqrt{\frac{m_N \xi}{\hbar^2 E_{2+}}} \delta g_{gb}^2.$$
 (25)

with $g_{gb} = \frac{1}{2}(g_p + g_n)$, the g factor of the 2^+ level of the ground-state band. The comparison to the experimental scissors mode strength is depicted in the lower part of Fig. 27. The agreement is very satisfactory except for some nuclei with A > 180. The strong deformation dependence is generated by the interplay of E_{2+}

 $^{^2}$ More precisely the moment of inertia of isoscalar and isovector motion differ by a factor $4NZ/A^2 \simeq 0.96$ (Lo Iudice and Richter, 1993).

and δ . Indeed the experimentally established quadratic dependence of the scissors-mode strength on the ground-state deformation is easily derived from Eq. (25) recalling (Bohr and Mottelson, 1975) that the moment of inertia J_{gb} is roughly proportional (albeit much larger) to the superfluid moment of inertia J_{lig}

$$J_{gb} \propto J_{liq} \simeq J_{rig} \delta^2,$$
 (26)

where J_{rig} stands for the moment of inertia of a rigid rotor.

f. Sum rule relation between magnetic dipole and octupole strength In the spirit of the IBM-2 non-energy weighted M1 sum rule (Ginocchio, 1991), a relation between magnetic dipole and octupole strength could be derived starting from energy-weighted sum rules (Heyde et al., 1994). It has been possible to obtain an approximate, yet very simple relation

$$\frac{\sum_{f} B(M1; 0_{1}^{+} \to 1_{f}^{+}) E_{x}(1_{f}^{+})}{\sum_{f} B(M3; 0_{1}^{+} \to 3_{f}^{+}) E_{x}(3_{f}^{+})} \cong \left[\frac{6(g_{\pi} - g_{\nu})}{7(\Omega_{\pi} - \Omega_{\nu})}\right]^{2}. \tag{27}$$

This relation establishes a link between different parameters of the IBM-2, the gyromagnetic boson factors and octupole boson moments, and hence imposes constraints on choosing them when fitting to other spectroscopic data. One can, on the other hand, also use the relation starting from common values for these parameters $(g_{\pi}, g_{\nu}, \Omega_{\pi}, \Omega_{\nu})$ as derived from a phenomenological and/or microscopic starting point and deduce an estimate for the summed M3 strength, whenever information on the summed M1 dipole strength is available. This has been discussed by De Coster et al. (1995), where an estimate of the summed M3 strength is presented in the mass region 144 < A < 164. Even though it contains a number of approximations, in the absence of systematics on the M3 strength, the above method might be a first guide for further experimental studies.

There are very few studies on M3 transitions carried out by now, both theoretically and experimentally. On the theoretical side, M3 transitions were investigated within the framework of the IBM-2 by Scholten *et al.* (1984) and within the context of a schematic RPA study for heavy deformed nuclei by Lo Iudice (1988). There has been an early experimental search for M3 strength in 164 Dy by Bohle *et al.* (1987b) that made use of both the electron accelerators at Darmstadt and Amsterdam. Only an upper limit for such strength has been derived. A more systematic search for magnetic octupole strength is called for in the light also of the above approximate connection with the summed M1 strength.

B. Spin-flip mode: experimental evidence and theoretical description

1. Qualitative nature of the magnetic dipole response

The magnetic dipole operator, for a system of protons and neutrons reads

$$T(M1) = \sqrt{\frac{3}{4\pi}} \sum_{i} \left\{ g_l(i)\hat{l}_i + g_s(i)\hat{s}_i \right\} \mu_N, \qquad (28)$$

with the usual orbital and spin g_l, g_s factors for neutrons and protons (see also Eq. (7)). Using the isospin labels $t_z(i) = \pm \frac{1}{2}$ for neutron and proton, respectively, the magnetic dipole operator can be split into an isoscalar and isovector term in the following way

$$T(M1) = \sqrt{\frac{3}{4\pi}} (g_J \hat{J} + g_S \hat{S}) \mu_N + T(M1, IV).$$
 (29)

Since \hat{J} denotes the total angular momentum operator, this term does not induce any M1 transitions and because $g_S = [(g_s^{\pi} + g_s^{\nu}) - 1]/2$, the isoscalar spin part only contributes in a minor way to M1 transitions. This is a consequence of the opposite signs in the proton and neutron spin g_s factors resulting in a value of $g_S = 0.38$.

The isovector part of the M1 operator T(M1, IV)

$$T(M1, IV) = \sqrt{\frac{3}{4\pi}} \left\{ \frac{1}{2} (\hat{L}_{\pi} - \hat{L}_{\nu}) + \frac{1}{2} (g_s^{\pi} - g_s^{\nu}) (\hat{S}_{\pi} - \hat{S}_{\nu}) \right\} \mu_N$$
(30)

splits into two pieces: the first part, describing the relative angular momentum between protons and neutrons, generates the scissors orbital motion whereas the second part, a spin-flip part, is nothing else but the ΔT_z =0 component of the Gamow-Teller operator. This term can strongly enhance spin-flip M1 transitions because of the large factor $\frac{1}{2}(g_s^\pi - g_s^\nu)$ in front (with a numerical value 4.72 using free g_s factors).

That the simple picture of the nuclear magnetic dipole response is approximately correct is shown in Fig. 28 by using the three nuclei 56 Fe, 156 Gd and 238 U as examples. As pointed out above, the mean excitation energy of the orbital mode scales approximately with deformation as $E_x \simeq 66\delta A^{-1/3}$ MeV. The spin M1 strength obtained in inelastic proton scattering (Frekers et~al., 1990) lies at $E_x \simeq 41A^{-1/3}$ MeV and thus exhibits a shell-model like excitation energy dependence. As seen in Fig. 28, the spin strength represents the largest fraction of the M1 strength.

It is the residual particle-hole interaction acting in the spin-isospin channel that shifts this spin M1 strength to higher energies as can be seen e.g. in schematic model (De Coster and Heyde, 1991a; De Coster et al., 1992; Zawischa et al., 1990; Zawischa and Speth, 1994) and in RPA studies (Sarriguren et al., 1996, 1993, 1994).

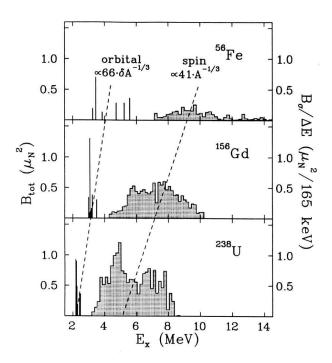


FIG. 28 The nuclear orbital and spin magnetic dipole response in a medium-heavy, a heavy and a very heavy nucleus, derived from experiments with electromagnetic and hadronic probes, respectively (Richter, 1994).

The orbital part of the M1 strength is hardly moved by this force component and so one expects to observe in the experiments a rather good overall separation of the orbital dipole magnetic excitations, at the lower energy end of 2.5-4 MeV, from the higher-lying spin magnetic dipole excitations.

The experimental detection of spin strength in the energy region above 4 MeV needs a probe that is particularly sensitive to the spin part of the nuclear current. Intermediate-energy scattering of (polarized) protons at small forward angles should be the optimal selective reaction to carry out such a search. First experiments performed by a Darmstadt/Münster/TRIUMF collaboration on ¹⁵⁴Sm, ¹⁵⁸Gd, and ¹⁶⁸Er used 200 MeV protons at an angle of 3.4° covering final states up to 12 MeV. Analyzing those data clearly showed in all nuclei the presence of extra strength sitting on the tail of the IVGDR (Fig. 29) with a double-hump structure (Frekers et al., 1990; Richter, 1995). The double-hump has centroids around 6 and 8.5 MeV and widths of about 1.5 and 2 MeV, respectively. Even more detailed substructure becomes visible (see insert in Fig. 29). Such a pronounced splitting and fragmentation of magnetic dipole strength has not been observed as yet in spherical nuclei. The selectivity of the (p, p') reaction could be demonstrated at the same time: the orbital M1 tranition at 3.19 MeV strongly excited in (e, e') and (γ, γ') reactions was not observed with an upper limit $B(M1) < 0.1 \mu_N^2$.

In order to obtain an even more complete insight in

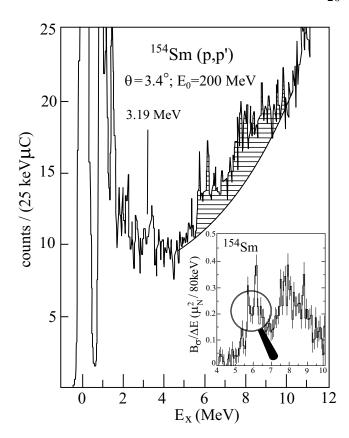


FIG. 29 Forward angle inelastic proton scattering spectrum taken at 200 MeV incident energy on 154 Sm. The shaded area constitutes the spin magnetic dipole giant resonance. In the insert, the extracted B(M1) strength (in units $\mu_N^2/80$ keV) distribution is shown (adapted from Frekers *et al.* (1990)).

the structure of these excitations, angular distributions have been taken for these three nuclei (Fig. 30). No strong Z nor A dependence shows up. A DWBA fit for a $\Delta S=1$, $\Delta L=0$ transition considering a neutron spin-flip $1h_{11/2} \rightarrow 1h_{9/2}$ or a proton $1g_{9/2} \rightarrow 1g_{7/2}$ transition has been performed. These orbitals are clearly the dominant ones in this mass region and for deformed nuclei; even the Nilsson states are dominated by these particular spherical components. A strength $B(M1)=10.5(2.0)~\mu_N^2$ could be extracted, a value in line with expectations for the theoretical spin-flip strength (see e.g. Fig. 17).

Complementary experiments detecting the transverse spin-flip probability S_{nn} have been carried out (Wörtche, 1994). Thereby the probability that an incoming proton, interacting with the target nucleus, will leave with its spin flipped ($\Delta S=1$ process) is measured. The results for ¹⁵⁴Sm are displayed in Fig. 31 where, besides the cross-section in the energy interval 4-32 MeV, the corresponding spin-flip probability S_{nn} is given. Here one notices the presence of increased S_{nn} values in the region of the observed spin M1 strength (6 to 8.5 MeV), confirming the spin-flip character of the structures located on the low-energy tail of the GDR, the latter be-

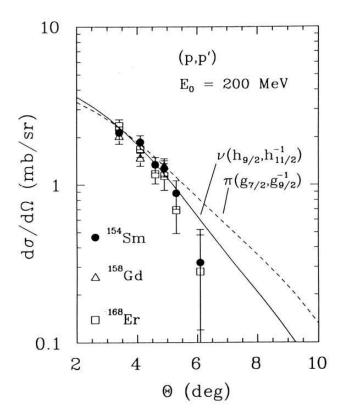


FIG. 30 Angular distribution of the summed double-humped structures observed in ¹⁵⁴Sm, ¹⁵⁸Gd, and ¹⁶⁸Er. The dashed and full curves result from DWBA calculations based on the indicated particle-hole excitations (Richter, 1991).

ing split in two fragments, due to nuclear deformation. This figure also shows the presence of an isovector giant quadrupole resonance (IVGQR) derived from a multipole decomposition around 23 MeV, an excitation energy where it is has also been detected in other heavy nuclei (Harakeh and van der Woude, 2001). The analysis of the S_{nn} values is carried out consistently with the DWBA analysis of the angular distributions.

In recent years it has become clear that the E1 response in heavy nuclei generally exhibits a local concentration of strength - called the pygmy dipole resonance (PDR) - well below the IVGDR overlapping with the excitation energy region of the spin-flip resonance. The PDR has been observed at a variety of shell closures (see e.g. Kneissl et al. (2006) and references therein) but not yet in heavy deformed nuclei. This raises the question whether part of the strength attributed to the spin-flip M1 resonance is in fact of E1 nature. While this problem needs further experimental investigation, there are immediately two arguments in favor of the present interpretation: the angular distributions shown in Fig. 30 are distinct from those of Coulomb-excited E1 transitions, and the S_{nn} values in the bottom part of Fig. 31 display a local maximum of the spin-flip strength.

Before concentrating on the theoretical description, we

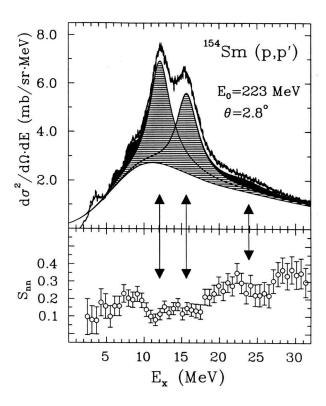


FIG. 31 Differential cross-section and transverse spin-flip probability for inelastic polarized proton scattering on $^{154}\mathrm{Sm}$. The hatched areas show the double-humped GDR. Visible on the low-energy side of the the GDR is the spin-flip M1 resonance between 5-12 MeV excitation energy and at $E_x=23.4$ MeV for the IVGQR. The arrows visualize the connection between the electric resonances and dips in the spin-flip probability (Richter, 1995).

show finally the M1 response for a set of nuclei spanning a wide region of deformed rare-earth nuclei (Fig. 32). In all of these nuclei, a particularly stable pattern is emerging: at the lower energy side, at energies 2.5-4 MeV, a concentration of orbital magnetic dipole strength shows up with a ratio $\sqrt{B_l/B_\sigma} \approx 4$, where B_l and B_σ denote the reduced transition strength of the orbital and spin part of the magnetic dipole operator (Eq. 30) and $B(M1) = (\sqrt{B_l} \pm \sqrt{B_\sigma})^2$ (Willis et al., 1989). Higher up, starting at 5.5 MeV up to almost 10 MeV, a rather broad and extended region with a clear double-hump structure in most of these nuclei appears in which $\sqrt{B_l/B_\sigma} \leq 1$. Figure 32 therefore reflects the magnetic dipole response for strongly deformed nuclei in the rare-earth region to electromagnetic and hadronic probes. The low-energy part corresponds to an orbital magnetic dipole structure, the scissors mode, and the higher part is the spin-flip part mainly caused by proton and neutron single-particle transitions between spin-orbit partners. Whereas the energy of the spin-flip M1 mode will be localized at the energy of the gap in closed shells for spherical nuclei, in the region where deformation sets in one expects splitting of the various Nilsson energy levels causing a spreading

of the spin-flip strength around the spherical centroid energy. If this argument is correct, one should obtain an $A^{-1/3}$ excitation-energy dependence for the observed peaks of M1 strength throughout the whole mass region (cf. Fig. 28). This seems indeed the case and is demonstrated in Fig. 33, in which the double-hump M1 spin strength in deformed rare-earth and actinide nuclei is connected (Richter, 1995) to the detailed knowledge of the spin-flip strength in the doubly-magic ²⁰⁸Pb nucleus (Laszewski et al., 1988). Indications of a similar splitting is observed in medium-mass nuclei (Djalali et al., 1982) except for the stable Zr isotopes (Crawley et al., 1982), where the forward-angle (p, p') cross sections exhibit a single bump at $E_x \approx 9$ MeV identified as a spin-flip M1 resonance. This may be related to the special shell structure at Z = 40, where the high-j orbital near the Fermi surface $(\pi 1g_{9/2})$ is essentially unoccupied and the corresponding $1g_{9/2} \rightarrow 1g_{7/2}$ transition suppressed.

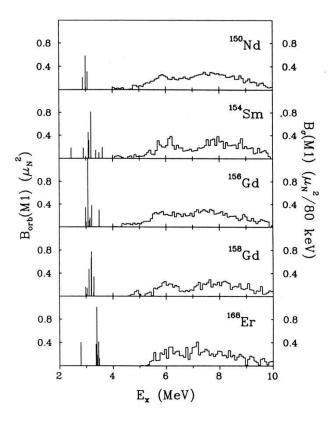


FIG. 32 Magnetic dipole response of several deformed rareearth nuclei determined by inelastic electron, photon and proton scattering (Richter, 1995).

2. Theoretical description

When trying to study the systematics of centroid energy and strength of the spin-flip transitions in rareearth nuclei, a first approximation is to look at the unperturbed M1 strength originating from a deformed

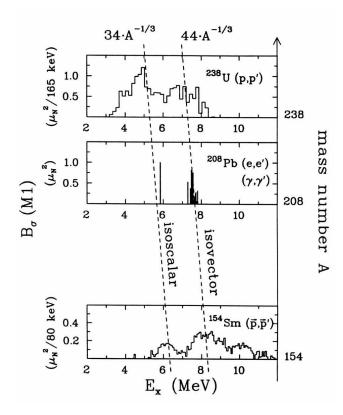


FIG. 33 Spin magnetic dipole strength distributions in $^{238}\mathrm{U},^{208}\mathrm{Pb}$ and $^{154}\mathrm{Sm}.$ The center of gravity of the excitation energy of the two peaks representing the main strength follows a simple $A^{-1/3}$ law, characteristic for spin-flip excitations between spin-orbit partners. The experimental strength distribution for $^{208}\mathrm{Pb}$ has been combined from inelastic electron and photon scattering experiments (Richter, 1995).

single-particle model. In carrying out this procedure, De Coster and Heyde (1991a) have studied the summed spin M1 strength throughout the whole rare-earth region from 140 Ce up to 198 Pt. The strongest values are obtained at the end of the major shell near Z=82 and N=126 through proton $1h_{11/2} \rightarrow 1h_{9/2}$ and neutron $1i_{13/2} \rightarrow 1i_{11/2}$ transitions. The unperturbed energy of these transitions is situated in the energy region 4-10 MeV. With the residual interaction switched on, the M1 strength will be redistributed but the total strength should not change much from the unperturbed case.

A comparison of the experimental strength distribution in ¹⁵⁴Sm with a number of QRPA and QTDA studies has been carried out (De Coster and Heyde, 1991a; Hilton et al., 1998; Sarriguren et al., 1993; Zawischa et al., 1990; Zawischa and Speth, 1990), see Fig. 34. The theoretical results have been folded with a Gaussian of variable width in order to facilitate comparison. One has to conclude that the agreement between experiment and theory is still on a qualitative level. The position of the two peaks does not vary a lot but the relative strength of the peaks is changing in

a rather important way pointing out the sensitivity of the calculations to both the underlying single-particle structure as well as to the residual interactions used. The QRPA calculations that come closer to the data (Hilton et al., 1998; Sarriguren et al., 1993) have been carried out after the experiments were performed. One also notices that the difference between QRPA and QTDA (De Coster and Heyde, 1991a; Zawischa et al., 1990; Zawischa and Speth, 1990) are not dramatic pointing out that ground-state correlations do not seem to play a major role in determining both the energy and the strength for these spin-flip transitions.

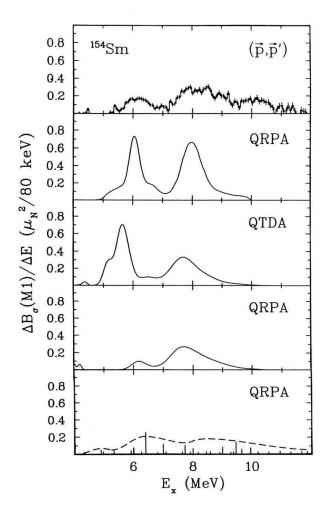


FIG. 34 Experimental and theoretical spin magnetic dipole strength distributions in ¹⁵⁴Sm. Underneath the experimental data, theoretical predictions from various calculations are given (in descending order): QRPA from Zawischa et al. (1990); Zawischa and Speth (1990), QTDA from De Coster and Heyde (1991a); De Coster et al. (1992), QRPA from Sarriguren et al. (1993), and QRPA from Hilton et al. (1998).

The specific double structure of the strength distributions is related to the residual spin-spin interaction which changes the unperturbed picture in an important way (De Coster *et al.*, 1992). Besides a shift of the spin

strength to higher energies, as expected from schematic p-h models studying isovector excitations, a rather clear separation into a proton-like and a neutron-like collective spin mode remains. As a result, in $^{154}\mathrm{Sm}$, the lower peak mainly originates from the proton $1g_{9/2} \rightarrow 1g_{7/2}$ and the $1h_{11/2} \rightarrow 1h_{9/2}$ excitations whereas the second, higher-lying peak is mainly due to the $1h_{11/2} \rightarrow 1h_{9/2}$ and the $1i_{13/12} \rightarrow 1i_{11/2}$ spherical components. These simple p-h configurations act as doorway states for the fragmentation of the resonance in analogy to the discussion for the scissors mode (cf. Fig. 20).

For larger strengths of the spin-spin proton-neutron interaction as used by Sarriguren et al. (1996, 1993, 1994), the spin strength becomes concentrated more into a full isoscalar and isovector part with proton and neutron configurations strongly mixed. Zawischa and Speth (1990) obtain results somewhat intermediate between the two more extreme cases of very weak coupling and strong coupling between the individual proton and neutron spinflip M1 configurations. The higher peak shows a structure that is reminiscent of a genuine giant spin-flip (or Gamow-Teller) mode of isovector character. In comparing both the incoherent sum of the separate proton and neutron contributions with the actual calculation where interference effects do play an important role, it seems like the lower part is mainly of proton character but also an isoscalar part is present. They come to the conclusion that for the higher peak in ¹⁵⁴Sm, using a Landau-Migdal residual interaction, the neutron contributions play the dominant role and come close to the results of De Coster and Heyde (1991a).

It is beyond discussion that the starting points, i.e. different single-particle deformed potentials (Nilsson, deformed Woods-Saxon, deformed Hartree-Fock mean field) and different residual interactions, lead to results that differ in an important way in their interpretation of the nature of the double-peak structure (proton and neutron vs. isoscalar and isovector), see also Lipparini and Richter (1984). Experiments that are sensitive to the proton-to-neutron content in exciting those states - like inelastic π^{\pm} scattering - can most probably solve this issue and, at the same time, give invaluable information concerning the proton-neutron part of the spin-spin component in the effective residual two-body interaction.

In conclusion, the general structure and evolution of spin-flip M1 strength can be studied using schematic models too (De Coster and Heyde, 1991b; De Coster $et\ al.$, 1992; Zawischa and Speth, 1990, 1994) and these results are in general consistent with those from the more detailed QRPA studies. For not too strong spin-spin proton-neutron coupling, while considering a two-level model (or a four level model) in the rare-earth region, the strength becomes concentrated, separately, into a pure proton and neutron collective spin-flip state. With increasing strength, all components eventually contribute into an isovector mode at the higher energy and an isoscalar part at the lower energy side, albeit with the

neutron configurations and proton configurations dominating in these two modes, respectively.

C. Magnetic dipole strength at higher excitation energy: prediction and experimental hints

Experiments have succeeded in studying the response of the nucleus to medium energy protons towards much higher excitation energies. This was illustrated already in Fig. 31 in which the cross section and the S_{nn} transverse spin-flip probability in ¹⁵⁴Sm up to an excitation energy of 32 MeV has been shown. Besides the dominant giant electric dipole resonance state, split by deformation, on the lower side, the spin-flip M1 strength has been detected and is discussed in detail in Sec. III.B. On the higher energy tail though, excess strength is observed, which can be described by a Lorentzian centered at $E_x = 23.4 \text{ MeV}$ with a width of 6.8(6) MeV. These parameters as well as an exhaustion of the corresponding energy-weighted sum rule of 76(11)\% agree well with the (scarce) systematics (Harakeh and van der Woude, 2001) of the isovector giant electric quadrupole resonance (IVGQR). The arrows indicate dips in the S_{nn} behaviour which are directly related to the electric character of the strong states as compared to the other spin magnetic excitations.

This is probably the $K^{\pi}=1^+$ component of the IVGQR, which is split into various K components for strongly deformed nuclei, and taken as the genuine manifestation of a classical scissors motion. Lo Iudice and Richter (1989) have pointed out that the lower RPA 1⁺ solution does not collect the whole M1 strength. They have shown that a non-negligible fraction is obtained at higher energy. For A=164 (N=Z), they obtain a value of 25 MeV for the excitation energy of a high-energy mode with a corresponding strength of $B(M1)=4.5~\mu_N^2$. The quadrupole component then acquires, through the M1 transition, an additional scissors characteristic.

As was discussed in Sect. III.A.3, making a comprehensive analysis of M1 excitations in atomic nuclei, a relation between the energy and strength of a low- and highenergy scissors state was indicated i.e. $\omega_h B_h = 4\omega_l B_l$ which gives interesting information on both the expected excitation energy and the M1 strength for a high-energy mode. Such a mode will be mainly built out of $2\hbar\omega$ quasiparticle excitations for which the M1 strength becomes concentrated into a single strong state above 20 MeV. The issue of how well such a strong state at that high energy will remain intact is not clearly settled. A number of calculations (Hamamoto and Nazarewicz, 1992, 1994; Nojarov et al., 1995; Zawischa, 1998; Zawischa and Speth, 1994), using schematic or more realistic forces come to different conclusions. No jarov et al. (1995) and Hamamoto and Nazarewicz (1992) obtain a large concentration in a strong peak if they truncate the 2gp model space up to 20 MeV, in line with calculations

within a schematic picture in which the higher-lying M1 strength also remains concentrated (see Fig. 35). Using a more extended space spreads out this strength considerably. Anyhow, at this energy, the resonance is highly unbound which will induce further spreading making it a difficult task to unambiguously detect and measure the amount of M1 strength. Thus, in view of the largely different theoretical results, the question of a still observable concentrated strong state will be difficult to solve. Moreover, as illustrated in Fig. 31, there are considerable experimental difficulties arising from a strong background mainly due to quasifree scattering, whose exact shape and strength are unknown and had be approximated by a semiempirical approach (Lisantti $et\ al.$, 1984).

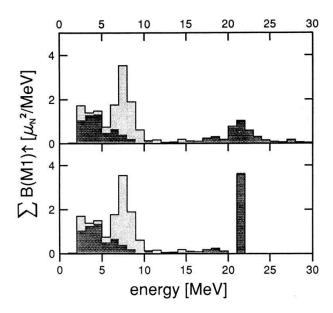


FIG. 35 Prediction of the full M1 strength distribution (summed in bins of 1 MeV) including the high-energy part in 160 Gd. The dark-hatched zone represents spin strength, the light-hatched one orbital strength. In the lower part, only 2qp configurations up to 20 Mev are taken into account. In the upper part, a much higher energy cut-off for these 2qp configurations is imposed. (Reprinted with permission from Nojarov et al. (1995). ©(2009) Am. Phys. Soc.)

The identification of the IVGQR permits an interesting test of the energy-weighted M1 sum rule given in Eq. (16) relating the summed M1 strength on the l.h.s. to the difference of the isoscalar and isovector E2 summed strengths on the r.h.s. of Eq. (16). With the experimental numbers for 154 Sm (Ziegler et~al., 1993) one obtains for the energy-weighted M1 sum rule a value of $7.71 \pm 0.44~\mu_N^2$ MeV whereas the right-hand part (the difference of summed E2 strength) becomes $9.32 \pm 0.31~\mu_N^2$ MeV. This is a rather good indication that below 4 MeV - the region where the sum of M1 strength was carried out - the M1 sum rule is exhausted already by 80%, leaving room for about 20% at the high energy part of M1 strength.

IV. MAGNETIC DIPOLE EXCITATIONS IN HEAVY ODD-MASS NUCLEI

A. Experimental results and systematics

Naturally, the issue of what will happen in odd-mass nuclei when a single nucleon (proton or neutron) is coupled to the scissors mode in the even-even underlying core system comes up. From the concept of particle-core coupling and considering the low-lying isoscalar quadrupole and octupole vibrational excitations in spherical and transitional nuclei, ample evidence for all states and a sharing of electromagnetic strength amongst the particlecore coupled multiplet members has been shown (Bohr and Mottelson, 1975). In the situation where the scissors mode M1 strength at the low-energy region is already spread out over a large energy span (2.5 - 4 MeV), it is clear that an experimental verification of (i) the presence of particle(hole)-scissors coupled configurations, and more compelling, (ii) if the total M1 strength summed up in the appropriate energy interval in the odd-mass nucleus is consistent with the observed M1 strength in the even-even adjacent nuclei, will not be easy.

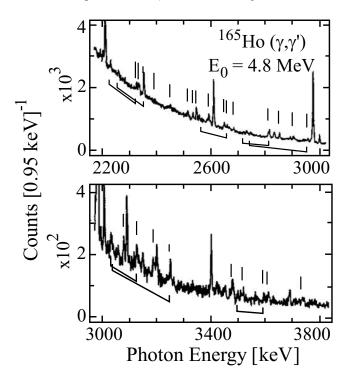


FIG. 36 Spectra of the $^{165}\text{Ho}(\gamma,\gamma')$ reaction in the energy region $E_{\gamma}=2.2-3.8$ MeV observed with a Euroball Cluster detector placed under 130° with respect to the incident beam with an endpoint energy of 4.8 MeV. Only the strongest transitions assigned to ^{165}Ho are marked with lines; brackets connecting two peaks indicate decay branches to low-lying excited states. Other transitions are due to the calibration standard ^{27}Al or result from background sources (Huxel *et al.*, 1999).

The problem is clearly one of detecting all the M1 strength, in particular the M1 strength residing in the

background of many and complex configurations. On the other hand, also the challenge for a good description from the theoretical side is not an easy one too. From a more phenomenological approach and taking the core M1 strength to be concentrated in one state, one will clearly not be able to correctly reproduce the strong fragmentation, however, the summed strength puts a constraint on this kind of model studies. From a more microscopic approach, odd-mass nuclei can be studied using a Quasiparticle-Phonon Nuclear Model (QPNM) (Soloviev, 1992). Here, one needs to take into account the fact that the QRPA phonons themselves are partly constructed from the quasi-particle configurations one is coupling to. This so-called Pauli blocking has been treated for deformed nuclei (Soloviev, 1992). For spherical oddmass nuclei, a detailed study of the transition strength from core-coupled configurations provided quantitative evidence for Pauli blocking (Scheck et al., 2008). Ultimately, one aims at exact shell-model calculations but for the strongly deformed rare-earth region, this is at present outside reach.

The experimental work covering a large part of the deformed rare-earth region has mainly been carried out by groups at Stuttgart, Köln and Darmstadt, using inelastic photon scattering through the excited states in these oddmass nuclei. The first search in the odd-proton ¹⁶⁵Ho nucleus, using photon scattering with an endpoint of about 2.5 MeV did result in appreciable amounts of M1strength (Huxel et al., 1992) albeit to be associated with transitions amongst single one-quasiparticle proton excitations in this particular nucleus. Partly due to the low endpoint energy, no clear evidence for the presence of M1strength into the mixed-symmetric configurations, to be expected beyond 2.5 MeV, was detected. Using a higher endpoint energy of 4.8 MeV and an EUROBALL cluster module (von Neumann-Cosel, 1997), spectra of much higher quality could be measured at the S-DALINAC (Huxel et al., 1999) and we give an illustration for the case of ¹⁶⁵Ho (Fig. 36).

The first and quite clear indication for such M1 excitations was obtained in the ¹⁶³Dy odd-neutron nucleus (Bauske et al., 1993). Here, a concentration of M1 strength near to 3 MeV excitation energy was detected and the summed strength fits with the observed M1 strength in the neigboring even-even $^{162,164}\mathrm{Dy}$ nuclei. Subsequent experiments on ¹⁶¹Dy and ¹⁵⁷Gd (Margraf et al., 1995), however, showed a huge fragmentation of strength in the latter nucleus. This is quite difficult to understand in the light of the proximity to ¹⁶¹Dy (just a difference of 2 protons and 2 neutrons). Further experiments on some key nuclei in order to build up systematics in this region of the rare-earth nuclei were performed on ¹⁵⁵Gd and ¹⁵⁹Tb (the latter an odd-proton nucleus), and on the heavier ¹⁶⁷Er (Schlegel *et al.*, 1996) and ¹⁶⁵Ho, ¹⁶⁹Tm (Huxel et al., 1999) nuclei.

Bringing these data together in Fig. 37, one observes that starting from $^{155}\mathrm{Gd}$, passing over the odd-mass Dy nuclei and progressing towards heavier nuclei, M1

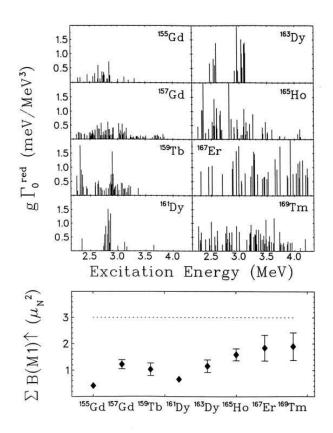


FIG. 37 Comparison of the magnetic dipole strength in the odd-mass rare-earth nuclei (Enders et~al., 1997). Upper part: distribution of reduced ground-state decay widths. Lower part: summed B(M1) strengths assuming magnetic dipole character of all observed transitions in the energy range between 2.5 and 3.7 MeV. Large differences in total strength, fragmentation and the number of detected ground-state transitions is observed. Data are from Bauske et~al.~(1993); Huxel et~al.~(1999); Margraf et~al.~(1995); Nord et~al.~(1996); Schlegel et~al.~(1996).

strength seemed to become more concentrated, precisely in those regions that were expected from the knowledge of the M1 scissors mode strength in the nearby even-even nuclei, (Enders et al., 1997). In Fig. 38 the full systematics of the Gd nuclei combining mass-even and -odd isotopes is shown (Kneissl et al., 1996; Nord et al., 1996). More recent experimental data on 151,153 Eu, and with increased sensitivity, on 163 Dy and 165 Ho, have been provided by Nord et al. (2003). Conclusions from all these data are (i) the fragmentation pattern is at best rather badly understood, and (ii) even worse, only about half to one-third of the M1 strength observed in the even-even nuclei (when summing the strength in the odd-mass nuclei in the interval 2.5-3.7 MeV) could be detected firmly (see lower part of Fig. 37).

One notable exception is the study of ¹⁶⁷Er (Schlegel *et al.*, 1996), where experiments have been carried out with endpoint energies going up to 5.8 MeV at the S-DALINAC. The summed strength reaches

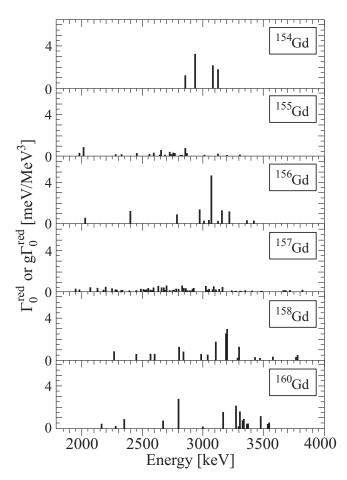


FIG. 38 Ground-state decay width distributions for 154,155,156,157,158,160 Gd extracted from photon scattering experiments. For the even-even nuclei all $\Delta K=1$ transitions are shown. In a number of cases, the M1 character has been determined by Compton polarimetry. Information from (e,e') form factor measurements has also been considered (Reprinted with permission from Nord $et\ al.\ (1996)$. ©(2009) Am. Phys. Soc.)

3.49(1.15) μ_N^2 , a value that is at variance with many of the former experiments by a factor of about three. This experiment gave a first hint that one needed to look for M1 fragments at energies higher than was first thought. As we shall discuss below, the calculated M1 strength using the interacting-boson fermion model (Iachello and Van Isacker, 1991) very well accounts for this total strength although in the theoretical study only two major peaks are obtained below 4 MeV.

B. Missing strength: experimental problem and its solution

The comparison between odd-mass and even-even nuclei immediately poses the question: where has the M1 strength gone in the odd-mass nuclei? The search was on for the observation of a large part of M1 strength residing in a very large number of complex states but with very small B(M1) values and hidden in the background of the

spectra. A detailed statistical model analysis of the highquality data in ¹⁶⁵Ho and ¹⁶⁹Tm, obtained with an EU-ROBALL cluster module, indeed revealed that a significant part of the M1 strength is carried by the background states (Enders et al., 1998, 1997; Huxel et al., 1999). It has been shown (Enders et al., 1997) that the statistical assumptions underlying the fluctuation analysis approach are also capable to explain the large variations in the measured dipole distributions seen in Fig. 37. Monte-Carlo distributions have been generated taking into account the properties of M1 and E1 distributions in the even-even neighboring nuclei and allowing for the energy dependence of the experimental sensitivity limits. These results are presented in Fig. 39 and compared to the data (Enders et al., 1997; Huxel et al., 1999). Overall, the large variations of the total number of observed levels and the summed dipole strengths can be simultaneously reproduced in a very satisfactory manner.

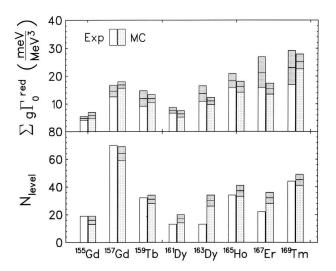


FIG. 39 Number of ground-state dipole transitions and their summed reduced width in deformed odd-mass rare-earth nuclei. Open bars refer to the data. Dotted bars stand for Monte-Carlo generated strength distributions based on a statistical approach. Uncertainties of the random spectra have been estimated by 1000-fold repetition of the calculation. Error bars denote a 1σ deviation (Huxel et al., 1999).

Subsequent experiments with unrivaled sensitivity confirmed these results for the cases of ¹⁶³Dy and ¹⁶⁵Ho (Nord et al., 2003). The experiments showed a wealth of previously unresolved weak transition as demonstrated in Fig. 40, where the strength distribution deduced for ¹⁶³Dy is compared to the first measurement (Bauske et al., 1993). The sum of the reduced dipole strength is roughly doubled. However, the fragmentation pattern into a few rather strong and many very weak transitions in ¹⁶³Dy seems to be peculiar, since a fluctuation analysis cannot explain the still missing strength. On the other hand, for ¹⁶⁵Ho good agreement with the previous work (Huxel et al., 1999) was obtained when combining the strength of resolved and unresolved tran-

sitions. It may also be noted that a very recent NRF measurement on 235 U also shows good agreement of the total M1 strength deduced from a comparable statistical analysis with that in the even-mass neighbor 236 U (Yevetska *et al.*, 2010).

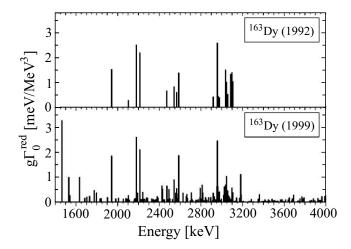


FIG. 40 Ground-state decay width distributions for 163 Dy from Bauske *et al.* (1993) and Nord *et al.* (2003). The latter experiment had an order-of-magnitude improved sensitivity. While transition strength agree well for prominent excitations, many previously unresolved weak transitions are visible.

So, as a conclusion for the present-day situation on scissors states and scissors M1 strength in odd-mass nuclei one can say that in the deformed odd-mass rare-earth (and probably also actinide) nuclei, the mode seems to be present with a strength expected from the even-even systematics but a significant part - which can change quite importantly from nucleus to nucleus depending on the respective level densities and the photon scattering endpoint energy - escapes detection in the photon scattering experiments because of the very large fragmentation.

C. Theoretical description

As remarked above, coupling an odd particle or hole (proton or neutron) to the collective modes of the eveneven core nucleus generally results in the observation of core-coupled multiplets (Bohr and Mottelson, 1975). Because of the subsequent fragmentation of M1 scissors strength, it is clear that in comparing theoretical results with data, at most indications for the total summed M1 strength will be the guiding principle to judge the level of agreement.

Within the context of the interacting boson-fermion model (IBFM), a sum rule has been derived by Ginocchio and Leviatan (1997) in which they have studied the coupling of a single j-shell particle (the unnatural parity orbital in fact) to the underlying scissors mode. In the limit of good F spin and large boson number N, the resulting new sum rule, has been compared with a similar

sum rule derived for even-even nuclei (Ginocchio, 1991; von Neumann-Cosel et al., 1995). These results yield upper limits and in the case of the odd-neutron $1i_{13/2}$ particle coupled to the scissors for the nuclei 161 Dy and 167 Er as some extremes (see the data), it is observed that the theoretical sum rule in 167 Er is consistent with the data but for 161 Dy definitely largely overestimates them.

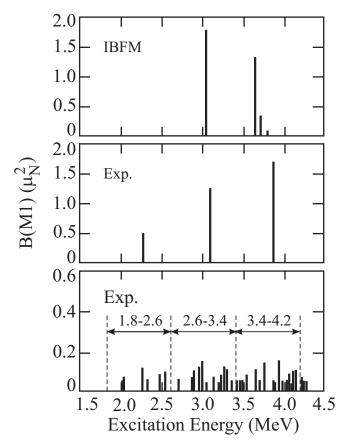


FIG. 41 Bottom to top: experimental M1 strength distribution in 167 Er, strength distribution summed up in the indicated energy bins, comparison with an IBFM calculation, and expected splitting of the energy spectrum through coupling of the unpaired particle (Schlegel *et al.*, 1996).

A number of interesting results have been derived by Van Isacker and Frank (1989) and by Frank et al. (1991) using the more general IBFM. Analytical results could still be derived under the assumption of good F spin and considering a single-j shell. Calculations were performed for ¹⁶⁹Tm and ¹⁶⁵Ho (Huxel et al., 1992). A more detailed study has also been carried out for ¹⁶³Dy and 167 Er. In the case of 167 Er, where the $1i_{13/2}$ oddneutron determines the ground-state structure, this orbital only is considered and again comparisons with summed strengths are possible, as depicted in Fig. 41. The multiplet structure resulting from coupling the oddparticle to the scissors excitation always underestimates fragmentation by far. Within the algebraic formulation, Devi and Kota (1992a,b, 1996) have been studying group-theoretical reductions for odd-A nuclei now including the g boson. Because of the very fact of introducing an extra boson degree of freedom, the effect of fragmentation of course increases.

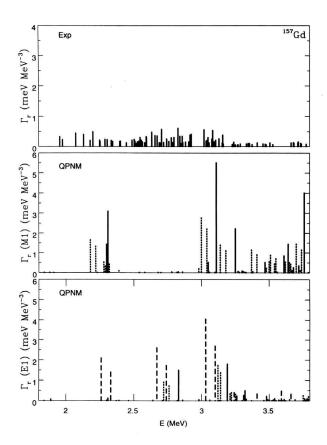


FIG. 42 Experimental ground-state reduced width distribution in 157 Gd (top), together with the QPNM predictions (Soloviev, 1992) for M1 (middle) and E1 (bottom) transitions to K=1/2,3/2,5/2 final states displayed by full, dashed and dotted lines, respectively (Soloviev $et\ al.$, 1997c).

Microscopic studies in the deformed rare-earth region have also been carried out using the coupling of a 1qp excitation with the underlying collective phonon structure. Calculations have been performed by Raduta and Lo Iudice (1989) and Raduta and Delion (1990), using a coherent-state formalism to describe the collective phonons. A more extensive calculation of this kind exists also for a number of odd-mass nuclei (Soloviev et al., 1996, 1997a,c). In the particular case of ¹⁵⁷Gd illustrated in Fig. 42, M1 strength is localized mainly between 2 and 2.5 and between 3 and 3.7 MeV, clearly overestimating the observed M1 strength. The strongest concentration of M1 strength is situated near 3 MeV with a subsequent large fragmentation (i) due to the fact that from the J; K ground state M1 excitations are possible into J+1, J, J-1; K-1 and J+1; K+1 excited states, and (ii) the 1qp-phonon coupling mechanism redistributes strength too. The calculations indeed give support to the presence of scissors particle-core coupled

configurations in the odd-mass strongly-deformed rareearth nuclei, albeit with a varying character of splitting and further fragmentation over a background of complex microscopic configurations.

V. MAGNETIC DIPOLE EXCITATIONS IN LIGHT AND MEDIUM-HEAVY NUCLEI

A. Experimental data

Like in heavier nuclei, most experimental information on ground-state isovector magnetic dipole transitions comes from inelastic photon and electron scattering experiments. A particularly interesting technique for the study of even-mass nuclei is electron scattering at 180° (Fagg, 1975) because of the dramatic suppression of the background due to the radiative tail of elastic scattering. If combined with a large-acceptance spectrometer it represent a powerful tool for the study of the M1 response (Lüttge et al., 1995, 1996b). Data from (γ, γ') in lighter nuclei are limited. Summaries are provided e.g. by Berg and Kneissl (1987) and Raman et al. (1991). The most exhaustive studies of M1 strength distributions are available from (e, e') data covering all stable nuclides in the p-shell, N = Z and Z + 2 nuclides in the sd-shell, the stable Ca isotope chain, the N=28 isotones, the open-shell nuclei ^{46,48}Ti, ⁵⁰Cr, ⁵⁶Fe, and finally ⁵⁸Ni.

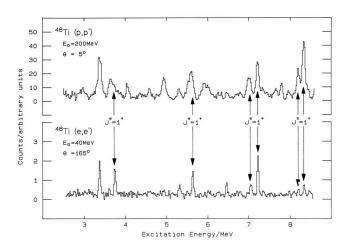


FIG. 43 High-resolution (p, p') and (e, e') spectra in ^{48}Ti . The $J^{\pi} = 1^{+}$ states are marked by arrows (Richter, 1990).

Complementary information on the spin part of the M1 strength stems from inelastic proton scattering experiments. As an example, in Fig. 43 spectra of the (p,p') and (e,e') reactions off 48 Ti under kinematics favoring M1 excitations are compared (Richter, 1990). All transitions identified to have M1 character (marked by arrows) are seen in both spectra although with different relative intensities due to the interference of orbital and spin strength in the latter reaction. Another case, 56 Fe, highlighting the close resemblance of spectra obtained

with both probes (Richter, 1994) is displayed in Fig. 44. Here, all transitions observed above $E_x = 6$ MeV possess M1 character and represent the spin-M1 Gamow-Teller resonance. At lower excitation energies mostly 2^+ states are populated except for the prominent transition seen in the (e, e') spectrum at about 3.5 MeV, which again carries most of the orbital M1 strength.

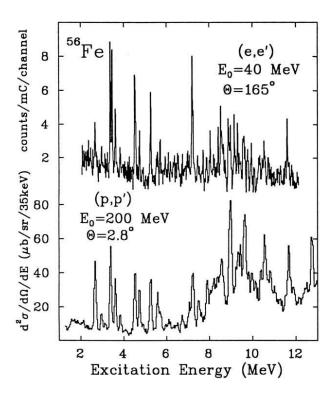


FIG. 44 High-resolution inelastic electron and proton scattering spectra in 56 Fe measured at TRIUMF. Lines above 6 MeV correspond to the excitation of $J^{\pi} = 1^{+}$ states under the kinematic conditions of the two experiments (Richter, 1994).

B. Theoretical description: the shell model and random phase approximation

Contrary to the problems encountered to describe the magnetic dipole excitation modes in strongly deformed heavy nuclei within a shell-model framework, for the light nuclei large-scale shell-model calculations have been performed and used to study M1 excitation properties in detail. Concentrating on mixed-symmetry states, it were in particular Zamick and coworkers who studied such excitations for nuclei in the $1f_{7/2}$ shell-model region (Liu and Zamick, 1987a,b,c; Zamick, 1985, 1986a,b). Let us treat ⁴⁸Ti with 2 proton particles and 2 neutron holes outside of the closed shell nucleus ⁴⁸Ca as an example (analogous arguments hold for other nearby nuclei). The wave functions within the $1f_{7/2}$ model space solely are

expanded as

$$\Psi(J^{\pi}) = \sum_{L_p, L_n} D_J(L_p, L_n) | (1f_{7/2})_{L_p}^2, (1f_{7/2})_{L_n}^{-2}; JM \rangle,$$
(31)

in which the notation is self-explanatory and where the coefficients $D_J(L_p, L_n)$ denote the amplitudes that the two protons couple to L_p and the two neutron holes to L_n . The M1 operator now induces transitions from the 0^+ ground state into the 1^+ states with a corresponding B(M1) value of

$$B(M1) = \frac{3}{4\pi} \left(g_j^{\pi} - g_j^{\nu} \right)^2 \times \left[\sum_{L} D_0(L, L) D_1(L, L) \right]^2 L(L+1) \mu_N^2,$$
(32)

in which g_j^π and g_j^ν are the conventional single-particle gyromagnetic factors $g_j^\rho = [(2j-1)g_l^\rho + g_s^\rho]/2j$, and $g_j^\rho = [(2j+3)g_l^\rho - g_s^\rho]/2(j+1)$, for $j=l+\frac{1}{2}$ and $l-\frac{1}{2}$, respectively, and $\rho=\pi,\nu$ (Brussaard and Glaudemans, 1977). One can even derive a sum rule for the strength into all possible final 1^+ states

$$\sum B(M1) = \frac{3}{4\pi} \left(g_j^{\pi} - g_j^{\nu} \right)^2 \times \left| \sum_{L} D_0(L, L) \right|^2 L(L+1) \ \mu_N^2.$$
 (33)

For the case of ⁴⁸Ti the mixed-symmetric 2⁺ state results from the mixed-symmetric combination of the $L_p = 0, L_n = 2$ and $L_p = 2, L_n = 0$ components whereas the 1^+ scissors counterpart originates from the combinations $L_p = 2$, $L_n = 2$; $L_p = 4$, $L_n = 4$; etc. While the use of a single $1f_{7/2}$ orbital keeps the M1 strength very much concentrated in a single strong excitation, gradually increasing the shell-model space with the inclusion of the $1f_{5/2}, 2p_{3/2}$ and $2p_{1/2}$ orbitals opens the way into many new states and fragmentation starts to set in, as illustrated in Fig. 45 (Liu and Zamick, 1987c). One notices that the low-lying $J^{\pi} = 1^+$ state is always well separated from the other 1⁺ states at higher excitation energy. It is mainly of orbital nature and can be associated with an experimental state at $E_x = 3.74$ MeV which carries a strength of $B(M1)\uparrow=0.52(8)~\mu_N^2$ (Guhr et al., 1990). We also emphasize that within this single $1f_{7/2}$ shell-model space, the orbital M1 strength is a measure of dynamical quadrupole correlations in the ground state since it depends on the $(2_p^+2_n^+)$ configuration admixture in the ground state.

An important issue in the study of scissors mode excitations in the open-shell nuclei situated in the Ti, Cr, Fe region is the orbital-to-spin ratio for the low-lying 1^+ states. In contrast to heavy nuclei, only a few orbitals are determining the structure of the wave functions and therefore, a non-negligible spin contribution will be present, even in the lowest 1^+ state. This can be studied

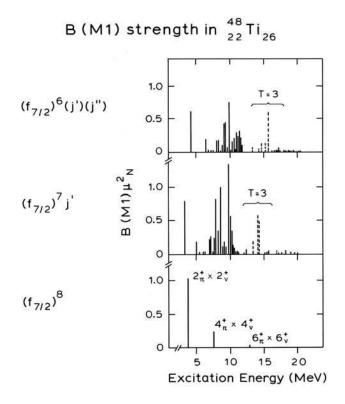


FIG. 45 Magnetic dipole strength distributions for ⁴⁸Ti as calculated for the model spaces given on the l.h.s. of the figure. (Reprinted with permission from Liu and Zamick (1987c). ©(2009) Am. Phys. Soc.)

in a most illustrative way by comparing (e, e') and (p, p') experiments (cf. Figs. 43 and 44). However, a quantitative analysis of the (p, p') data is hampered by the dependence of the extracted spin-M1 strengths on the choice of the effective projectile target interaction, which can lead to variations up to about 40% (Hofmann *et al.*, 2007).

Thus, it is important to find other means to disentangle the spin and orbital parts. It has been discussed e.g. by Abdelaziz and Elliott (1987) that the GT matrix element in β decay might be used to estimate the spin contribution to a collective isovector M1 transition. Such matrix elements can also be measured in charge-exchange reactions populating analog states in the odd-odd neighboring nuclei. A wealth of high-resolution data on the GT strength distributions has recently become available (Frekers, 2006; Fujita et al., 2008) and the dependence on the effective projectile-target interaction in hadronic reactions can partly be circumvented by normalizing to β -decay results. However, isospin selection rules limit the applicability to special cases (some of which are discussed below). Electron scattering form factors present another method to derive bounds on the relative importance of orbital versus spin magnetism in a number of transitions. We illustrate in Fig. 46 two form factors in ⁴⁸Ti for transitions to the 3.74 MeV and 7.22 MeV 1⁺ states, respectively (Guhr et al., 1990). Whereas the first transition

seems to proceed through the $1f_{7/2}$ orbital only (recoupling), the second transition is consistent with a spin-flip transition $1f_{7/2} \rightarrow 1f_{5/2}$.

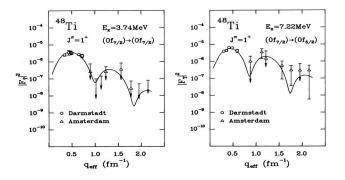


FIG. 46 Inelastic electron scattering form factors for a predominantly orbital (left-hand side) and predominantly spin-flip (right-hand side) M1 transition (Richter, 1990).

A systematic study of the orbital-to-spin ratio was carried out by Oda and Muto (1987) in which a dependence on both the model space and the use of effective g factors was explored. The result is that including excitations from the $1f_{7/2}$ orbital into the higher-lying $1f_{5/2}, 2p_{3/2}, 2p_{1/2}$ orbitals increases the orbital part over the spin part in a systematic way for a number of Ti, Cr, and Fe isotopes. Moreover, quenching the g factors from the free-nucleon values further reduces the spin strength and reinforces the model space extension. The latter result was also derived independently by Heyde and Sau (1984); see also Heyde (1989) for an illustration. Studies of the M1 response for the light Ti nuclei have been carried out using the QRPA approach with very similar results and conclusions concerning the orbital-to-spin ratio (Faessler and Nojarov, 1988; Faessler et al., 1989; Nojarov et al., 1987, 1991, 1992).

On the other hand, truncations of the large-scale shell-model space and choosing a specific proton-neutron force may lead to a symmetry-based approach to study For light nuclei, the SU(3) shell-M1 properties. model has been used and applied to both sd- and fpshell nuclei (Chaves and Poves, 1986; Poves et al., 1989; Retamosa et al., 1990). A comparison of shell-model and IBM calculation was performed for the light Sc, Ti, V fp-shell nuclei (Abdelaziz et al., 1988). A particular symmetry-dictated truncation to realistic shellmodel calculations, emphasizing the importance of S,D, and G pairs, was used for 54,56 Cr and $^{56-60}$ Fe (Halse, 1990, 1991b). Moreover, a pseudo SU(3) model was suggested to describe rotational properties in this mass region (Halse, 1991a).

Recent computational progress allows shell-model studies of the M1 strength in large model spaces to describe details of the fragmentation of the mode. For example, unrestricted calculations in the full fp model space are possible now for 46,48 Ti (Fearick $et\ al.$, 2006). As an example of the state-of-the-art, a study

of the stable N=28 isotones $^{48}\mathrm{Ca},~^{50}\mathrm{Ti},~^{52}\mathrm{Cr},~^{54}\mathrm{Fe}$ (Langanke et al., 2004) is discussed, whose experimental M1 strength distributions have been measured in great detail (Sober et al., 1985; Steffen et al., 1980). In Fig. 47, the results for ⁵²Cr are shown together with calculations based on two widely used shell-model interactions called KB3G (Poves et al., 2001) and GXPF1 (Honma et al., 2004) derived in a G-matrix approach from nucleonnucleon interaction potentials. Because of the N=28shell closure a spherical ground state can be expected. Correspondingly, no low-lying orbital transitions are observed. In the energy region above 6 MeV, presented in Fig. 47, a resonance structure arising from spin-flip transitions is visible. The shell-model results are quite successful in reproducing the features of the strength distribution qualitatively and also quantitatively when analyzing the resonance centroid and total strength, but in detail differences remain.

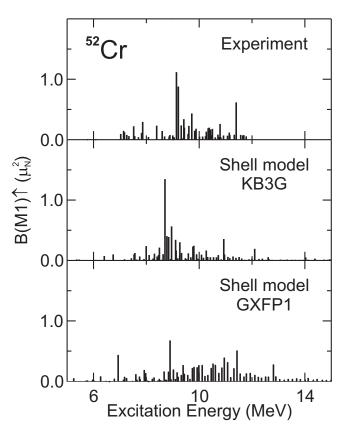


FIG. 47 M1 strength distribution in 52 Cr from (top to bottom) high-resolution (e,e') experiments (Sober et~al.,~1985) and large-scale shell-model calculations using the KB3G (Poves et~al.,~2001) and GXPF1 (Honma et~al.,~2004) interactions, respectively.

The comparison made in Fig. 47 raises an important and nontrivial question: how can one quantify the degree of correspondence between data and calculation? One possible way may be the extraction of scales characterizing the fine structure of the mode by means of a wavelet analysis (Shevchenko et al., 2008).

Fine structure has been shown to be a global phenomenon of giant resonances (Kalmykov et al., 2006; Shevchenko et al., 2009, 2004). A recent application of this method to M1 strength distributions in fp-shell nuclei (Petermann et al., 2010) indeed reveals considerable differences in the characteristic scales extracted from the M1 strength functions obtained with different effective interactions including those shown in Fig. 47, and in the case of 52 Cr the results from the KB3G calculation are found to be closer to the data.

The capability of large-scale shell model calculations to describe the interference of spin and orbital parts has been investigated in a detailed study of the electron scattering form factor of the prominent M1 transition at $E_x=3.449$ MeV in $^{56}{\rm Fe}$, which contains spin and orbital matrix elements of comparable size (Fearick et~al., 2003). While different effective shell-model interactions describe the strength reasonably well, the predicted q dependence differs considerably. Clearly, the spin-orbit interplay remains a challenge to shell-model studies even in very large model spaces.

It also became clear in the above studies that a reduction of the g factors from the free-nucleon values is generally increasing the orbital-to-spin matrix element ratio. An independent approach to understand the qfactor quenching has been carried out in this region of medium-heavy and light nuclei, concentrating on the comprehensively studied N=28 nuclei. Shell-model calculations require in all cases a reduction of the spin part of the magnetic dipole operator. A consistent description for the stable N=28 isotones can be reached using a value $g_s^{eff} = 0.75(2)g_s^{free}$ (von Neumann-Cosel et al., 1998). The required reduction is remarkably close to the quenching factor 0.744(15) obtained from a recent shellmodel analysis of GT β -decay transitions in the lower fpshell region (Martínez-Pinedo et al., 1996). Indeed, the most important mechanism responsible for the quenching is, viz. the mixing with two-particle two-hole configurations at high excitation energies, is expected to be the same as in the GT case (Bertsch and Hamamoto, 1982; Ichimura et al., 2006).

C. Some astrophysical implications

The knowledge of the magnetic dipole strength in fp-shell nuclei is also crucial in supernova modeling. It permits to determine cross sections of inelastic neutrino-nucleus scattering, a process whose importance for supernova dynamics was recognized only recently (Hix et~al., 2003). Under the conditions of a supernova type II in massive stars, neutrino-nucleus reactions are dominated by GT transitions (Langanke and Martínez-Pinedo, 2003). The description of inelastic scattering processes requires knowledge of the $T_0 \rightarrow T_0$ isospin component of the GT strengths, where T_0 denotes the ground-state isospin. Except for an overall factor relating the weak and electromagnetic interaction

this is nothing but the spin part of the M1 strength. Orbital strength is negligible in the stable N=28 isotones due to the shell closure; thus, the measured M1 distribution represents to a good approximation the needed GT_0 strength. Figure 48 displays differential cross sections of inelastic neutrino scattering on 52 Cr calculated under this assumption for two typical neutrino energies $E_{\nu}=15$ and 25 MeV (solid line).

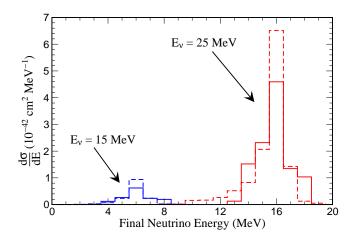


FIG. 48 (Color online) Differential inelastic neutrino cross sections for $^{52}\mathrm{Cr}$ and initial neutrino energies $E_{\nu}=15~\mathrm{MeV}$ and 25 MeV. The solid histograms are obtained from the M1 data, the dashed ones from shell-model calculations. The bumps represent the GT_0 strength shifted by the centroid energy of the resonance. The final neutrino energies are given by $E_f=E_{\nu}-E_x$ (Langanke et al., 2004).

Since the reaction network calculations in a supernova require information on the GT₀ strength in many nuclei, one has to rely on model calculations. The set of highly precise data in the N=28 isotones (Sober et al., 1985) was used to demonstrate the capability of large-scale shell-model calculations to describe M1 strength distributions (Langanke et al., 2004). If applied to the present problem, differential neutrino scattering cross sections shown by the dashed line in Fig. 48 result. These are indeed in good agreement with the data, and it was concluded that present-day shell-model calculations can provide the necessary GT₀ strengths needed as input to the supernova simulations. Inclusion of inelastic neutrinonucleus scattering increases the neutrino opacities noticeably and strongly reduces the high-energy tail of the neutrino spectrum emitted in the neutrino burst at shock breakout. Relatedly, the expected event rates for the observation of such neutrinos by earthbound detectors are reduced by up to about 60% (Langanke et al., 2008).

D. Selected problems in light and medium-heavy nuclei

In light nuclei, magnetic dipole transitions have been measured in almost all stable nuclei and detailed theoretical studies (mostly shell-model) have been performed. We do not attempt an exhaustive discussion of these results. Rather, we highlight a few topics accessible only in light nuclei or extending beyond the range of physics questions discussed so far.

1. Quenching of the spin M1 strength: the case of ^{48}Ca

A classical example of a clear-cut isovector M1transition with a remarkably large transition strength $B(M1)\uparrow=3.9(3)~\mu_N^2$ was found in ⁴⁸Ca in electron scattering experiments (Steffen et al., 1980). It was interpreted as a rather pure neutron $1f_{7/2} \rightarrow 1f_{5/2}$ excitation (Richter, 1985). The (e, e') form factor of this transition serves as a another prime example to study the problem of quenching of the spin part of the isovector M1response. Since GT and M1 strengths are quenched by comparable amounts one can expect that the responsible processes are the same. Two major mechanisms are expected to contribute (Ericson and Weise, 1988). On the one hand, a polarization of the nuclear core by the nucleon undergoing a spin-isospin transition occurs, which leads to a virtual excitation of high-lying states by the tensor force in second order; hence it is named secondorder core polarization. Alternatively, the GT and M1operators are modified by virtual Δ -hole excitations, analogous to the Lorenz-Lorentz correction in dielectrics (Delorme et al., 1976). Both mechanisms lead to different redistributions of the strength. Thus, by measuring the response up to high excitation energies, the former mechanism was shown to dominate for GT transitions (Ichimura et al., 2006). The study of the (e, e') form factor permits to delineate the contributions in the case of the M1 response (Richter, 1991). The most refined calculations (Takayanagi et al., 1988), although still not able to fully explain the amount of quenching, confirm the dominant role of core polarization leading to mixing with 2p2h states over the Δ -hole part, which contributes about 10% to the strength reduction. This value is in agreement with an analysis of the Δ -hole contributions to the quenching of GT strength (Ichimura et al., 2006).

2. Cross-shell transitions in $^{36,38,40}\mathrm{Ar}$

Magnetic dipole strength distributions in 36 Ar and 38 Ar deduced from electron scattering experiments (Foltz et al., 1994) reveal marked differences. Calculations in a $0\hbar\omega$ model space (sd shell) with a phenomenological interaction, generally successful in the description of M1 and GT strengths (Brown and Wildenthal, 1988), work well for 36 Ar (with a closed neutron sd shell) but fail completely for 38 Ar (with two neutrons in the fp shell), indicating the importance of $sd\rightarrow fp$ cross-shell contributions. This problem has recently been studied by Lisetskiy et al. (2007) based on an effective interaction including the coupling of sd- to the $1f_{7/2}2p_{3/2}$ - orbitals (Caurier et al., 2001). Such cross-shell calculations present a limit of present-day computational capabilities

and still require a significant truncation of the model space. The first results support the importance of an inclusion of cross-shell transitions and make specific predictions for the even more neutron-rich 40 Ar. Some information on M1 strength in 40 Ar has recently been reported (Li et al., 2006), but a measurement of the full M1 strength distribution would be important.

3. l-forbidden transitions

In sd-shell nuclei, an effective shell-model M1 operator has been determined by an empirical fit to the large body of data on magnetic and M1 transitions (Brown and Wildenthal, 1987). The deviations from the bare operator are incorporated in correction factors for the spin and orbital parts and an induced-tensor term. Microscopic calculations (Arima et al., 1987; Towner, 1987) are in good agreement except for an isovector tensor correction. Tensor corrections are generally weak and therefore buried in the dominant spin strength for most M1 transitions. However, experimental information on the tensor correction terms can be obtained from l-forbidden transitions $(1d_{3/2} \leftrightarrow 2s_{1/2})$ in the sd shell). The term 'l-forbidden' refers to a selection rule for the one-body operator of M1 or GT transitions which does not allow a change of the orbital quantum number. The higher-order corrections to the l-forbidden transitions are theoretically expected to be dominated by Δ resonance admixtures into the nuclear wave functions (Arima et al., 1987; Towner, 1987) and they are a unique observable in this respect. The problem has been studied extensively in $1d_{3/2} \rightarrow 2s_{1/2}$ single-hole transitions in A = 39 nuclei. One finds an order of magnitude larger M1 strength (Grundey et al., 1981) relative to the GT strength (Hagberg et al., 1994), while the microscopic results predict the tensor correction governing the strength to be the same. However, the interpretation could be blurred by weak cross-shell admixtures of the type discussed in Sec. V.D.2. Therefore, data away from the end of the sd shell are important. One such example is an (e,e') study of the *l*-forbidden transition to the 1^+ state in 32 S at $E_x = 7.003$ MeV (Reitz et al., 1999). The form factor exhibits an anomalous momentum-transfer dependence compared to allowed M1 transitions because its finite strength results from higher-order terms only. The shell-model analysis reconfirms the discrepancy between empirical and microscopic approaches to determine the tensor correction, and the problem remains unresolved so

Of course, l-forbidden transitions are not restricted to the case $1d_{3/2} \leftrightarrow 2s_{1/2}$ but can appear between all pairs of shell-model orbitals with quantum numbers (n,l,j=l+1/2) and (n-1,l+2,j'=(l+2)-1/2), where n is the radial quantum number, and l and j are the orbital and total angular momenta, respectively. Experimentally, the single-particle energies of corresponding pairs of states show a near degeneracy in many nu-

clei. This led to the concept of pseudospin symmetry, where the doublet structure is expressed in terms of a 'pseudo'-orbital angular momentum $\tilde{l} = l + 1$, in which the two levels represent spin-orbit partners with a 'pseudo'spin $\tilde{s} = 1/2$. While pseudospin symmetry was empirically established 40 years ago (Arima et al., 1969; Hecht and Adler, 1969), a deeper understanding has been lacking. Relativistic corrections (Bohr and Mottelson, 1969) have been suggested as a possible source and applied with some success to describe magnetic moments of pseudospin partners near the N=82 shell closure (Heyde et al., 1977). Recently, pseudospin symmetry has been interpreted as a relativistic SU(2) symmetry of the Dirac-Hamiltonian which occurs when the attractive scalar and repulsive vector nuclear mean fields cancel (Ginocchio, 2005). Evaluating this concept, Ginocchio (1999) derived a relation between the magnetic moments of the pseudospin partners and the strength of the lforbidden M1 transition between them. Application to data near a variety of magic numbers reveals overall good correspondence with a few marked deviations (von Neumann-Cosel and Ginocchio, 2000).

4. Enhancement of magnetic dipole strength by meson exchange currents

Direct signatures of mesonic exchange currents (MEC) in experimental observables are usually restricted to fewbody nuclear systems (Ericson and Weise, 1988) with the exception of magnetic dipole properties (moments and transition strengths) as discussed in the previous subsection. One such example is discussed here. It is well established in sd-shell nuclei that full $0\hbar\omega$ shell model calculations with an effective operator are able to describe the M1 and GT matrix elements (Arima et al., 1987; Brown and Wildenthal, 1987; Towner, 1987). In selfconjugate even-even nuclei with ground-state spin and isospin $J; T = 0^+; 0$ the set of final states populated by isovector M1, GT_- and GT_+ transitions forms a triplet of isobaric analogue states. Their transition strengths are directly related, if spin-orbital interference effects are negligible. This is certainly not the case for individual transitions but holds on the level of 10% when studying full strength distributions of sd-shell nuclei (Hino et al., 1988) because of the sign variations of the mixing term.

When comparing M1 and GT strength distributions in ²⁴Mg, excellent agreement of the GT strengths amongst each other and with the shell model result. However, the same calculations significantly underpredict the M1strength. Such an enhancement of the experimental M1 strength can be traced back (Richter et al., 1990) to MEC contributions. To make this clear it is convenient to describe the M1 and GT strengths schematically in the following form

$$B(M1) = C[M_{\sigma} + M_{l} + M_{\Delta} + M_{V}^{MEC}]^{2} , \quad (34)$$

$$B(GT) = [M_{\sigma} + M_{\Delta} + M_{A}^{MEC}]^{2} . \quad (35)$$

$$B(GT) = [M_{\sigma} + M_{\Lambda} + M_{\Lambda}^{MEC}]^2 . \tag{35}$$

Here, M_l and M_{σ} are the orbital and spin matrix elements, and M_{Δ} stands for the contribution of Δ -isobar admixtures to the strength. The numerical factor C before the square brackets in Eq. (34) equals to 2.643 μ_N^2 using free nucleon g factors. Neglecting the orbital part, the main difference between M1 and GT excitations lies in the MEC contributions, which are of vector type for the former and of axial-vector type for the latter. Since axial vector currents are strongly suppressed because of the conservation of G parity (Towner, 1987), deviations of the ratio $R(M1/GT) = \sum B(M1)/2.643 \sum B(GT)$ from unity point towards an enhancement of the M1 strength by vector-type MEC contributions. Besides ²⁴Mg, a clear enhancement was also observed in ²⁸Si (Lüttge *et al.*, 1996a; von Neumann-Cosel et al., 1997). In 32 S the situation is less clear (Hofmann et al., 2002) because (e, e')form factors indicate significant orbital admixtures in some of the strongest transitions, and the experimental information on the M1 strength distribution is limited to an excitation energy of 12 MeV and therefore incomplete. Another problem noted in Richter et al. (1990) was that spin-M1 strengths in selfconjugate sd-shell nuclei deduced from forward-angle (p, p') data (Crawley et al., 1982) are systematically about 20% larger than the corresponding GT± strengths. However, this discrepancy can probably be resolved utilizing the latest experimental developments allowing true 0° measurements combined with high energy resolution (Tamii et al., 2009).

5. Isoscalar and isovector M1 transitions in $^{12}\mathrm{C}$ and isospin mixing

Isospin is an approximate symmetry in nuclei broken by the long-range Coulomb force but also by small charge-dependent components of the nuclear interaction. In light nuclei, Coulomb effects are weak and excited states possess a well defined isospin quantum number T experimentally known in many cases. This allows to study isospin mixing between states of the same J^{π} but different T. Evidence for isospin mixing beyond the Coulomb force has been claimed from the observation of very large isospin mixing matrix elements but later it was realized that the predictions exhibit a strong dependence on the poorly known radial wave functions of the involved single-particle states (Auerbach, 1983).

A unique testing ground are the M1 transitions to the pair of J^{π} ; $T = 1^{+}$; 0 and J^{π} ; $T = 1^{+}$; 1 states in ¹²C at 12.71 and 15.11 MeV, respectively. These are of $1p_{3/2} \rightarrow$ $1p_{1/2}$ spin-flip character. Form factors of both transitions (albeit weak for the isoscalar case) at low momentum transfer have been measured with high precision in inelastic electron scattering (von Neumann-Cosel et al., 2000). Analysis in a two-state model determines not only the mixing amplitudes but also the relative sign through the q dependence of the form factors. The resulting Coulomb matrix element $\langle H_c \rangle = 118(8)$ keV, determined with unequaled precision, is large but can be fully explained by

Coulomb mixing (Harney et al., 1986).

VI. ISOVECTOR MAGNETIC DIPOLE TRANSITIONS IN VIBRATIONAL NUCLEI

A. Introduction

In most of this review we have concentrated on the magnetic dipole orbital and spin response in stable, deformed nuclei, which has been studied using both electromagnetic and hadronic scattering off the nuclear ground state (Sec. I.B). In particular, the scissors 1⁺ mode was shown to be excited with a summed strength that scales with the square of the nuclear deformation. Therefore, with decreasing deformation entering the region near closed shells, the scissors mode as well as a stable intrinsic quadrupole deformation will cease to be formed. This mass region is characterized by small amplitude quadrupole vibrational oscillations as the major degree of freedom, with typical $B(E2; 2_1^+ \to 0_1^+)$ strength of the order a few tens of Weisskopf units (W.u.). The low-energy nuclear structure properties result in a first excited 2^{+}_{1} phonon excitation which is an in phase motion (or symmetric mode) in the proton and neutron collective motion, also called isoscalar (IS) mode. Multiphonon states can then be constructed, as shown in Fig. 3, in which the proton and neutron motion combines into symmetric (or IS) excitations. However, multiphonon states can also arise rise from proton and neutron motion combining into non-symmetric (or IV) excitations. Besides the 1⁺ states, which is the counterpart of the scissors mode in the vibrational nuclei, also $0^+, 2^+, 3^+, 4^+$, states result.

The appearance of IV proton-neutron excitations has been proposed in the context of the IBM-2 (Arima et al., 1977; Iachello, 1984; Iachello and Arima, 1987; Otsuka and Ginocchio, 1985). This approach points out that the isovector excitations appear in a natural way by combining the lowest-lying proton and neutron 2_1^+ d-boson configurations (Iachello, 1984; Van Isacker et al., 1986) into states of mixed-symmetry (MS) character (Iachello, 1984). Besides, the shell model constitutes a microscopic framework in order to describe excitations that are non-symmetric in its proton and neutron coordinates (Boelaert et al., 2007b; Heyde and Sau, 1986; Holt et al., 2007; Lisetskiy et al., 2000). Moreover, it is possible to describe isoscalar and isovector excitations within the framework of a quasiparticle-phonon model which defines RPA phonons, and then to construct states in a basis of one-, two- and three-phonon components. Since this approach has a microscopic (QRPA) underpinning, it allows to bridge the gap between a fully microscopic shell-model approach and the algebraic IBM-2 (Lo Iudice and Stoyanov, 2000, 2002, 2004, 2006; Lo Iudice et al., 2009, 2008).

In order to locate these mixed-symmetry (MS) states, one can use the particular structure of the M1 and E2 operators. We have already shown that the M1 magnetic

dipole operator can be separated into its isoscalar and isovector parts (see Sec. III.B.1). In view of the structure of the isovector part, one expects strong magnetic dipole transitions in the decay of the mixed-symmetry states into the low-lying symmetric states. Likewise, one can separate the electric quadrupole operator T(E2)

$$T(E2) = e_{\pi} \sum_{i=1}^{Z} r_{i,\pi}^2 Y_2(\hat{r}_{i,\pi}) + e_{\nu} \sum_{i=Z+1}^{A} r_{i,\nu}^2 Y_2(\hat{r}_{i,\nu}), \quad (36)$$

with e_{π} and e_{ν} the proton and neutron effective charges into an isoscalar and isovector part

$$T(E2) = \frac{e_{\pi} + e_{\nu}}{2} T(E2, IS) + \frac{e_{\pi} - e_{\nu}}{2} T(E2, IV). \quad (37)$$

Here, T(E2, IS) and T(E2, IV) are the symmetric and antisymmetric combinations of the proton and neutron parts of the E2 operator. Because of the specific symmetry character of the IS and IV excitations, strong E2 transitions are expected between S and MS states, separately, but rather weak E2 transitions from MS to S states. These characteristics are highlighted in Fig. 3.

In view of the above discussion, the key signature, in order to assign mixed-symmetry character to a state, derives from the E2 and M1 decay properties: (i) strong M1 transitions (B(M1)) of the order of $\simeq 1\mu_N^2$) to low-lying symmetric states restricting to transitions between states with equal number of phonons, mainly, (ii) weak collective E2 transitions (with transition probabilities about 10% of the strong E2 transitions such as $2_1^+ \to 0_1^+$) to low-lying symmetric states, and (iii) strong collective E2 transitions amongst the MS states themselves.

B. Experimental results and theoretical description

¿From an experimental point of view, the study of MS states in nuclei of vibrational and transitional structure is rather different from the mapping of scissors mode 1⁺ excitations in deformed nuclei. In the latter case (see Sec. I.B), electron, photon and hadron scattering starting from the 0⁺ ground state in deformed nuclei allowed to determine orbital and spin-flip M1 strength. In the present situation, the identifying elements are both a strong M1 transition into the isoscalar (mostly the 2^+_1 state) state accompanied by a weak E2 transition into the 0^+ groundstate (with a magnitude of the order of a few % of the $B(E2; 2_1^+ \to 0_1^+)$ reduced transition probability). Thus, in order to obtain a unique characterization of the mixed-symmetry states, different types of experiments have to be carried out and combined. Typical experiments will need to probe the lifetime of a given level, the determination of J^{π} values, γ -decay branching ratios and the $\delta(E2/M1)$ mixing ratios and this for as complete a set of states with given J^{π} value. Classical γ -ray spectroscopic methods have to be used extensively as well photon-scattering experiments as e.g.

 γ decay following β decay that populates levels in nuclei under study. Determining the nuclear lifetimes needs typically the study of Doppler-shifted attenuation techniques (DSAM) in $(n,n'\gamma)$, light-ion induced reactions and Coulomb excitation in inverse kinematics. Direct excitation as used in electron scattering, photon scattering and Coulomb excitation on stable nuclei gives rise to lifetimes in a rather straightforward way. A detailed review of these techniques is presented by Kneissl $et\ al.$ (2006) and Pietralla $et\ al.$ (2008).

1. The Z \sim 40, N \sim 50 mass region

The nucleus ⁹⁴Mo forms a particularly suitable test case of the above schematic picture because it has 4 protons outside of the Z = 38 subshell closure and 2 neutrons outside of the N = 50 closed shell. With these building blocks, ideal conditions show up to form both symmetric and antisymmetric couplings of these pairs. Detailed experimental studies have identified proton-neutron mixed symmetry 2^+ and 1^+ states (Pietralla et al., 1999). In Fig. 49, the specific signatures of a MS 2^+ state (strong M1 transition into the 2^+_1 state, weakly collective B(E2) transition from the ground state) are shown. The data point towards the 2^+_3 state as the ideal MS candidate. Experiments in ⁹⁴Mo have furthermore shown (Pietralla et al., 2000) evidence for twophonon MS states built from combining a symmetric (2_1^+) and anti-symmetric (2_3^+) state, thus forming states with spins in the range 0⁺- 4⁺. Clear-cut identification of the 2^+ (Fransen et al., 2001) and 3^+ (Pietralla et al., 2000) members could be achieved and candidates for the other spins were identified (Fransen et al., 2003) based on the above discussed signatures of MS states. Independent evidence for one-phonon symmetric and mixed-symmetric states has been recently demonstrated in a combined study of high-resolution inelastic electron and proton scattering off MS 2⁺ states in ⁹⁴Mo (Burda *et al.*, 2007) when comparing with theoretical results derived from quasiparticle-phonon, shell-model and IBM-2 as shown in Fig. 50. Multi-phonon MS states have also been observed in the N=54 ⁹⁶Mo nucleus (Lesher *et al.*, 2007).

The study of the variation of the MS states, keeping the neutron number fixed at N=52 but changing the proton number, is quite interesting. In this spirit, experiments have been carried out in 96 Ru with Z=44 (Klein et al., 2002; Pietralla et al., 2001a) pointing out a strong similarity with 94 Mo. Moreover, evidence for a MS 1^+ state has been shown by Linnemann et al. (2005). A study of the 92 Zr nucleus, with the same number of neutrons i.e., N=52, is quite different because in the core nucleus, 90 Zr, two 0^+ states appear resulting from the presence of both $(1g_{9/2})_{0^+}^2$ and $(2p_{1/2})_{0^+}^2$ configurations. The 0^+ ground state therefore acquires extra binding energy which distorts the vibrational spectra as compared with the N=52 Mo and Ru nuclei. Still, photon scattering

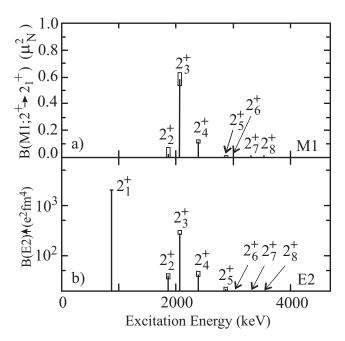


FIG. 49 Measured E2 and M1 strengths in order to identify the $2^+_{1,ms}$ state in ⁹⁴Mo. (a) $B(M1;2^+ \rightarrow 2^+_1)$ values for the seven lowest-lying identified nonyrast 2^+ states. (b) Corresponding $B(E2;0^+_1 \rightarrow 2^+_i)$ values. The error bars are displayed as boxes. (Reprinted with permission from Fransen *et al.* (2003). ©(2009) Am. Phys. Soc.)

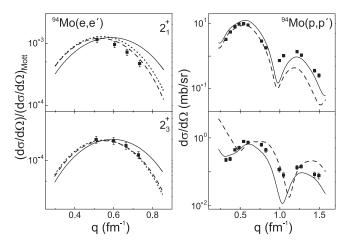


FIG. 50 Momentum-transfer dependence of the symmetric $(2_1^+, \text{ upper part})$ and mixed-symmetric $(2_3^+, \text{ lower part})$ one-phonon excitation cross-sections in ^{94}Mo in inelastic electron (left side) and proton (right side) scattering. The data (full squares) are compared to QPM (solid lines), shell-model(dashed lines) and IBM-2 (dotted lines) predictions (Burda $et\ al., 2007$).

(Werner et~al.,~2002) and $(n,n'\gamma)$ inelastic neutron scattering have enabled to observe 2^+ and 1^+ states with a MS character. Recent experiments (Elhami et~al.,~2007,~2008; Werner et~al.,~2008) have concentrated on the special situation in the Zr nuclei.

The nuclei with proton number 40 < Z < 50 and neu-

tron number close to N = 50 (here, N = 52), with in particular the nucleus ⁹⁴Mo, form an ideal testing ground for both the IBM-2, the QPM as well as the nuclear shell model. It was shown (Iachello, 1984) that the IBM-2 framework naturally predicts a class of states with MS character in the proton and neutron contributions. Detailed selection and intensity rules have been derived for E2 and M1 transitions in the various limits of the IBM-2 (Scholten et al., 1985b; Van Isacker et al., 1986). A comparison of the IBM-2 results has been carried out in ⁹⁴Mo as well as the adjacent N = 52 isotones Ru and Zr with a consistent description of MS states with both a one (2^+) and two-phonon $(1^+, 2^+, 3^+, 4^+)$ character as well as of their decay properties to lower-lying symmetric zero and one-phonon states (see e.g. Pietralla et al. (2008) for more details). Keeping within the context of the IBM-2, a particular scheme (called Q-phonon scheme) has been set up by Otsuka and Kim (1994), which allows for the description of these symmetric and MS excitations as phonons, applicable in the U(5), O(6) symmetry limits and in the transitional nuclei between these two limits.

The IBM-2 has a drawback because the operator only addresses the orbital part and specific spin contributions are only considered in an average way. Therefore, microscopic techniques are needed such as the standard shell model and quasi-particle phonon (QPM) approaches. The QPM approach has been applied with considerable success in the region of vibrational nuclei by Lo Iudice and Stovanov (2000, 2002, 2004, 2006) and Lo Iudice et al. (2008). In the Z = 40, N = 50 region, it turns out that the 2_1^+ RPA phonon has mainly a symmetric structure in the interchange of proton and neutron labels (or is F-spin symmetric in the IBM-2 language) whereas the 2^+_2 RPA phonon is antisymmetric (or of Fspin MS nature) to a good approximation. The QPM eigenvalue problem is then solved in a basis including up to three-phonon states and considering many phonons of different J^{π} nature. These results give support to the IBM-2 calculations carried out in which these s and dbosons are the only building blocks. Detailed results are given for ⁹⁴Mo (Lo Iudice and Stoyanov, 2000, 2002) and for ⁹²Zr (Lo Iudice and Stoyanov, 2004, 2006).

Large-scale shell-model calculations have been carried out for ⁹⁴Mo (Lisetskiy et al., 2000), ⁹⁶Ru (Werner et al., 2002) and 92 Zr (Klein *et al.*, 2002) in the Z = 40, N = 50mass region using a surface delta interaction and treating all valence protons and neutrons outside of the ⁸⁸Sr core with Z=38 and N=50. Here, both the orbital and spin matrix elements contribute, the latter part being non-negligible for M1 transitions and moments. Like in the IBM-2 and QPM, very much the same structure shows up. The specific characteristics of M1 and E2decay characterizing excited states give a microscopic underpinning to the concepts of symmetric and mixed symmetric excitations as used in the Q-phonon classification. The fingerprints that characterize the decay of MS states (see Section VI.A) are clearly observed in the results (Lisetskiy et al., 2000). Shell-model calculations for the N=52 nuclei have been performed within the same model space but now using matrix elements derived from the low-momentum $V_{low\ k}$ nucleon-nucleon interaction (Holt $et\ al., 2007$) giving a rather good reproduction of the experimentally observed results. In Fig. 51, besides the calculated total M1 strength the orbital and spin contributions, which interfere constructively, are given separately.

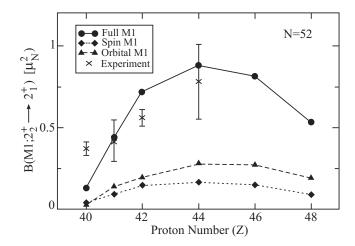


FIG. 51 Evolution of the total, orbital, and spin $B(M1; 2_2^+ \to 2_1^+)$ values for the N=52 isotones. The results of recent shell-model calculations are compared with the data. (Reprinted with permission from Holt *et al.* (2007). ©(2009) Am. Phys. Soc.)

2. Nuclei near other doubly-closed shell regions

a. The region near $Z \sim 50$ The Cd and Te nuclei with 2 proton holes (2 proton particles) away from the Z=50Sn closed core, combined with the neutron filling of the 50 < N < 82 neutron shell form an interesting region to expect mixed-symmetric states. In this mass region, a variety of experiments have been carried out, including the $(n, n'\gamma)$ reaction (Bandyopadhyay et al., 2003; Garrett et al., 1996) for 112,114Cd, photon scattering, eventually combined with Compton polarimetry to deduce the parity unambiguously (Gade et al., 2003; Kohstall et al., 2005; Lehmann et al., 1999) for $^{108-116}$ Cd, β -decay (Linnemann et al., 2007) for 106 Cd, and recoil-distance Doppler-shift measurement (RDDS) after fusion-evaporation reactions (Boelaert et al., 2007a) for ^{102,104}Cd. Evidence for the presence of MS 2⁺ states, slightly above 2 MeV excitation energy, as well as for MS 1⁺ states, at the higher excitation energy near 3 MeV, have been obtained in almost all of these nuclei.

Shell-model calculations have also been reported for the light Cd nuclei with A = 98-106, using the same core as before (⁸⁸Sr), treating all the available valence protons (10 in the case of Cd) and neutrons (Boelaert *et al.*, 2007b). A detailed mapping of shell-model states onto

MS states for both 2^+ and 1^+ states was performed using as criterium, strong M1 transitions to the 0_1^+ and 2_1^+ symmetric states combined with weak E2 transitions to these same states, and strong E2 transitions in between the MS $2^+, 1^+$ states. Likewise, in the even-even $^{122-130}{\rm Te}$ nuclei, an experimental search for MS 1^+ and 2^+ states was performed by Schwengner *et al.* (1997a,b) using photon inelastic scattering, and in ¹²⁴Te by Georgii et al. (1995) using a variety of reactions. This resulted in the detection of candidates for mixedsymmetric 2⁺ states, slightly above 2 MeV excitation energy. More recently, Hicks et al. (2008) have provided detailed results on fragmentation of MS 2⁺ states in the $^{122-130}\mathrm{Te}$ nuclei. The deformation dependence of the g.s. scissors mode strength in these isotopes could by successfully reproduced by QRPA calculations (Guliyev et al., 2002).

b. The rare-earth region: $54 < Z \le 60$ and $72 < N \le 82$ Early evidence for the presence of MS 2⁺ states in rareearth nuclei resulted from an analysis by Hamilton et al. (1984). He showed that in nuclei with two neutrons outside of the N=82 closed neutron shell, with an even number of protons filling the 50 < Z < 82 proton shell, 2⁺ states near 2 MeV would show up. Experiments at the ILL Grenoble using $\gamma - \gamma$ directional correlation experiments (from $\delta\text{-mixing ratios}$ and branching ratios) allowed to find in $^{140}\text{Ba},~^{142}\text{Ce}$ and ^{144}Nd a 2^+_3 state with the typical M1 and E2 branching into the lowerlying symmetric excitations. More recently, these nuclei have been studied using NRF and $(n, n'\gamma)$ scattering, even extending up to ¹⁴⁸Sm (Gade et al., 2004, 2000; Hicks et al., 1998; Li et al., 2005; Mukhopadhyay et al., 2008; Vanhoy et al., 1995) showing evidence for MS 2⁺ states, in most cases exhibiting fragmentation over a number of 2⁺ states near or just above to 2 MeV excitation energy. Likewise, experiments have been carried out in the N = 80 nuclei 134 Xe, 136 Ba and ¹³⁸Ce using the same techniques and Coulomb excitation (Pietralla et al., 1998a, 2008; Rainovski et al., 2006; Scheck et al., 2004; Williams et al., 2009). They again show the presence of MS 2⁺ states and fragmentation, the latter changing quite drastically with the proton number increasing from Z = 56 to Z = 58. As demonstrated in Fig. 52 it is even possible to extract the strength of the residual proton-neutron interaction from the energy splitting between lowest symmetric and mixed-symmetric 2^+ state in these N = 80 nuclei (Ahn et al., 2009). The N = 80 nuclei have been extensively studied by Lo Iudice et al. (2008) within the QPM approach and in a large-scale shell-model study (Sieja et al., 2009). These calculations describe rather well both the variation in excitation energy as well as the changing fragmentation pattern, moving from ¹³²Xe to ¹⁴⁰Nd.

A number of nuclei in which both the proton number and neutron number is steadily increasing moving away from the Z=50 and N=82 shell clo-

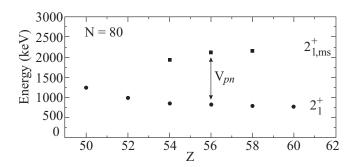


FIG. 52 Excitation energies of the 2_1^+ and 2_{ms}^+ states in N=80 isotones. The splitting is a measure of the residual protonneutron interaction (Ahn *et al.*, 2009).

sure, such as 126,128,130 Xe, 134 Ba, 136 Ce, have been studied using techniques as discussed before (Ahn *et al.*, 2007; Bettermann *et al.*, 2009; Fazekas *et al.*, 1992; Gade *et al.*, 2000; Wiedenhöver *et al.*, 1997). In these nuclei, the signature of a MS 2^+ state shows up consistently slightly above 2 MeV excitation energy.

c. The $A \approx 60$ region The region near the doubly-magic nucleus ^{56}Ni with proton and neutron hole or particle pairs outside of the Z=N=28 core may well give rise to MS couplings of proton and neutron building blocks. The nucleus 56 Fe forms an ideal testing case, and has been studied by Eid et al. (1986),using γ -decay studies, and by Hartung et al. (1989) in electron scattering. Clear-cut evidence for fragmentation of MS 2^+ strength has been observed around 2.6-2.9 MeV. In nearby nuclei such as 54 Cr and 66 Zn, candidates for MS 2^+ strength have been detected near 3 MeV excitation energy (Gade et al., 2002; Lieb et al., 1988). It is interesting to note the increase in energy from the heavier nuclei, where the typical energies are closer to 2 MeV.

d. Heavy nuclei in the vicinity of ^{208}Pb The idea of lowlying 2_{ms}^+ excitations appearing in regions where the number of protons and neutrons forms a stable, closed shell, has been shown to be a general property all through the nuclear mass region. Therefore, the region around Z=82 and N=126 should be a most interesting region one in order to explore the appearance of states which exhibit a MS character in the protons and neutrons. Early evidence was shown by Ahmad et al. (1989) for 2^+ states near 1.5 MeV in 200 Hg. Likewise a 2^+ state near 2.2 MeV was observed in 196 Pt (von Brentano et al., 1996; Jewell et al., 1997). This region has by now not been studied in a systematic way but the Hg, Pt and also the Po, Rn nuclei with neutron numbers close to N=126 should form an ideal testing ground but require radioactive beams.

C. Summary

To sum up this chapter, we have presented evidence for isovector proton-neutron excitations in which small amplitude quadrupole oscillations form the basic building blocks. These isovector excitations result in a natural way in both shell-model, collective (geometrical and algebraic) and quasi-particle-phonon theoretical approaches. The characteristic fingerprint of strong M1 transitions between MS and symmetric collective states, associated with weak E2 transitions from MS states into the symmetric collective states, has allowed to identify the MS states. They are most clearly observed when a given nucleus only contains a few proton particles (holes) and neutron particle (holes)outside of closed shells. The $Z \approx 40$, $N \approx 50$ region is one of the best studied regions showing besides the lowest MS 2⁺ state, more complex MS $1^+, 2^+, 3^+$ and 4^+ states (Zr, Mo, Ru nuclei). More recently, the presence of MS states has been accumulated in nuclei near $Z \approx 50$ (such as the Cd, Te nuclei), in the rare-earth region (54 < Z < 60, 72 < N < 82) and also for lighter nuclei near the Z=N=28 closed shells. In view of the very general characteristic of these excitations, one can expect them to appear in nuclei adjacent to any doubly-closed shell nucleus.

VII. SCISSORS MODES IN OTHER MANY-BODY SYSTEMS

A. Rotational magnetic excitations in other fermion systems

1. Deformed metallic clusters

Mass spectra for a particular class of metallic clusters were produced more than 20 years ago (Knight et al., 1984). They exhibit large abundance peaks at N = $8, 20, 40, 58, 92, 138, \dots$ These clusters were produced in the expansion of an inert gas (typically Argon or Xenon) through a 0.1-0.2 mm wide nozzle, in which the random thermal motion of the atoms is converted into a uniform translational motion, thereby also causing a cooling of the inert gas. Introducing atomic Na vapor into this system results into large clusters with an overall broad size distribution. Details of the mechanism are described by de de Heer (1993). The peculiar observation for Na clusters, and later shown to exist also for Ag, Au, and Cs, points towards extra stability associated with the delocalized motion of atomic 3s electrons, bound in a spherical potential (see Bjørnholm et al. (1990, 1992); de Heer et al. (1987) for a detailed and extensive discussion). The observation of quantal effects in clusters of atoms established a deep connection between various fields in physics such as electronic motion in atoms and nucleonic motion in nuclei.

The Hamiltonian describing the neutral cluster consisting of N nuclei with Z electrons each, is fully determined through the Coulomb force, but is generally too complex

to be solved exactly. In simple metals, though, such as Ag, Al, Na, ..., the separation into valence electrons and core electrons (well bound and localized) leads to the simplification of xN interacting electrons (x the number of valence electrons per atom) moving in the field caused by the N ions. A further step results in completely ignoring the nuclear motion, thus leading to an electronic Hamiltonian of the form

$$H_{el} = \sum_{k=1}^{xN} \{ \frac{\mathbf{p}_k^2}{2m} + V_I(\mathbf{r}_k) \} + \frac{1}{2} \sum_{l \neq k=1}^{xN} \frac{e^2}{|\mathbf{r}_k - \mathbf{r}_l|}, \quad (38)$$

with the ionic potential $V_I(\mathbf{r}_k)$ defined as

$$V_I(\mathbf{r}_k) = -\sum_{i=1}^N \frac{xe^2}{|\mathbf{r}_k - \mathbf{R}_i|}.$$
 (39)

The latter potential most often is replaced by some pseudo-potential. At the end, a rather drastic but efficient approximation consists of averaging out the ionic structure and replacing the corresponding charge distribution by a "constant" background charge in a finite (spherical, deformed, vibrating and/or rotating) volume. This defines the so-called *jellium* model as used in the description of metallic bulk and surface properties (Brack, 1992; Brack et al., 1991a,b).

Collective dipole excitations are well known in alkalimetal clusters (de Heer, 1993; de Heer $et\ al.$, 1987) which correspond to the classical surface-plasmon oscillations (Ashcroft and Mermin, 1976) of the electron cloud against the positively charged ions forming the cluster. The relative motion of protons versus neutrons in atomic nuclei, giving rise to the electric giant dipole mode is equivalent to the electric dipole mode that results in electron motion in atoms. For deformed metallic clusters, a magnetic excitation of orbital nature was predicted by Lipparini and Stringari (1989b) at an energy much lower than the classical plasmon frequency. A macroscopic illustration of this new magnetic excitation derives from a displacement field ${\bf u}$ of the electron motion in the valence cloud. The suggested form is

$$\mathbf{u} = \hat{\omega}\mathbf{x}\mathbf{r} + \frac{\delta}{1 + \delta/3}\nabla(xy),\tag{40}$$

with $\hat{\omega}$ the unit vector in the z-direction. The cluster is described with a deformed electron density profile

$$\rho_e = \rho_0 \left(\frac{x^2}{R_x^2} + \frac{y^2}{R_y^2} + \frac{z^2}{R_z^2} \right), \tag{41}$$

with $R_x = R_z$ and the deformation δ defined by the expression

$$\delta = \frac{3}{2} \frac{R_y^2 - R_x^2}{R_y^2 + 2R_x^2}. (42)$$

This displacement field satisfies the condition $\nabla \cdot \mathbf{u} = 0$. The first term solely corresponds to a rigid rotation of the electrons with respect to the jellium background and implies a scissors mode (Fig. 53a) with a restoring force that is due to the Coulomb attraction between the electron cloud and the ion positive charged background. This is very much identical to the proton-neutron symmetry term causing the restoring force in case of atomic nuclei. Including the quadrupole term in the displacement field i.e. $\nabla(xy)$, the corresponding motion (Fig. 53b) corresponds to a rotation within a spheroidal rigid surface with a velocity field such that $\mathbf{v} \cdot \mathbf{n} \mid_{surface} = 0$. In the limit

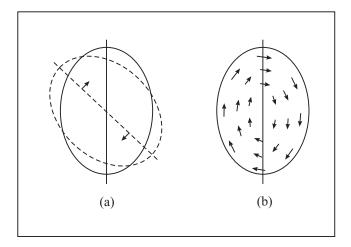


FIG. 53 Displacement field for the magnetic M1 low-lying rotational state corresponding to (a) a rigid rotation of the electrons with respect to the jellium background and (b) a rotation within a rigid surface. The direction of the unit vector $\hat{\omega}$ (z direction) points out of the plane. (Reprinted with permission from Lipparini and Stringari (1989a). ©(2009) Am. Phys. Soc.)

of small deformation, one can determine the frequency of the magnetic mode $\omega_{M1} = \sqrt{K/\theta}$ with K the energy change originating from the displacement field and θ the collective mass parameter (related to the moment of inertia)

$$\omega_{M1} = \delta \sqrt{\frac{4\varepsilon_F}{mr_s^2}} N^{-1/3}. \tag{43}$$

In this section VII, we use throughout the convention $\hbar = c = 1$. In deriving this result, the approximate relation

$$\langle r^2 \rangle = \frac{3}{5} r_s^2 N^{2/3},\tag{44}$$

was used. In the particular case of Na clusters, with $r_s=4$ a.u. and $\varepsilon_F=3.1$ eV, the frequency becomes $\omega_{M1}=\delta 4.6N^{-1/3}$ eV. In the range of clusters with N=10-100 and for typical deformations $\delta=0.2-0.4$, the frequency amounts to 0.2- 0.6 eV, much lower than the dipole plasmon frequency of 3.4 eV. This collective state carries considerable M1 strength $B(M1)\approx\omega_{M1}\theta~\mu_B^2,$ or, using explicit expressions for the frequency and the moment of inertia (Lipparini, 2003), becomes

$$B(M1) = \frac{4}{5}\delta\sqrt{\varepsilon_F m r_s^2} N^{4/3} \mu_B^2, \tag{45}$$

which, in the particular case of Na clusters, reduces to the approximate result $B(M1) \approx \delta N^{4/3} \mu_B^2$ with μ_B the Bohr magneton. The particular results for both ω_{M1} and B(M1) have a very similar form as the expressions derived in Eqs. (2) and (3) for the nuclear case.

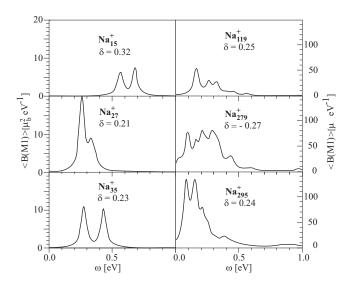


FIG. 54 Energy distribution of the M1 strength for Na clusters ranging from N=15 to N=295. The deformation parameter δ characterizing the specific cluster is also given. (Reprinted with permission from Nesterenko *et al.* (1999). 0(2009) Am. Phys. Soc.)

The above semi-classical results can also be derived starting from microcopic calculations. schematic two-level RPA model has been studied by Lipparini and Stringari (1989a) with results that corroborate the macroscopic approach. More detailed studies of the orbital magnetic dipole mode were carried out by Nesterenko et al. (1999), based on a self-consistent RPA approach (Kleinig et al., 1998; Nesterenko et al., 1997). Using a phenomenological Woods-Saxon potential, the results confirm those obtained by Lipparini and Stringari (1989b) discussed before. We present in Fig. 54 the salient features of the M1 strength distribution as derived for Na clusters, in which the N dependence (moving from N = 15 towards N = 295) becomes particularly clear. In the heavy clusters with $N\sim300,$ the M1 strength reaches very big values of the order of $350-400~\mu_B^2$ with the strength remaining concentrated in a rather narrow energy interval. Therefore, it is in the heavy clusters that the collective nature of the M1 mode becomes particularly clear with a $N^{-1/3}$ frequency dependence as pointed out in Eq. (43). This distinguishes the magnetic mode from the regular electric dipole surface-plasmon mode, which is much better documented experimentally. As was pointed out before, the strength in exciting this magnetic dipole mode is so much weaker than the electric dipole strength that it has not been observed experimentally as yet, contrary to the nuclear case as discussed in the major part of the present paper.

2. Orbital current modes in deformed quantum dots

Recent advances in semiconductor technology have allowed to build nanostructures with varying shapes. In these systems, electrons are laterally confined at the semiconductor boundary and form a two-dimensional (2D) quantum dot (see Alhassid (2000)). Until recently, the vast majority of theoretical and experimental efforts focused on systems with circular symmetry. Many of the properties of these dots are very well accounted for when imposing a parabolic potential as confinement, or, invoking the concept of a jellium disk (see subsection VII.A.1). Experiments addressing deformed nanostructures using Raman scattering and far-infrared spectroscopy (Austing et al., 1999; Demel et al., 1990; Schüller et al., 1996; Sikorski and Merkt, 1989; Strenz et al., 1994) have been at the origin of theoretical studies relaxing on the circular symmetry (Hirose and Wingreen, 1999; Koskinen et al., 1997; Puente and Serra, 1999).

With the breaking of the rotational symmetry of the system (metallic cluster, quantum dot), in particular by introducing a quadrupole distortion, it turns out that a low-energy orbital current excitation (OCE) is generated in the 2D dot with an energy dependence $N^{-1/2}$ in contrast with the $N^{-1/3}$ dependence for three-dimensional system (Lipparini, 2003). Here, N denotes the number of electrons confined in the elliptic quantum dot. Interesting here to remark is the fact that the frequency goes to zero with $N \to \infty$.

One can now perform essentially the same analysis as was used in the description of deformed metallic clusters, but now with an elliptic two-dimensional charge distribution given as

$$\rho_e = \rho_0 \left(\frac{x^2}{R_x^2} + \frac{y^2}{R_y^2} \right), \tag{46}$$

with R_x and R_y the ellipse radii and a deformation parameter η , defined by the expression

$$\eta = \frac{R_y^2 - R_x^2}{R_y^2 + R_x^2}. (47)$$

The kinetic energy density has an expression analogous to the charge distribution given in Eq. (46). An OCE follows, using a displacement field similar to the one discussed in Sec. VII.A.1, but for the 2D case which is expressed as (Serra $et\ al.$, 1999)

$$\mathbf{u} = \hat{\omega} \mathbf{x} \mathbf{r} + \eta \nabla (xy) . \tag{48}$$

The electronic motion becomes a collective rotating flow along the ellipse contour lines. The frequency for this orbital mode turns out to be

$$\omega_{OCE} \approx \frac{\eta}{\sqrt{1 - \eta^2}} \sqrt{\frac{16\varepsilon_F}{3mr_s^2}} N^{-1/2} , \qquad (49)$$

where r_s stands for the Wigner-Seitz radius (Ashcroft and Mermin, 1976). Carrying out the same analysis but retaining only the quadrupole deformation generating displacement term $\mathbf{u} = \nabla(xy)$, the frequency of the corresponding quadrupole charge-density excitation (QCDE) becomes

$$\omega_{QCDE} \approx \sqrt{2}\omega_0,$$
 (50)

where ω_0 is the average of the frequencies in x and y direction for the confining parabolic potential. This frequency is much larger than the frequency of the orbital excitation. A most interesting illustration of the effects of both modes is obtained when evaluating the magnetic orbital response. This M1 response $\langle \tilde{L}_z \rangle$ has been calculated (Serra et al., 1999), as a function of time, by modifying the electron orbitals using the displacement operator of Eq. (48) (left-side part of Fig. 55) as well as the corresponding M1 strength (right-side part of Fig. 55). Here, the cases of (i) a pure rotational perturbation (rotation), (ii) orbital perturbation (OCE or scissors), and (iii) pure quadrupole perturbation are shown, respectively. These results are in line with the simple discussion of the frequencies as presented before. The M1 orbital strength is divided in two distinct regions: one at the higherenergy side, which is associated with the quadrupole distortion and one at the low-energy end, associated with the orbital excitation. An in-depth study is presented by Austing et al. (1999). The distribution of scissors M1strength in these elliptic quantum dots is quite similar to the situation in strongly deformed nuclei in which the M1strength is also separated into a low-energy orbital (the scissors mode) part and the higher-lying $K^{\pi}=1^+$ component of the isovector giant-quadrupole resonance (see Sec. III.C).

Besides these orbital charge-density excitations, it is also possible to describe spin-density oscillations in quantum dots. When the spin components oscillate in phase, they describe the density modes, however, when oscillating out of phase, spin modes can be created. A particular interesting case is obtained as alternating rotation of spin up and spin down densities in opposite direction, This spin-twist mode is very soft (Puente and Serra, 1999), well below the spin dipole oscillation modes.

3. Other Fermi systems

It turns out that scissor modes can also be realized in a superfluid Fermi gas (Minguzzi and Tosi, 2001). Confining the Fermi gas inside a spherical harmonic trap and solving the equations of motion for the density and concentration fluctuations gives rise to a single scissors frequency for the superfluid situation and results in two scissors frequencies in a normal Fermi gas.

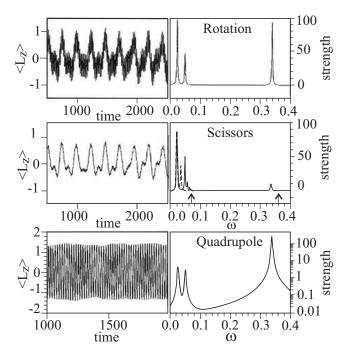


FIG. 55 Results for the time evolution of an elliptic quantum dot with N=20 electrons, for a deformation value of $\eta{=}0.28$, and this using three different initial perturbations: pure rotation, orbital and quadrupole distortions. In the left-side panels, the simulated M1 signal, expressed as $\langle \hat{L}_z \rangle$, is plotted as a function of time. The right-side panels show the corresponding M1 excitation strength. The middle right-side panel also shows the independent electron strength function (dashed line) and the little arrows indicate the position of the solutions given by Eqs. (49) and (50). (Reprinted with permission from Serra et al. (1999). ©(2009) Am. Phys. Soc.)

Recently, Hatada et al. (2005) have suggested the possibility that axially symmetric atoms in crystals with ionic bonding can exhibit a scissors excitation. Its signature in this case is the existence of a low-lying collective excitation resulting from precessing atoms around the anisotropy axis of the crystal cells. This excitation has a magnetic dipole character and could be observed through the absorption of incoming photons which should exhibit a differential dichroism. An extension to cover also crystals with cubic symmetry was presented by Hatada et al. (2009a).

Other ways of observing scissors modes in crystals start from a recent experiment that studied magnetic properties of rare-earth systems (van der Laan et al., 2008). If one considers crystals in which the internal electrostatic field is small with respect to the electron spin-orbit coupling in the atoms, the so-called "spin-orbit locking" situation, an applied external magnetic field will rotate both the spin and charge density profiles simultaneously. Switching off the magnetic field, the atoms will start oscillating around the axes of the crystal cells (Hatada et al., 2009b). Experiments are proposed that may be sensitive enough to detect the photons emitted when deexciting the scissors excitation (Hatada et al.,

2010).

B. Scissors modes of a trapped Bose-Einstein condensate

By now, there exists a vast literature on trapped Bose-Einstein condensates (Giorgini et al., 2008; Pitaevskii and Stringari, 2003). Superfluidity in these condensates is one of the most spectacular consequences. It is, however, not easy to obtain unambiguous evidence for the superfluid characteristics. Because superfluidity will affect the moment of inertia of the trapped condensate (such as a reduction over the classical rigid value), one could expect that a study of rotational properties of such condensates can give rise to experimental evidence for the existence of superfluidity. Guéry-Odelin and Stringari (1999) have studied the oscillatory behavior caused by rotating a condensate with respect to the symmetry axis when trapped in a deformed external potential of parabolic type. They concentrate in particular on the superfluid effects in the condensate. The restoring force associated with such a rotation in the xy-plane is proportional to δ^2 , with the trapping potential given by the expression

$$V_{ext}(\mathbf{r}) = \frac{m}{2}\omega_x^2 x^2 + \frac{m}{2}\omega_y^2 y^2 + \frac{m}{2}\omega_z^2 z^2,$$
 (51)

with moreover

$$\omega_x^2 = \omega_0^2 (1+\delta), \omega_y^2 = \omega_0^2 (1-\delta).$$
 (52)

The mass parameter, determined by the moment of inertia, in the superfluid case becomes proportional to δ^2 too (Lipparini, 2003; Rowe, 1970). As a result, even when the deformation of the external potential approaches zero, the frequency keeps a finite value. It is only in the absence of superfluidity that the moment of inertia regains its rigid value and therefore, a low-frequency will characterize the oscillatory motion. The outcome of the theoretical study (see Guéry-Odelin and Stringari (1999); Lipparini (2003)) is that a sudden rotation of the trap symmetry axis by a small angle θ_0 will perturb the condensate from its equilibrium shape and, if the angle θ_0 is not too large, will start a scissors-like motion³ in the xyplane. This is illustrated in a schematic way in Fig. 56 (right-hand side) in which both the trapping potential and the condensate are drawn. An idea of the dimensions of such condensates is also given. On the left-hand side in Fig. 56, we compare with the analogous situation in strongly deformed atomic nuclei, in which protons and neutrons can give rise to a scissors motion. Here too, the dimension of the system is given in order to stress the large difference in scales but keeping essentially the same

³ The expression scissors mode has been taken over from nuclear physics by the atomic physics and BEC community.

physics. Under the above conditions one obtains

$$\theta(t) = \theta_0 cos(\omega_{sc}t), \tag{53}$$

in which $\omega_{sc}=\sqrt{2}\omega_0$, or $\omega_{sc}=\sqrt{\omega_x^2+\omega_y^2}$. In the absence of a superfluid condensate and entering a high-temperature regime, analytic solutions become possible (Guéry-Odelin and Stringari, 1999) which, for small rotation angles of the trap axis, are described by a differential equation for $\theta(t)$. This equation now allows for different solutions, propagating at high and low frequencies. In a collisionless regime, the higher frequency becomes $\omega_+{=}\omega_x+\omega_y$ and can be identified with an irrotational quadrupole oscillation. The lower frequency $\omega_-{=}|\;\omega_x-\omega_y\;|$, however, corresponds to the rotational mode of the system, a component which is absent in the superfluid case.

Protons Protons Neutrons Trapped BEC Bose-Einstein Condensate Trap Potential

FIG. 56 Schematic comparison of the scissors motion in atomic nuclei with protons and neutrons generating the scissors motion, on a length scale of 10 fm (left side), and in Bose-Einstein condensates trapped in an external potential, the latter acting on a length scale of 100 μ m (right side).

Maragò et al. (2001, 2000) have reported on a clear observation of the scissors mode of a Bose-Einstein condensed gas of ⁸⁷Rb atoms brought into a magnetic trap. The authors highlighted the very importance of the discovery of the scissors mode in atomic nuclei stating "The experimental discovery of the scissors mode (Bohle et al., 1984b), first predicted in a geometrical model (Lipparini and Stringari, 1983; Lo Iudice and Palumbo, 1978), has been one of the most exciting findings in nuclear physics during the past two decades".

The scissors mode has been excited by a sudden rotation of the deformed trapping potential (with the constraints of having $\omega_x \approx \omega_z > \omega_y$). The condensate, cooled to well below the critical temperature for the Rb gas contains of the order of 10^4 atoms. The time evolution of the scissors oscillation (Fig. 57) exhibits a single undamped mode corresponding to 265.6 ± 0.8 Hz, which agrees perfectly with the theoretical value of 265 ± 2 Hz as deduced from $\omega_{sc} = \sqrt{\omega_x^2 + \omega_y^2}$. This observation thus provides an unambiguous demonstration of the superfluid nature the Rb condensate. In Fig. 57, we show the classical frequency and the tilt angle for the trapped Bose-

Einstein condensate of ⁸⁷Rb atoms, which correspond to $\nu = 265.6 \pm 0.8 \text{ s}^{-1}$ and $\alpha = 7.2^{\circ}$, respectively.

It is instructive to compare the condensate result with the scissors mode in the deformed nucleus ¹⁵⁶Gd (Bohle et al., 1984b). We can extract both the frequency (using the experimental energy of the 1⁺ state at E_x =3.075 MeV) and a classical tilt angle identifying the restoring force with the symmetry energy and making use of the expression $\alpha = (CJ_{intr}/\hbar^2)^{-1/4}$ deduced from the proton-neutron collective model (De Franceschi et al., 1984; Nojarov et al., 1986) discussed in Sec. III.A.2. This results in $\nu = 7 \times 10^{20} \,\mathrm{s}^{-1}$ and $\alpha \sim 6^{\circ}$, respectively. Thus, comparable tilt angles are found characterizing the scissors motion, albeit in different regimes of physics. A striking difference between the two systems, however, is the fact that in the condensate the deformation of the trap can be varied by choosing appropriate frequencies as well as the temperature giving a large range to study the intricate properties of quantum fluids and their transition from the irrotational (superfluid case) to the rotational regime. In the atomic nucleus, the deformation is fixed for a given (Z, N) combination.

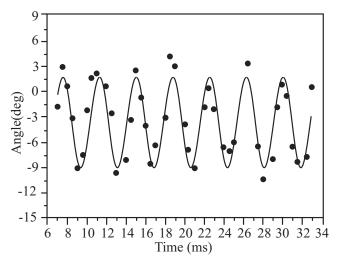


FIG. 57 The classical frequency and the tilt angle for a trapped Bose-Einstein condensate of 87 Rb atoms. (Reprinted with permission from Maragò *et al.* (2000). ©(2009) Am. Phys. Soc.)

Compared to the observation of scissors modes for trapped Bose-Einstein condensates, scissors-like oscillations have also been observed in a quantum-degenerate mixture of two such condensates consisting of different atomic species, i.e., ⁴¹K and ⁸⁷Rb (Modugno *et al.*, 2002). In this case, the scissors mode is induced by interactions between the two atomic species present in the trap. Therefore, this situation is even closer to the nuclear physics case where proton and neutron fluids are present.

C. Rounding up

The above theoretical and experimental studies in both fermionic (metallic clusters, quantum dots, Fermi gas, crystals) systems in which the electronic motion in some 'external' potential has been studied and in bosonic systems (Bose-Einstein condensates in the superfluid phase trapped in an external anisotropic potential) show the existence of scissors modes with properties in line of those observed in atomic nuclei. The arguments can even be turned around: the observation of low-lying scissors modes in these systems is a proof of the superfluid characteristic of proton and and neutron fluids in the nuclear rotational motion.

VIII. CONCLUSIONS AND OUTLOOK

A. Conclusions

The study of the response of nucleons, moving inside the atomic nucleus, to external (electromagnetic and hadronic) probes in the magnetic dipole channel, discussed in depth in the present article, can be crystallized and summarized in a succinct way. Here, we take Figs. 1-3 as our guide.

- (i) At the lower energy side, orbital nucleonic motion is strongly excited by the $\hat{L}_{\pi} - \hat{L}_{\nu}$ part of the M1 operator (contra-rotational or scissors-like motion of protons versus neutrons). This does not show up, however, as a single collective state but the orbital strength is fragmented within a rather limited energy interval between 2.5 and 4 MeV. In light nuclei, the strength distribution looks more simple since it is contained mainly in the $2^+(\pi) \otimes 2^+(\nu)$, $4^+(\pi) \otimes 4^+(\nu)$, ... configurations. In strongly deformed rare-earth nuclei, there is more spreading of strength which has been detected using (e, e') and (γ, γ') reactions and not or very weakly with proton scattering. A number of interesting results are connected to the fact that the non-energy weighted M1 sum rule strength in the interval 2.5-4 MeV exhausts the larger part of the orbital M1strength. Moreover, it was discovered that this summed M1 strength correlates strongly with deformation and thus with the $B(E2; 0_1^+ \rightarrow 2_1^+)$ strengths, indicating saturation when progressing through the strongly deformed rare-earth region.
- (ii) Spin M1 strength is concentrated at higher excitation energies because the spin-flip part of the M1 operator is mainly a $1\hbar\omega$ excitation in the shell model and has been studied throughout the region of deformed rarearth and actinide nuclei. In light nuclei, the strength is particularly associated with the spin-flip transition between spin-orbit partners. In the heavier nuclei, it is the spin-spin isospin dependent part of the residual interaction $(\vec{\sigma} \cdot \vec{\sigma} \vec{\tau} \cdot \vec{\tau})$ that rules the concentration or fragmentation of spin M1 strength.
- (iii) At still higher excitation energies $(2\hbar\omega)$, theoretical predictions indicate the presence of a $K^{\pi} = 1^{+}$

component of the isovector giant quadrupole resonance (IVGQR), a state that could rightly be associated with a collective scissors mode. Because of the high excitation energy, no systematic experimental studies exist yet exploring such a mode. Knowledge of the IVGQR strength, connected to this response, allows the evaluation of an energy-weighted sum rule to constrain the orbital M1 strength, which is exhausted to more than 80%.

- (iv) In nuclei with just a few valence protons and neutrons outside of closed shells, it has been experimentally proven that low-lying mixed-symmetry (isovector) 2^+ excitations exist. They are characterized by strong M1 decay into the first excited 2^+_1 state, a symmetric (isoscalar) mode in the proton and neutron motion, accompanied by weak E2 decay into the 0^+ ground state. The quadrupole degree of freedom dominates the low-energy structure, resulting in energy spectra with a vibrational character. However, isovector combinations can also give rise to a 1^+ state which is related to the scissors 1^+ state as it appears in strongly deformed nuclei.
- (v) In odd-mass nuclei, the study of the M1 strength distribution is more complicated because of the odd-particle (-hole) coupling to the 1^+ modes which induces a large fragmentation. There occurred a number of early problems in accounting for the observation of the M1 strength but the present situation is such that, through very careful studies of the highly-fragmented background structures, the full summed M1 strength in odd-mass nuclei is consistent with the summed M1 strength obtained in the adjacent even-even nuclei. Much more work needs to be done from the side of theoretical studies in order to understand the major mechanisms that can explain the fragmentation and the sometimes sudden important changes in the observed fragmentation when going from nucleus to nucleus.
- (vi) In the even-even nuclei, a number of general features have resulted from detailed experimental studies of the M1 response over many nuclei, spanning the region from light to very heavy nuclei. One of the most important observations is a strong correlation of the orbital magnetic dipole response with other multipoles, in particular with the E2 strength but also, be it in an indirect way, with the nuclear charge radii, i.e. with the E0 strength. This connection has been formulated in a more quantitative way using various M1sum rules that are proportional to the ground-state expectation value of the quadrupole-quadrupole force and thus lead to new E2 sum rules. A former review article by Lipparini and Stringari (1989a) already pointed out such a connection. The microscopic understanding of this intimate relation between the M1 properties on one side and the quadrupole and monopole properties on the other side are more indirect. It has been shown that the quadrupole deformation of the nucleus spreads the individual single-particle states (breaking the spherical symmetry) and, with pairing included, the shell-model description of M1 strength is indirectly connected to the E2 ground-state deformation characteristics of the nu-

cleus. This gives at the same time a natural explanation of the M1 orbital summed strength saturation in the midshell region of the rare-earth nuclei. It is a consequence of the fact that the deformed rare-earth nuclei also exhibit a saturation in the quadrupole deformation value (about $\delta \simeq 0.25-0.30$).

(vii) At the present stage, the magnetic M1 strength seems to be rather well understood and there appears a clear-cut concentration of orbital strength at lower energies using various theoretical approaches: shell-model, QRPA and QPM calculations, collective geometric and algebraic models all account rather well with the experimental situation on a qualitative level. This is a highly comforting situation with respect to the theoretical understanding of how M1 strength builds up at the lower energy side as mainly orbital in character. Collective model approaches, however, by their specific nature, fail in accounting for the detailed fragmentation. This holds for both the geometrical and algebraic collective models. In order to bridge the gap between the various theoretical approaches, one will have to address and explore the possibilities of large-scale shell-model calculations as a way to extract collective effects starting from a microsocpic basis.

B. Outlook and future perspectives

Having formulated a number of concluding elements on how the nucleus responds when excited with electromagnetic and hadronic probes in the magnetic dipole channel, a number of very clear-cut problems remain to be solved in the coming years. We also present a number of topics that will deserve intensive thought and future experimental efforts such as to bring us closer to a quantitative level of understanding the magnetic dipole channel (and some related other multipoles).

- (i) It will be particularly interesting to connect the scissors 1⁺ mode, which was shown to be strongly excited using electron and photon scattering off deformed nuclei, to the recent observation of 1⁺ states in nuclei when approaching closed shells. In the latter nuclei, the quadrupole mode determines the dominant low-lying proton-neutron isoscalar and isovector excitations. The exploration of a full class of isovector (also called mixedsymmetry states) excitations with spins ranging from 0^+ to 4⁺ has begun but is clearly in need of still more systematic studies. In this quest, it is of utmost importance to combine as many complementary probes as possible (Coulomb excitation, lifetime measurements, detailed γ spectroscopy measuring branching ratios and $\delta(E2/M1)$ mixing ratios,...) so as to be able to pin down the nuclear wave functions of these states.
- (ii) The spin-flip magnetic dipole response that has been studied using (polarized) proton scattering off nuclei in the energy region 5-10 MeV poses some specific challenges. A particularly important and unsolved problem, at present, concerns the precise character of the two

humps observed in the spin-flip M1 strength distribution in many deformed nuclei. Theoretical studies come to opposing conclusions and only a more detailed experimental survey of the precise charge character of the two humps (proton/neutron versus isoscalar/isovector) can resolve this issue. From the theoretical side and with the present-day numerical capacities to perform large-scale shell-model calculations, a precise survey of the M1 response in both the sd-shell nuclei, in particular along the N = Z line with nuclei like ²⁰Ne, ²⁴Mg, ···, ³⁶Ar, ⁴⁰Ca as well as in the heavier fp-shell nuclei would be very important, also in the light of existing high-quality data in some of these light and medium-heavy nuclei. Recent experimental developments (Tamii et al., 2009), which combine for the first time high-resolution measurements of the $(\vec{p}, \vec{p'})$ reaction at proton energies of several hundred MeV with measurements at 0°, should allow to tackle all of these questions. Indeed, the feasibility of such experiments has been proven for the heaviest nuclei as demonstrated in Fig. 58 for the example of ²⁰⁸Pb (von Neumann-Cosel et al., 2009).

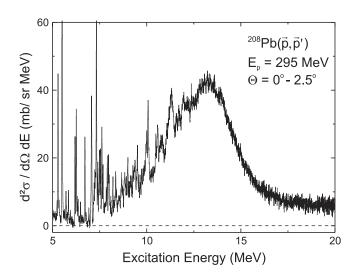


FIG. 58 Spectrum of the $^{208}{\rm Pb}(\vec{\rm p},\vec{\rm p}')$ reaction at E=295 MeV and $\Theta=0^{\circ}$ (von Neumann-Cosel et~al.,2009) measured with an energy resolution $\Delta E\simeq25~{\rm keV}$ (full width at half maximum) .

- (iii) The spin-flip magnetic dipole response which is connected to the axial-vector part of the Gamow-Teller operator, may well be extended towards higher excitation energies by using the fact that the spin-flip part probed in proton scattering is also connected to the neutrino (axial-vector term) scattering contribution. So, proton scattering may provide valuable information concerning neutrino scattering off nuclei in the giant resonance region and, subsequently, give insight in neutrino-nucleus cross-sections and their importance in supernova processes.
- (iv) There are other magnetic modes to be explored besides the dipole one, which formed the major part of the present review. There has been a search for magnetic quadrupole excitations of $J^{\pi}=2^{-}$ states in 48 Ca

and 90 Zr using high-resolution backward angle inelastic electron scattering (von Neumann-Cosel et al., 1999), extended to 58 Ni (Reitz et al., 2002). Macroscopically, the orbital M2 mode can be viewed as a vibrational counter-rotation of different fluid layers in the upper and lower hemisphere, hence the name "twist" mode (Holzwarth and Eckart, 1977, 1979). Like for the scissors mode, occurrence of an orbital M2 mode is a general feature of quantum many-body systems as discussed e.g. for the cases of metallic clusters by Nesterenko et al. (2000) and ultracold Fermi gases by Viñas et al. (2001). Therefore this mode deserves more intensive and systematic exploration.

(v) The proportionality of the summed M1 strength to the nuclear deformation, more in particular showing a δ^2 dependence, opens the possibility to make use of measured magnetic dipole strengths as a new fingerprint for exploring sudden shape-phase transitions. Since there appear a number of regions in the nuclear mass table that look like potential places where sudden changes in nuclear shape may occur, an in depth study of magnetic dipole excitations in those regions can lead to extra information.

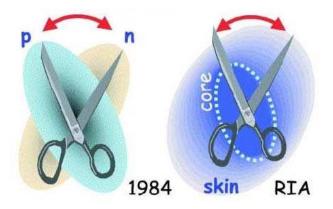


FIG. 59 (Color online) Schematic picture of the scissors mode in stable deformed nuclei (left) as compared to exotic nuclei with a weakly-bound neutron skin (right), taken from (RIA, 2003).

(vi) Last, but not least, a point of current interest is to study the scissors mode moving away from the valley of β -stability. Because the nuclei to be explored are highly unstable, one will have to resort to reactions in inverse kinematics such as Coulomb excitation, inverse proton scattering utilizing radioactive ion beam facilities like RIKEN, FAIR-GSI, SPIRAL II, and further on, FRIB. New phenomena are expected (RIA, 2003) in very neutron-rich nuclei when coupling a weakly bound neutron skin to a well-bound deformed core (cf. Fig. 59). A soft scissors mode due to the presence of a neutron skin, analogous to the soft electric dipole mode, has been proposed within both a geometrical approach by Van Isacker et al. (1992) at an energy approximately half of the scissors mode energy, and also within the IBM-2 (Caprio and Iachello, 2004; Warner and Van Isacker, 1997). Considering the plans at FAIR, one may even think about electron scattering on unstable nuclei. These and more exotic possibilities may materialize at future RIB facilities, as discussed e.g. in recent report on upgrading the NSCL facility at MSU (RIA, 2006). The new generation facilities should also allow to systematically explore mixed-symmetry states in the $N\approx 50$ and $N\approx 82$ region, by providing beams of high enough luminosity.

To sum up, in the present review we have shown that (i) the 1^+ orbital scissors mode at low excitation energy, (ii) the 1^+ spin-flip mode at higher excitation energy, and, (iii) the 1^+ component originating from the still higher-lying isovector quadrupole mode are universal realizations of the nuclear many-body system. It turns out that these modes are not specific to nuclei only. On the contrary, the scissors mode also shows up in other many-body systems. There is compelling evidence for M1 excitations in deformed metallic clusters, in elliptically deformed quantum dots and in Bose-Einstein condensates of superfluid nature, and we have stressed the fascinating interplay between nuclear physics and other highly correlated many-body systems.

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References

Abdelaziz, M., and J. P. Elliott, 1987, Phys. Lett. B 197, 505.
Abdelaziz, M., M. J. Thompson, J. P. Elliott, and J. A. Evans, 1988, J. Phys. G 14, 219.

Ahmad, S. T., W. D. Hamilton, P. V. Isacker, S. A. Ahmad, and S. J. Robinson, 1989, J. Phys. G 15, 93.

Ahn, T., M. Coqard, N. Pietralla, G. Rainowski, A. Costin, R. F. V. Janssens, C. J. Lister, M. Carpenter, S. Zhu, and K. Heyde, 2009, Phys. Lett. B 679, 19.

- Ahn, T., N. Pietralla, G. Rainovski, A. Costin, K. Dusling, T. C. Li, A. Linnemann, and S. Pontillo, 2007, Phys. Rev. C 75, 014313.
- Alaga, K. A. A. B., G., and B. Mottelson, 1955, Dan. Mat. Fys. Medd. 29(9), 1.
- Alhassid, Y., 2000, Rev. Mod. Phys. 72, 895.
- Allaart, K., E. Boeker, G. Bonsignori, M. Savoia, and Y. K. Gambhir, 1988, Phys. Rep. **169**, 209.
- Arima, A., M. Harvey, and K. Shimizu, 1969, Phys. Lett. B **30**, 517.
- Arima, A., and F. Iachello, 1975a, Phys. Lett. B 57, 39.
- Arima, A., and F. Iachello, 1975b, Phys. Rev. Lett. 35, 1069.
 Arima, A., T. Otsuka, F. Iachello, and I. Talmi, 1977, Phys. Lett. B 66, 205.
- Arima, A., K. Shimizu, W. Bentz, and H. Hyuga, 1987, Advances in Nuclear Physics, volume 18 (Plenum, New York).
- Peña Arteaga, D., and P. Ring, 2007, Prog. Part. Nucl. Phys. 59, 314.
- Peña Arteaga, D., and P. Ring, 2008, Phys. Rev. C 77, 034317.
- Ashcroft, E. W., and N. D. Mermin, 1976, Solid State Physics (Saunders College, Philadelphia).
- Auerbach, N., 1983, Phys. Rep. 98, 273.
- Austing, D. G., S. Sasaki, S. Tarucha, S. M. Reimann, M. Koskinen, and M. Manninen, 1999, Phys. Rev. B 60, 11514.
- Balbutsev, E. B., and P. Schuck, 2004, Nucl. Phys. A **731**, 256.
- Balbutsev, E. B., and P. Schuck, 2007, Ann. Phys. 322, 489.
 Bandyopadhyay, D., C. C. Reynolds, C. Fransen,
 N. Boukharouba, M. T. McEllistrem, and S. W. Yates,
 2003, Phys. Rev. C 67, 034319.
- Bauske, I., J. M. Arias, P. von Brentano, A. Frank,
 H. Friedrichs, R. D. Heil, R.-D. Herzberg, F. Hoyler,
 P. Van Isacker, U. Kneissl, J. Margraf, H. H. Pitz, et al.,
 1993, Phys. Rev. Lett. 71, 975.
- Baznat, M. I., and N. Y. Pyatov, 1975, Sov. J. Nucl. Phys. **21**, 365.
- Berg, U. E. P., 1984, J. Phys. (France) C 4, 359.
- Berg, U. E. P., K. Ackermann, K. Bangert, C. Blasing, W. Naatz, R. Stock, K. Wienhard, M. K. Brussel, T. E. Chapuran, and B. H. Wildenthal, 1984a, Phys. Lett. B 140, 191.
- Berg, U. E. P., C. Blasing, J. Drexler, R. D. Heil, U. Kneissl,
 W. Naatz, R. Ratzek, S. Schennach, R. Stock, T. Weber,
 H. Wickert, B. Fischer, et al., 1984b, Phys. Lett. B 149,
 59
- Berg, U. E. P., and U. Kneissl, 1987, Annu. Rev. Nucl. Part. Sci. 37, 33.
- Bertsch, G. F., and I. Hamamoto, 1982, Phys. Rev. C **26**, 1323.
- Bes, D. R., and R. A. Broglia, 1984, Phys. Lett. B 137, 141.
 Bettermann, L., C. Fransen, S. Heinze, J. Jolie, A. Linnemann, D. Mücher, W. Rother, T. Ahn, A. Costin, N. Pietralla, and Y. Luo, 2009, Phys. Rev. C 79, 034315.
- Beuschel, T., J. P. Draayer, D. Rompf, and J. G. Hirsch, 1998, Phys. Rev. C 57, 1233.
- Beuschel, T., J. G. Hirsch, and J. P. Draayer, 2000, Phys. Rev. C **61**, 054307.
- Bjørnholm, S., J. Borggreen, O. Echt, K. Hansen, J. Pedersen, and H. D. Rasmussen, 1990, Phys. Rev. Lett. 65, 1627.
- Bjørnholm, S., J. Borggreen, K. Hansen, T. Martin, H. D. Rasmussen, and J. Pedersen, 1992, in *Ecole Internationale Joliot-Curie de Physique Nucleaire*, edited by Y. Abgrall

- (IN2P3), p. 311.
- Boelaert, N., A. Dewald, C. Fransen, J. Jolie, A. Linnemann, B. Melon, O. Möller, N. Smirnova, and K. Heyde, 2007a, Phys. Rev. C 75, 054311.
- Boelaert, N., N. Smirnova, K. Heyde, and J. Jolie, 2007b, Phys. Rev. C **75**, 014316.
- Bohigas, O., 1991, in Les Houches Summer School on Theoretical Physics, Session L II: Chaos and Quantum Physics, edited by M. J. Giannoni, A. Voros, and J. Zinn-Justin (North-Holland, Amsterdam), p. 87.
- Bohle, D., G. Kilgus, A. Richter, C. W. de Jager, and H. de Vries, 1987a, Phys. Lett. B 195, 326.
- Bohle, D., G. Küchler, A. Richter, and W. Steffen, 1984a, Phys. Lett. B 148, 260.
- Bohle, D., A. Richter, U. E. P. Berg, J. Drexler, R. D. Heil, U. Kneissl, H. Metzger, R. Stock, B. Fischer, H. Hollick, and D. Kollewe, 1986, Nucl. Phys. A 458, 205.
- Bohle, D., A. Richter, C. W. de Jager, and H. de Vries, 1987b, Z. Phys. A 328, 463.
- Bohle, D., A. Richter, W. Steffen, A. E. L. Dieperink, N. L. Iudice, F. Palumbo, and O. Scholten, 1984b, Phys. Lett. B 137, 27.
- Bohr, A., and B. Mottelson, 1969, *Nuclear Structure*, volume 1 (Benjamin, New York).
- Bohr, A., and B. Mottelson, 1975, *Nuclear Structure*, volume 2 (Benjamin, New York).
- Brack, M., 1992, in *Ecole Internationale Joliot-Curie de Physique Nucleaire*, edited by Y. Abgrall (IN2P3), p. 325.
- Brack, M., O. Henzken, and K. Hanzen, 1991a, Z. Phys. D 19, 51.
- Brack, M., O. Henzken, and K. Hanzen, 1991b, Z. Phys. D **21**, 65.
- von Brentano, P., J. Eberth, J. Enders, L. Esser, R.-D. Herzberg, N. Huxel, H. Meise, P. von Neumann-Cosel, N. Nicolay, N. Pietralla, H. Prade, J. Reif, et al., 1996, Phys. Rev. Lett. **76**, 2029.
- von Brentano, P., A. Zilges, N. V. Zamfir, and R.-D. Herzberg, 1994, in *Symmetries in Science VII*, edited by B. Gruku and T. Otsuka (Plenum Press, New York), p. 123.
- Brown, B. A., and B. H. Wildenthal, 1987, Nucl. Phys. A 474, 290.
- Brown, B. A., and B. H. Wildenthal, 1988, Annu. Rev. Nucl. Part. Sci. 38, 29.
- Brown, G. E., and M. Bolsterli, 1959, Phys. Rev. Lett. 3, 472.
 Brussaard, P. J., and P. W. M. Glaudemans, 1977, Shell-model Applications in Nuclear Spectroscopy (North-Holland, Amsterdam).
- Burda, O., N. Botha, J. Carter, R. W. Fearick, S. V. Förtsch, C. Fransen, H. Fujita, J. D. Holt, M. Kuhar, A. Lenhardt, P. von Neumann-Cosel, R. Neveling, et al., 2007, Phys. Rev. Lett. 99, 092503.
- Caprio, M. A., and F. Iachello, 2004, Phys. Rev. Lett. 93, 242502.
- Casten, R. F., D. S. Brenner, and P. E. Haustein, 1987, Phys. Rev. Lett. 58, 658.
- Casten, R. F., K. Heyde, and A. Wolf, 1988, Phys. Lett. B 208, 33.
- Caurier, E., K. Langanke, G. Martínez-Pinedo, F. Nowacki, and P. Vogel, 2001, Phys. Lett. B 522, 240.
- Chaves, L., and A. Poves, 1986, Phys. Rev. C 34, 1137.
- Crawley, G. M., N. Anantaraman, A. Galonsky, C. Djalali, N. Marty, M. Morlet, A. Willis, J. C. Jourdain, and P. Kitching, 1982, Phys. Rev. C 26, 87.
- De Coster, C., and K. Heyde, 1989, Phys. Rev. Lett. 63, 2797.

- De Coster, C., and K. Heyde, 1991a, Phys. Rev. Lett. **66**, 2456.
- De Coster, C., and K. Heyde, 1991b, Nucl. Phys. A 529, 507.
 De Coster, C., K. Heyde, and A. Richter, 1992, Nucl. Phys. A 542, 375.
- De Coster, C., K. Heyde, S. Rombouts, and A. Richter, 1995, Phys. Rev. C 51, 3510.
- De Franceschi, G., F. Palumbo, and N. Lo Iudice, 1983, Lett. Nuovo Cimento 37, 61.
- De Franceschi, G., F. Palumbo, and N. Lo Iudice, 1984, Phys. Rev. C 29, 1496.
- Delorme, J., M. Ericson, A. Figureau, and C. Thévenet, 1976, Ann. Phys. (N.Y.) 102, 273.
- Demel, T., D. Heitmann, P. Grambow, and K. Ploog, 1990, Phys. Rev. Lett. **64**, 788.
- Devi, Y. D., and V. K. B. Kota, 1992a, Phys. Lett. B 287, 9.
 Devi, Y. D., and V. K. B. Kota, 1992b, Nucl. Phys. A 541, 173.
- Devi, Y. D., and V. K. B. Kota, 1996, Nucl. Phys. A 600, 20.Dieperink, A. E. L., 1983, Prog. Part. Nucl. Phys. 9, 121.
- Djalali, C., N. Marty, M. Morlet, A. Willis, J. C. Jourdain, N. Anantaraman, G. M. Crawley, A. Galonsky, and P. Kitching, 1982, Nucl. Phys. A 388, 1.
- Eid, S. A. A., W. D. Hamilton, and J. P. Elliott, 1986, Phys. Lett. B 166, 267.
- Eisenberg, J. M., and W. Greiner, 1987, Nuclear Theory, Vol. 1: Nuclear Models (North-Holland, Amsterdam), 3rd edition.
- Elhami, E., J. N. Orce, S. Mukhopadhyay, S. N. Choudry, M. Scheck, M. T. McEllistrem, and S. W. Yates, 2007, Phys. Rev. C 75, 011301(R).
- Elhami, E., J. N. Orce, M. Scheck, S. Mukhopadhyay, S. N. Choudry, M. T. McEllistrem, S. W. Yates, C. Angell, M. Boswell, B. Fallin, C. R. Howell, A. Hutcheson, et al., 2008, Phys. Rev. C 78, 064303.
- Elliott, J. P., 1985, Rep. Prog. Phys. 48, 171.
- Elliott, J. P., 1990, Prog. Part. Nucl. Phys. 25, 325.
- Elliott, P. J., 1958, Proc. Roy. Soc. A 245, 128;562.
- Elliott, P. J., 1963, Proc. Roy. Soc. A 272, 557.
- Enders, J., T. Guhr, N. Huxel, P. von Neumann-Cosel, C. Rangacharyulu, and A. Richter, 2000, Phys. Lett. B 486, 273.
- Enders, J., N. Huxel, U. Kneissl, P. von Neumann-Cosel, H. H. Pitz, and A. Richter, 1998, Phys. Rev. C 57, 996.
- Enders, J., N. Huxel, P. von Neumann-Cosel, and A. Richter, 1997, Phys. Rev. Lett. 79, 2010.
- Enders, J., H. Kaiser, P. von Neumann-Cosel, C. Rangacharyulu, and A. Richter, 1999, Phys. Rev. C 59, R1851.
- Enders, J., P. von Neumann-Cosel, C. Rangacharyulu, and A. Richter, 2005, Phys. Rev. C 71, 014306.
- Ericson, T., and W. Weise, 1988, Pions and Nuclei, International Series of Monographs on Physics, Vol. 74 (Clarendon, Oxford).
- Faessler, A., 1966, Nucl. Phys. 85, 653.
- Faessler, A., Z. Bochnacki, and R. Nojarov, 1986, J. Phys. G 12, L47.
- Faessler, A., and R. Nojarov, 1986, Phys. Lett. B 166, 367.Faessler, A., and R. Nojarov, 1987, Prog. Part. Nucl. Phys.
- Faessler, A., and R. Nojarov, 1988, Phys. Lett. B 215, 439.
- Faessler, A., R. Nojarov, and T. Taigel, 1989, Nucl. Phys. A 492, 105.
- Fagg, L. W., 1975, Rev. Mod. Phys. 47, 683.
- Fazekas, B., T. Belgya, G. Molnar, A. Veres, R. A. Gatenby,

- S. W. Yates, and T. Otsuka, 1992, Nucl. Phys. A **548**, 249. Fearick, R. W., G. Hartung, K. Langanke, G. Martínez-Pinedo, P. von Neumann-Cosel, and A. Richter, 2003, Nucl. Phys. A **727**, 41.
- Fearick, R. W., P. von Neumann-Cosel, A. Richter, S. J. Q. Robinson, and L. Zamick, 2006, J. Phys. Soc. Japan 75, 094201.
- Foltz, C. W., D. I. Sober, L. W. Fagg, H. D. Gräf, A. Richter, E. Spamer, and B. A. Brown, 1994, Phys. Rev. C 49, 1359.
- Frank, A., J. M. Arias, and P. Van Isacker, 1991, Nucl. Phys. A 531, 125.
- Fransen, C., N. Pietralla, Z. Ammar, D. Bandyopadhyay, N. Boukharouba, P. von Brentano, A. Dewald, J. Gableske, A. Gade, J. Jolie, U. Kneissl, S. R. Lesher, et al., 2003, Phys. Rev. C 67, 024307.
- Fransen, C., N. Pietralla, P. von Brentanon, A. Dewald, J. Gableske, A. Gade, A. Lisetsky, and V. Werner, 2001, Phys. Lett. B 508, 219.
- Freeman, S. J., R. Chapman, J. L. Durell, M. A. C. Hotchkis, F. Khazaie, C. J. Lisle, J. N. Mo, A. M. Bruce, R. A. Cunningham, P. V. Drumm, D. D. Warner, and J. D. Garrett, 1989, Phys. Lett. B 222, 347.
- Frekers, D., 2006, Prog. Part. Nucl. Phys. 57, 217.
- Frekers, D., H. J. Wörtche, A. Richter, R. Abegg, R. E. Azuma, A. Celler, C. Chan, T. E. Drake, R. Helmer, K. P. Jackson, J. D. King, C. A. Miller, et al., 1990, Phys. Lett. B 244, 178.
- Fujita, Y., B. Rubio, T. Adachi, F. Molina, A. Algora, G. P. A. Berg, P. von Brentano, J. Buscher, T. Cocolios, D. D. Frenne, C. Fransen, H. Fujita, et al., 2008, J. Phys. G 35, 014041.
- Gade, A., D. Belic, P. von Brentano, C. Fransen, H. von Garrel, J. Jolie, U. Kneissl, C. Kohstall, A. Linnemann, H. H. Pitz, M. Scheck, F. Stedile, et al., 2003, Phys. Rev. C 67, 034304.
- Gade, A., H. Klein, N. Pietralla, and P. von Brentano, 2002, Phys. Rev. C 65, 054311.
- Gade, A., N. Pietralla, P. von Brentano, D. Belic, C. Fransen, U. Kneissl, C. Kohstall, A. Linnemann, H. H. Pitz, M. Scheck, N. A. Smirnova, F. Stedile, et al., 2004, Phys. Rev. C 69, 054321.
- Gade, A., I. Wiedenhöver, J. Gableske, A. Gelberg, H. Meise, N. Pietralla, and P. von Brentano, 2000, Nucl. Phys. A 665, 268.
- von Garrel, H., P. von Brentano, C. Fransen, G. Friessner, N. Hollmann, J. Jolie, F. Käppeler, L. Käubler, U. Kneissl, C. Kohstall, L. Kostov, A. Linnemann, et al., 2006, Phys. Rev. C 73, 054315.
- Garrett, J. D., J. Q. Robinson, A. J. Foglia, and H. Q. Jin, 1997, Phys. Lett. B 392, 24.
- Garrett, P. E., H. Lehmann, C. A. McGrath, M. Yeh, and S. W. Yates, 1996, Phys. Rev. C 54, 2259.
- Garrido, E., E. Moya de Guerra, P. Sarriguren, and J. M. Udias, 1991, Phys. Rev. C 44, R1250.
- Georgii, R., T. von Egidy, J. Klora, H. Lindner, U. Mayerhofer, J. Ott, W. Schauer, P. von Neumann-Cosel, A. Richter, C. Schlegel, R. Schultz, V. A. Khitrov, et al., 1995, Nucl. Phys. A **592**, 307.
- Ginocchio, J. N., 1991, Phys. Lett. B 265, 6.
- Ginocchio, J. N., 1999, Phys. Rev. C 59, 2487.
- Ginocchio, J. N., 2005, Phys. Rep. 414, 165.
- Ginocchio, J. N., and A. Leviatan, 1997, Phys. Rev. Lett. 79, 813.
- Giorgini, S., L. P. Pitaevskii, and S. Stringari, 2008, Rev.

- Mod. Phys. 80, 1215.
- Govaert, K., W. Mondelaers, E. Jacobs, D. D. Frenne, K. Persijn, S. Pommé, M. L. Yoneama, R.-D. Herzberg, S. Lindenstruth, K. Huber, A. Jung, B. Starck, et al., 1994, Nucl. Instrum. Methods A 337, 265.
- Greiner, W., 1965, Phys. Rev. Lett. 14, 599.
- Greiner, W., 1966, Nucl. Phys. 80, 417.
- Grundey, T., A. Richter, G. Schrieder, E. Spamer, and W. Stock, 1981, Nucl. Phys. A 357, 269.
- Moya de Guerra, E., and L. Zamick, 1993, Phys. Rev. C **47**, 2604
- Guéry-Odelin, D., and S. Stringari, 1999, Phys. Rev. Lett. 83, 4452.
- Guhr, T., H. Diesener, A. Richter, C. W. de Jager, H. de Vries, and P. K. A. de Witt Huberts, 1990, Z. Phys. A 336, 159.
- Guhr, T., A. Müller-Groeling, and H. A. Weidenmüller, 1998, Phys. Rep. 299, 189.
- Guliyev, E., A. A. Kuliev, P. von Neumann-Cosel, and A. Richter, 2002, Phys. Lett. B 532, 173.
- Hagberg, E., T. K. Alexander, I. Neeson, V. T. Koslowsky, G. C. Ball, G. R. Dyck, J. S. Forster, J. C. Hardy, J. R. Leslie, H.-B. Mak, H. Schmeing, and I. S. Towner, 1994, Nucl. Phys. A 571, 555.
- Halse, P., 1990, Phys. Rev. C 41, 2340.
- Halse, P., 1991a, Nucl. Phys. A 526, 152.
- Halse, P., 1991b, Phys. Rev. C 44, 2467.
- Hamamoto, I., 1971, Nucl. Phys. A 177, 484.
- Hamamoto, I., and S. Åberg, 1984, Phys. Lett. B 145, 163.
- Hamamoto, I., and C. Magnusson, 1991, Phys. Lett. B **260**, 6.
- Hamamoto, I., and W. Nazarewicz, 1992, Phys. Lett. B 297, 25.
- Hamamoto, I., and W. Nazarewicz, 1994, Phys. Rev. C 49, 2489.
- Hamamoto, I., and C. Ronström, 1987, Phys. Lett. B 194, 6.Hamilton, W. D., A. Irbäck, and J. P. Elliott, 1984, Phys. Rev. Lett. 53, 2469.
- Haq, R. U., A. Pandey, and O. Bohigas, 1982, Phys. Rev. Lett. 48, 1086.
- Harakeh, M., and A. van der Woude, 2001, Giant Resonances: Fundamental High-Frequency Modes of Nuclear Excitation (Oxford University, Oxford).
- Harney, H. L., A. Richter, and H. A. Weidenmüller, 1986, Rev. Mod. Phys. 58, 607.
- Hartung, G., A. Richter, E. Spamer, H. Wörtche, C. Rangacharyulu, C. W. de Jager, and H. de Vries, 1989, Phys. Lett. B 221, 109.
- Hasse, R. W., and W. D. Myers, 1988, Geometrical Relationships of Macroscopic Nuclear Physics (Springer, Berlin).
- Hatada, K., K. Hayakawa, and F. Palumbo, 2005, Phys. Rev. B $\mathbf{71}$, 092402.
- Hatada, K., K. Hayakawa, and F. Palumbo, 2009a, arXiv:0909.1422 .
- Hatada, K., K. Hayakawa, and F. Palumbo, 2009b, arXiv:0905.2100 .
- Hatada, K., K. Hayakawa, and F. Palumbo, 2010, ar
Xiv:1004.2220 .
- Hecht, K. T., and A. Adler, 1969, Nucl. Phys. A 137, 129.
- de Heer, W. A., 1993, Rev. Mod. Phys. **65**, 611. de Heer, W. A., K. Selby, V. Kresin, J. Masui, M. Vollmer,
- A. Chatelain, and W. D. Knight, 1987, Phys. Rev. Lett. 59, 1805.
- Heil, R. D., H. H. Pitz, U. E. P. Berg, U. Kneissl, K. D. Hummel, G. Kilgus, D. Bohle, A. Richter, C. Wesselborg,

- and P. von Brentano, 1988, Nuc. Phys. A 476, 39.
- Heyde, K., 1989, Int. J. Mod. Phys. A 4, 2063.
- Heyde, K., and C. De Coster, 1991, Phys. Rev. C 44, R2262.
- Heyde, K., and C. De Coster, 1993, Phys. Lett. B 305, 322.
- Heyde, K., C. De Coster, and D. Ooms, 1994, Phys. Rev. C 49, 156.
- Heyde, K., C. De Coster, D. Ooms, and A. Richter, 1993, Phys. Lett. B 312, 267.
- Heyde, K., C. De Coster, A. Richter, and H. J. Wörtche, 1992, Nucl. Phys. A 549, 103.
- Heyde, K., C. De Coster, S. Rombouts, and S. J. Freeman, 1996, Nucl. Phys. A **596**, 30.
- Heyde, K., and J. Sau, 1984, Phys. Rev. C 30, 1355.
- Heyde, K., and J. Sau, 1986, Phys. Rev. C 33, 1050.
- Heyde, K., M. Waroquier, P. Van Isacker, and H. Vincx, 1977, Phys. Rev. C 16, 489.
- Hicks, S. F., C. M. Davoren, W. M. Faulkner, and J. R. Vanhoy, 1998, Phys. Rev. C 57, 2264.
- Hicks, S. F., J. R. Vanhoy, and S. W. Yates, 2008, Phys. Rev. C 78, 054320.
- Hilton, R. R., 1976, talk at International Conference on Nuclear Structure, JINR Dubna, unpublished.
- Hilton, R. R., 1995, in Proceedings of the International Conference on Perspectives of Nuclear Physics in the Late Nineties, edited by N. Dinh Dang, D. Feng, N. V. Giai, and N. D. Tu (World Scientific, Singapore), p. 86.
- Hilton, R. R., W. Höhenberger, and H. J. Mang, 1993, Phys. Rev. C 47, 602.
- Hilton, R. R., W. Höhenberger, and P. Ring, 1998, Eur. Phys. J. A 1, 257.
- Hino, M., K. Muto, and T. Oda, 1988, Phys. Rev. C **37**, 1328. Hirose, K., and N. S. Wingreen, 1999, Phys. Rev. B **59**, 4604.
- Hix, W. R., A. Mezzacappa, O. E. B. Messer, and S. W. Bruenn, 2003, J. Phys. G 29, 2523.
- Hofmann, F., C. Bäumer, A. M. van den Berg, D. Frekers, V. M. Hannen, M. N. Harakeh, M. de Huu, Y. Kalmykov, P. von Neumann-Cosel, V. Y. Ponomarev, S. Rakers, B. Reitz, et al., 2007, Phys. Rev. C 76, 014314.
- Hofmann, F., P. von Neumann-Cosel, F. Neumeyer, C. Rangacharyulu, B. Reitz, A. Richter, G. Schrieder, D. I. Sober, L. W. Fagg, and B. A. Brown, 2002, Phys. Rev. C 65, 024311.
- Holt, J. D., N. Pietralla, J. W. Holt, T. T. S. Kuo, and G. Rainovski, 2007, Phys. Rev. C 76, 034325.
- Holzwarth, D., and G. Eckart, 1977, Z. Phys. A 283, 219.
- Holzwarth, D., and G. Eckart, 1979, Nucl. Phys. A 325, 1.
- Honma, M., T. Mizusaki, and T. Otsuka, 1995, Phys. Rev. Lett. 75, 1284.
- Honma, M., T. Mizusaki, and T. Otsuka, 1996, Phys. Rev. Lett. 77, 3315.
- Honma, M., T. Otsuka, B. A. Brown, and T. Mizusaki, 2004, Phys. Rev. C 69, 034335.
- Huxel, A., P. von Brentano, J. Eberth, J. Enders, R.-D. Herzberg, P. von Neumann-Cosel, N. Nicolay, N. Pietralla, H. Prade, C. Rangacharyulu, J. Reif, A. Richter, et al., 1999, Nucl. Phys. A 645, 239.
- Huxel, N., W. Ahner, H. Diesener, P. von Neumann-Cosel, C. Rangacharyulu, A. Richter, C. Spieler, W. Ziegler, C. De Coster, and K. Heyde, 1992, Nucl. Phys. A 539, 478.
- Iachello, F., 1981, Nucl. Phys. A 358, 89c.
- Iachello, F., 1984, Phys. Rev. Lett. 53, 1427.
- Iachello, F., and A. Arima, 1987, The Interacting Boson Model (Cambridge University, New York).

- Iachello, F., and P. Van Isacker, 1991, The Interacting Boson-Fermion Model (Cambridge University, New York).
- Ichimura, M., H. Sakai, and T. Wakasa, 2006, Prog. Part. Nucl. Phys. 56, 446.
- Ikeda, A., and T. Shimano, 1993, Nucl. Phys. A 577, 573c. Iwasaki, S., and K. Hara, 1984, Phys. Lett. B 144, 9.
- Jewell, J. K., P. D. Cottle, K. W. Kemper, and L. A. Riley, 1997, Phys. Rev. C 56, 2440.
- Kalmykov, Y., T. Adachi, G. P. A. Berg, H. Fujita, K. Fujita, Y. Fujita, K. Hatanaka, J. Kamiya, K. Nakanishi, P. von Neumann-Cosel, V. Y. Ponomarev, A. Richter, et al., 2006, Phys. Rev. Lett. 96, 012502.
- Kasten, B., R. D. Heil, P. von Brentano, P. A. Butler, S. D. Hoblit, U. Kneissl, S. Lindenstruth, G. Müller, H. H. Pitz, K. W. Rose, W. Scharfe, M. Schumacher, et al., 1989, Phys. Rev. Lett. 63, 609.
- Klein, H., A. F. Lisetskiy, N. Pietralla, C. Fransen, A. Gade, and P. von Brentano, 2002, Phys. Rev. C 65, 044315.
- Kleinig, W., V. O. Nesterenkov, P.-G. Reinhard, and L. Serra, 1998, Eur. Phys. J. D 4, 343.
- Kneissl, U., N. Pietralla, and A. Zilges, 2006, J. Phys. G 32, R217.
- Kneissl, U., H. H. Pitz, and A. Zilges, 1996, Prog. Part. Nucl. Phys. **37**, 349.
- Knight, W. D., K. Clemenger, W. A. de Heer, W. A. Saunders, M. Y. Chou, and M. L. Cohen, 1984, Phys. Rev. Lett.
- Kohstall, C., D. Belic, P. von Brentano, C. Fransen, A. Gade, R.-D. Herzberg, J. Jolie, U. Kneissl, A. Linnemann, A. Nord, N. Pietralla, H. H. Pitz, et al., 2005, Phys. Rev. C **72**, 034302.
- Koskinen, M., M. Manninen, and S. M. Reimann, 1997, Phys. Rev. Lett. 79, 1389.
- Kurasawa, H., and T. Suzuki, 1984, Phys. Lett. B 144, 151. Kuyucak, S., and A. E. Stuchbery, 1995, Phys. Lett. B 348, 315.
- van der Laan, G., E. Arenholz, A. Schmehl, and D. G. Schlom, 2008, Phys. Rev. Lett. 100, 067403.
- Langanke, K., and G. Martínez-Pinedo, 2003, Rev. Mod. Phys. 75, 819.
- Langanke, K., G. Martínez-Pinedo, B. Müller, H.-T. Janka, A. Marek, W. R. Hix, A. Juodagalvis, and J. M. Sampaio, 2008, Phys. Rev. Lett. **100**, 011101.
- Langanke, K., G. Martínez-Pinedo, P. von Neumann-Cosel, and A. Richter, 2004, Phys. Rev. Lett. 93, 202501.
- Laszewski, R. M., R. Alarcon, D. S. Dale, and S. D. Hoblit, 1988, Phys. Rev. Lett. **61**, 1710.
- Lehmann, H., A. Nord, A. E. de Almeida Pinto, O. Beck, J. Besserer, P. von Brentano, S. Drissi, T. Eckert, R.-D. Herzberg, D. Jäger, J. Jolie, U. Kneissl, et al., 1999, Phys. Rev. C 60, 024308.
- Lesher, S. R., C. J. McKay, M. Mynk, D. Bandyopadhyay, N. Boukharouba, C. Fransen, J. N. Orce, M. T. McEllistrem, and S. W. Yates, 2007, Phys. Rev. C 75, 034318.
- Li, T. C., N. Pietralla, C. Fransen, H. von Garrel, U. Kneissl, C. Kohstall, A. Linnemann, H. H. Pitz, G. Rainovski, A. Richter, M. Scheck, F. Stedile, et al., 2005, Phys. Rev. C 71, 044318.
- Li, T. C., N. Pietralla, A. P. Tonchev, M. W. Ahmed, T. Ahn, C. Angell, M. A. Blackston, A. Costin, K. J. Keeter, J. Li, A. Lisetskiy, S. Mikhailov, et al., 2006, Phys. Rev. C 73, 054306.
- Lieb, K. P., H. G. Börner, M. S. Dewey, J. Jolie, S. J. Robinson, S. Ulbig, and C. Winter, 1988, Phys. Lett. B 215,

- 50
- Linnemann, A., P. von Brentano, J. Eberth, J. Enders, A. Fitzler, C. Fransen, E. Guliyev, R.-D. Herzberg, L. Käubler, A. A. Kuliev, P. von Neumann-Cosel, N. Pietralla, et al., 2003, Phys. Lett. B **554**, 15.
- Linnemann, A., C. Fransen, M. Gorska, J. Jolie, U. Kneissl, P. Knoch, D. Mücher, H. H. Pitz, M. Scheck, C. Scholl, and P. von Brentano, 2005, Phys. Rev. C 72, 064323.
- Linnemann, A., C. Fransen, J. Jolie, U. Kneissl, P. Knoch, C. Kohstall, D. Mücher, H. H. Pitz, M. Scheck, C. Scholl, F. Stedile, P. von Brentano, et al., 2007, Phys. Rev. C 75,
- Lipparini, E., 2003, Modern Many-Particle Physics: Atomic Gases, Quantum Dots and Quantum Fluids (World Scientific, Singapore).
- Lipparini, E., and A. Richter, 1984, Phys. Lett. B 144, 13.
- Lipparini, E., and S. Stringari, 1983, Phys. Lett. B 130, 139. Lipparini, E., and S. Stringari, 1989a, Phys. Rep. 175, 103.
- Lipparini, E., and S. Stringari, 1989b, Phys. Rev. Lett. 63, 570.
- Lisantti, J., J. R. Tinsley, D. M. Drake, I. Bergqvist, L. W. Swenson, D. K. McDaniels, F. E. Bertrand, E. E. Gross, D. J. Horen, and T. P. Sjoreen, 1984, Phys. Lett. B 147,
- Lisetskiy, A. F., E. Caurier, K. Langanke, G. Martínez-Pinedo, P. von Neumann-Cosel, F. Nowacki, and A. Richter, 2007, Nucl. Phys. A 789, 114.
- Lisetskiy, A. F., N. Pietralla, C. Fransen, R. V. Jolos, and P. von Brentano, 2000, Nucl. Phys. A 677, 100.
- Liu, H., and L. Zamick, 1987a, Nucl. Phys. A 467, 29.
- Liu, H., and L. Zamick, 1987b, Phys. Rev. C 36, 2064.
- Liu, H., and L. Zamick, 1987c, Phys. Rev. C 36, 2057.
- Lo Iudice, N., 1988, Phys. Rev. C 38, 2895.
- Lo Iudice, N., 1996a, Nucl. Phys. A 605, 61.
- Lo Iudice, N., 1996b, Phys. Rev. C 53, 2171.
- Lo Iudice, N., 1997, Phys. Part. Nucl. 28, 556.
- Lo Iudice, N., 1998, Phys. Rev. C 57, 1246.
- Lo Iudice, N., 2000, Rivista Nuovo Cimento 23, 1.
- Lo Iudice, N., and F. Palumbo, 1978, Phys. Rev. Lett. 41, 1532.
- Lo Iudice, N., and F. Palumbo, 1979, Nucl. Phys. A 326, 193.
- Lo Iudice, N., and A. Richter, 1989, Phys. Lett. B 228, 91. Lo Iudice, N., and A. Richter, 1993, Phys. Lett. B **304**, 193.
- Lo Iudice, N., and C. Stoyanov, 2000, Phys. Rev. C 62,
- 047302.
- Lo Iudice, N., and C. Stoyanov, 2002, Phys. Rev. C 65, 064304.
- Lo Iudice, N., and C. Stoyanov, 2004, Phys. Rev. C 69, 044312.
- Lo Iudice, N., and C. Stoyanov, 2006, Phys. Rev. C 73, 037305.
- Lo Iudice, N., C. Stoyanov, and N. Pietralla, 2009, Phys. Rev. C **80**, 024311.
- Lo Iudice, N., C. Stoyanov, and D. Tarpanov, 2008, Phys. Rev. C 77, 044310.
- Löbner, K. E. G., M. Vetter, and V. Hönig, 1970, Nucl. Data Tables A 7, 495.
- Lüttge, C., C. Hofmann, J. Horn, F. Neumeyer, A. Richter, G. Schrieder, E. Spamer, A. Stiller, D. I. Sober, S. K. Matthews, and L. W. Fagg, 1995, Nucl. Instrum. Methods Phys. Res. A **366**, 325.
- Lüttge, C., P. von Neumann-Cosel, F. Neumeyer, C. Rangacharyulu, A. Richter, G. Schrieder, E. Spamer, D. I. Sober, S. K. Matthews, and B. A. Brown, 1996a, Phys.

- Rev. C 53, 127.
- Lüttge, C., P. von Neumann-Cosel, F. Neumeyer, and A. Richter, 1996b, Nucl. Phys. A 606, 183.
- Maragò, O., G. Hechenblaikner, E. Hodby, and C. Foot, 2001, Phys. Rev. Lett. 86, 3938.
- Maragò, O. M., S. A. Hopkins, J. Arlt, E. Hodby, G. Hechenblaikner, and C. J. Foot, 2000, Phys. Rev. Lett. 84, 2056.
- Margraf, J., A. Degener, H. Friedrichs, R. D. Heil, A. Jung, U. Kneissl, S. Lindenstruth, H. H. Pitz, H. Schacht, U. Seemann, R. Stock, C. Wesselborg, et al., 1990, Phys. Rev. C 42, 771.
- Margraf, J., T. Eckert, M. Rittner, I. Bauske, O. Beck, U. Kneissl, H. Maser, H. H. Pitz, A. Schiller, P. von Brentano, R. Fischer, R.-D. Herzberg, et al., 1995, Phys. Rev. C 52, 2429.
- Margraf, J., R. D. Heil, U. Kneissl, U. Maier, H. H. Pitz, H. Friedrichs, S. Lindenstruth, B. Schlitt, C. Wesselborg, P. von Brentano, R.-D. Herzberg, and j. . P. v. . . p. . . n. . y. . . d. . P. p. . A. Zilges, A., ????
- Martínez-Pinedo, G., A. Poves, E. Caurier, and A. P. Zuker, 1996, Phys. Rev. C 53, R2602.
- McCullen, J. D., B. F. Bayman, and L. Zamick, 1964, Phys. Rev. 134, B515.
- Minguzzi, A., and M. P. Tosi, 2001, Phys. Rev. A 63, 023609.Mizusaki, T., M. Honma, and T. Otsuka, 1996, Phys. Rev. C 53, 2786.
- Mizusaki, T., T. Otsuka, and M. Sugita, 1991, Phys. Rev. C 44, R1277.
- Mizusaki, T., T. Otsuka, Y. Utsuno, M. Honma, and T. Sebe, 1999, Phys. Rev. C 59, R1846.
- Modugno, G., M. Modugno, F. Riboli, G. Roati, and M. Inguscio, 2002, Phys. Rev. Lett. 89, 190404.
- Mukhopadhyay, S., M. Scheck, B. Crider, S. N. Choudry, E. Elhami, E. Peters, M. T. McEllistrem, J. N. Orce, and S. W. Yates, 2008, Phys. Rev. C 78, 034317.
- Nakatsukasa, T., K. Matsuyanagi, I. Hamamoto, and W. Nazarewicz, 1994, Nucl. Phys. A 573, 333.
- Nesterenko, V. O., W. Kleinig, V. V. Gudkov, N. Lo Iudice, and J. Kvasil, 1997, Phys. Rev. A 56, 607.
- Nesterenko, V. O., W. Kleinig, F. F. de Souza Cruz, and N. Lo Iudice, 1999, Phys. Rev. Lett. 83, 57.
- Nesterenko, V. O., J. R. Marinelli, F. F. de Souza Cruz, W. Kleinig, and P.-G. Reinhard, 2000, Phys. Rev. Lett. 85, 3141.
- von Neumann-Cosel, P., 1997, Prog. Part. Nucl. Phys. 38, 213
- von Neumann-Cosel, P., T. Adachi, C. A. Bertulani, J. Carter, M. Dozono, H. Fujita, K. Fujita, Y. Fujita, H. Hashimoto, K. Hatanaka, M. Itoh, Y. Kalmykov, et al., 2009, in Proceedings of the 13th International Symposium on Capture Gamma-Ray Spectroscopy and Related Topics, edited by J. Jolie, A. Zilges, N. Warr, and A. Blazhev (AIP, Vol. 1090), p. 404.
- von Neumann-Cosel, P., and J. N. Ginocchio, 2000, Phys. Rev. C 62, 014308.
- von Neumann-Cosel, P., J. N. Ginocchio, H. Bauer, and A. Richter, 1995, Phys. Rev. Lett. **75**, 4178.
- von Neumann-Cosel, P., H.-D. Gräf, U. Krämer, A. Richter, and E. Spamer, 2000, Nucl. Phys. A 669, 3.
- von Neumann-Cosel, P., F. Neumeyer, S. Nishizaki, V. Y. Ponomarev, C. Rangacharyulu, B. Reitz, A. Richter, G. Schrieder, D. I. Sober, T. Waindzoch, and J. Wambach, 1999, Phys. Rev. Lett. 82, 1105.
- von Neumann-Cosel, P., A. Poves, J. Retamosa, and

- A. Richter, 1998, Phys. Lett. B 443, 1.
- von Neumann-Cosel, P., A. Richter, Y. Fujita, and B. D. Anderson, 1997, Phys. Rev. C 55, 532.
- Nojarov, R., Z. Bochnacki, and A. Faessler, 1986, Z. Phys. A 324, 289.
- Nojarov, R., and A. Faessler, 1990, Z. Phys. A 336, 151.
- Nojarov, R., A. Faessler, and O. Civitarese, 1987, Phys. Lett. B 183, 122.
- Nojarov, R., A. Faessler, and M. Dingfelder, 1995, Phys. Rev. C 51, 2449.
- Nojarov, R., A. Faessler, and P. O. Lipas, 1991, Nucl. Phys. A 533, 381.
- Nojarov, R., A. Faessler, and P. O. Lipas, 1992, Nucl. Phys. A 537, 707 (erratum).
- Nord, A., J. Enders, A. E. de Almeida Pinto, D. Belic, P. von Brentano, C. Fransen, U. Kneissl, C. Kohstall, A. Linneman, P. von Neumann-Cosel, N. Pietralla, H. H. Pitz, et al., 2003, Phys. Rev. C 67, 034307.
- Nord, A., A. Schiller, T. Eckert, O. Beck, J. Besserer, P. von Brentano, R. Fischer, R.-D. Herzberg, D. Jäger, U. Kneissl, J. Margraf, H. Maser, et al., 1996, Phys. Rev. C 54, 2287.
- Oda, M. H., T., and K. Muto, 1987, Phys. Lett. B **190**, 14.
- Otsuka, M. H., T., and T. Mizusaki, 1999, J. Phys. G 25, 699.
- Otsuka, T., 1990, Nucl. Phys. A 507, 129c.
- Otsuka, T., 1992, Hyperfine Int. 74, 93.
- Otsuka, T., and J. N. Ginocchio, 1985, Phys. Rev. Lett. **54**, 777.
- Otsuka, T., M. Honma, and T. Mizusaki, 1998, Phys. Rev. Lett. 81, 1588.
- Otsuka, T., and K.-H. Kim, 1994, Phys. Rev. C 50, R1768.
- Otsuka, T., X. W. Pan, and A. Arima, 1990, Phys. Lett. B **247**, 191.
- Otten, E., 1989, Treatise on Heavy-Ion Science, Vol. 8 (Plenum, New York).
- Petermann, I., K. Langanke, G. Martínez-Pinedo, P. von Neumann-Cosel, F. Nowacki, and A. Richter, 2010, Phys. Rev. C 81, 014308.
- Pietralla, N., C. J. Barton III, R. Krücken, C. W. Beausang, M. A. Caprio, R. F. Casten, J. R. Cooper, A. A. Hecht, H. Newman, J. R. Novak, and N. V. Zamfir, 2001a, Phys. Rev. C 64, 031301(R).
- Pietralla, N., Z. Berant, V. N. Litvinenko, S. Hartman, F. F. Mikhailov, I. V. Pinayev, G. Swift, M. W. Ahmed, J. H. Kelley, S. O. Nelson, R. Prior, K. Sabourov, et al., 2001b, Phys. Rev. Lett. 88, 012502.
- Pietralla, N., P. von Brentano, A. Gelberg, T. Otsuka, A. Richter, N. Smirnova, and I. Wiedenhöver, 1998a, Phys. Rev. C 58, 191.
- Pietralla, N., P. von Brentano, R.-D. Herzberg, U. Kneissl, N. Lo Iudice, H. Maser, H. H. Pitz, and A. Zilges, 1998b, Phys. Rev. C 58, 184.
- Pietralla, N., P. von Brentano, and A. F. Lisetsky, 2008, Prog. Part. Nucl. Phys. 60, 225.
- Pietralla, N., C. Fransen, D. Belic, P. von Brentano, C. Frießner, U. Kneissl, A. Linnemann, A. Nord, H. H. Pitz, T. Otsuka, I. Schneider, V. Werner, et al., 1999, Phys. Rev. Lett. 83, 1303.
- Pietralla, N., C. Fransen, P. von Brentano, A. Dewald, A. Fitzler, C. Frießner, and J. Gableske, 2000, Phys. Rev. Lett. 84, 3775.
- Pitaevskii, L., and S. Stringari, 2003, Bose-Einstein Condensation, International Series of Monographs on Physics, Vol. 116 (Oxford Science, Clarendon).

- Poves, A., J. Retamosa, and E. Moya de Guerra, 1989, Phys. Rev. C **39**, 1639.
- Poves, A., J. Sánchez-Solano, E. Caurier, and F. Nowacki, 2001, Nucl. Phys. A 694, 157.
- Puente, A., and L. Serra, 1999, Phys. Rev. Lett. 83, 3266.
- Raduta, A. A., and D. S. Delion, 1990, Nucl. Phys. A **513**, 11.
- Raduta, A. A., and N. Lo Iudice, 1989, Z. Phys. A 334, 403.
 Rainovski, G., N. Pietralla, T. Ahn, C. J. Lister, R. V. F.
 Janssens, M. P. Carpenter, S. Zhu, and C. J. Barton III, 2006, Phys. Rev. Lett. 96, 122501.
- Raman, S., L. W. Fagg, and R. S. Hicks, 1991 (World Scientific, Singapore), International Review of Nuclear Physics, Vol. 7, p. 355.
- Rangacharyulu, C., A. Richter, H. J. Wörtche, W. Ziegler, and R. F. Casten, 1991, Phys. Rev. C 43, R949.
- Reitz, B., A. M. van den Berg, D. Frekers, F. Hofmann, M. de Huu, Y. Kalmykov, H. Lenske, P. von Neumann-Cosel, V. Y. Ponomarev, S. Rakers, A. Richter, G. Schrieder, et al., 2002, Phys. Lett. B 532, 179.
- Reitz, B., F. Hofmann, P. von Neumann-Cosel, F. Neumeyer, C. Rangacharyulu, A. Richter, G. Schrieder, D. I. Sober, and B. A. Brown, 1999, Phys. Rev. Lett. 82, 291.
- Retamosa, J., J. M. Udias, A. Poves, and E. Moya de Guerra, 1990, Nucl. Phys. A 511, 221.
- RIA, 2003, The Intellectual Challenges of RIA: A White Paper from the RIA Users Community (http://www.orau.org/ria/pdf/intell.pdf).
- RIA, 2006, Isotope Science Facility at Michigan State University (http://www.nscl.msu.edu/future/nsclwhitepaper2006).
- Richter, A., 1983, in *Proceedings of the International Conference on Nuclear Physics, Vol. 2*, edited by P. Blasi and R. A. Ricci (Tipografica Compositori, Bologna), p. 189.
- Richter, A., 1985, Prog. Part. Nucl. Phys. 13, 1.
- Richter, A., 1990, Nucl. Phys. A 507, 99c.
- Richter, A., 1991, Nucl. Phys. A 522, 139c.
- Richter, A., 1993, Nucl. Phys. A 553, 417c.
- Richter, A., 1994, in Proceedings of the 4th International Spring Seminar on Nuclear Physics 'The Building Blocks of Nuclear Structure', edited by A. Covello (World Scientific, Singapore), p. 335.
- Richter, A., 1995, Prog. Part. Nucl. Phys. 34, 261.
- Richter, A., and W. Knüpfer, 1980, in Electron and Pion Interactions with Nuclei at Intermediate Energies, edited by W. Bertozzi, S. Costa, and C. Schaerf (Harwood Academic, Newark, New Jersey), p. 241.
- Richter, A., A. Weiss, O. Häusser, and B. A. Brown, 1990, Phys. Rev. Lett. 65, 2519.
- Ring, P., and P. Schuck, 1980, *The Nuclear Many-Body Problem* (Springer, Heidelberg).
- Rohozinski, S. G., and W. Greiner, 1985, Z. Phys. A **322**, 271.
- Rompf, D., T. Beuschel, J. P. Draayer, W. Scheid, and J. G. Hirsch, 1998, Phys. Rev. C 57, 1703.
- Rowe, D., 1970, *Nuclear Collective Motion* (Methuen and Co., London).
- Sambataro, M., and A. Dieperink, 1981, Phys. Lett. B **107**, 249.
- Sambataro, M., O. Scholten, A. E. L. Dieperink, and G. Picciotto, 1984, Nucl. Phys. A 423, 333.
- Sarriguren, P., E. Moya de Guerra, and R. Nojarov, 1996, Phys. Rev. C 54, 690.
- Sarriguren, P., E. M. de Guerra, R. Nojarov, and A. Faessler,

- 1993, J. Phys. G 19, 291.
- Sarriguren, P., E. M. de Guerra, R. Nojarov, and A. Faessler, 1994, J. Phys. G **20**, 315.
- Savran, D., S. Müller, A. Zilges, M. Babilon, M. W. Ahmed, J. H. Kelley, A. Tonchev, W. Tornow, H. R. Weller, N. Pietralla, J. Li, I. V. Pinayev, et al., 2005, Phys. Rev. C 71, 034304.
- Scheck, M., S. N. Choudry, E. Elhami, M. T. McEllistrem, S. Mukhopadhyay, J. N. Orce, and S. W. Yates, 2008, Phys. Rev. C 78, 034302.
- Scheck, M., H. von Garrel, N. Tsoneva, D. Belic, P. von Brentano, C. Fransen, A. Gade, J. Jolie, U. Kneissl, C. Kohstall, A. Linnemann, A. Nord, et al., 2004, Phys. Rev. C 70, 044319.
- Schlegel, C., P. von Neumann-Cosel, A. Richter, and P. Van Isacker, 1996, Phys. Lett. B **375**, 21.
- Scholten, O., A. E. L. Dieperink, K. Heyde, and P. Van Isacker, 1984, Phys. Lett. B **149**, 279.
- Scholten, O., K. Heyde, and P. Van Isacker, 1985a, Phys. Rev. Lett. 55, 1866.
- Scholten, O., K. Heyde, P. Van Isacker, J. Jolie, J. Moreau, M. Waroquier, and J. Sau, 1985b, Nucl. Phys. A 438, 41.
- Scholtz, F. G., R. Nojarov, and A. Faessler, 1989, Phys. Rev. Lett. 63, 1356.
- Schüller, C., G. Biese, K. Keller, C. Steinebach, D. Heitmann, P. Grambow, and K. Eberl, 1996, Phys. Rev. B **54**, R17304.
- Schwengner, R., G. Rusev, N. Benouaret, R. Beyer, M. Erhard, E. Grosse, A. R. Junghans, J. Klug, K. Kosev, L. Kostov, C. Nair, N. Nankov, et al., 2007, Phys. Rev. C 76, 034321.
- Schwengner, R., G. Winter, W. Schauer, M. Grinberg, F. Becker, P. von Brentano, J. Eberth, J. Enders, T. von Egidy, R.-D. Herzberg, N. Huxel, L. Käubler, et al., 1997a, Nucl. Phys. A 620, 277.
- Schwengner, R., G. Winter, W. Schauer, M. Grinberg, F. Becker, P. von Brentano, J. Eberth, J. Enders, T. von Egidy, R.-D. Herzberg, N. Huxel, L. Käubler, et al., 1997b, Nucl. Phys. A 624, 776(E).
- Serra, L., A. Puente, and E. Lipparini, 1999, Phys. Rev. B 60, R13966.
- Shevchenko, A., O. Burda, J. Carter, G. R. J. Cooper, R. W. Fearick, S. V. Förtsch, H. Fujita, Y. Fujita, Y. Kalmykov, D. Lacroix, J. J. Lawrie, P. von Neumann-Cosel, et al., 2009, Phys. Rev. C 79, 044305.
- Shevchenko, A., J. Carter, G. R. J. Cooper, R. W. Fearick, Y. Kalmykov, P. von Neumann-Cosel, V. Y. Ponomarev, A. Richter, I. Usman, and J. Wambach, 2008, Phys. Rev. C 77, 024302.
- Shevchenko, A., J. Carter, R. W. Fearick, S. V. Förtsch, H. Fujita, Y. Fujita, Y. Kalmykov, D. Lacroix, J. J. Lawrie, P. von Neumann-Cosel, R. Neveling, V. Y. Ponomarev, et al., 2004, Phys. Rev. Lett. 93, 122501.
- Shimizu, N., T. Otsuka, T. Mizusaki, and M. Honma, 2001, Phys. Rev. Lett. **86**, 1171.
- Shriner Jr., J. F., G. E. Mitchell, and T. von Egidy, 1991, Z. Phys. A 338, 309.
- Sieja, K., G. Martínez-Pinedo, L. Coquard, and N. Pietralla, 2009, Phys. Rev. C 80, 054311.
- Sikorski, C., and U. Merkt, 1989, Phys. Rev. Lett. **62**, 2164.Smith, B. H., X.-W. Pan, D. H. Feng, and M. Guidry, 1995, Phys. Rev. Lett. **75**, 3086.
- Sober, D. I., B. C. Metsch, W. Knüpfer, G. Eulenberg, G. Küchler, A. Richter, E. Spamer, and W. Steffen, 1985, Phys. Rev. C **31**, 2054.

- Soloviev, V. G., 1992, Theory of the Atomic Nucleus: Quasiparticles and Phonons (IOP, Bristol).
- Soloviev, V. G., A. V. Sushkov, and N. Y. Shirikova, 1996, Phys. Rev. C 53, 1022.
- Soloviev, V. G., A. V. Sushkov, and N. Y. Shirikova, 1997a, Prog. Part. Nucl. Phys. 38, 53.
- Soloviev, V. G., A. V. Sushkov, and N. Y. Shirikova, 1997b, Phys. Rev. C 56, 2528.
- Soloviev, V. G., A. V. Sushkov, N. Y. Shirikova, and N. Lo Iudice, 1997c. Nucl. Phys. A 613, 45.
- Steffen, W., H.-D. Gräf, W. Gross, D. Meuer, A. Richter, E. Spamer, O. Titze, and W. Knüpfer, 1980, Phys. Lett. B 95, 23.
- Strenz, R., U. Bockelmann, F. Hirler, G. Abstreiter, G. Böhm, and G. Weimann, 1994, Phys. Rev. Lett. 73, 3022.
- Sugawara-Tanabe, K., and A. Arima, 1989, Phys. Lett. B **229**, 327.
- Suzuki, T., and D. Rowe, 1977, Nucl. Phys. A 289, 461.
- Takayanagi, K., K. Shimizu, and A. Arima, 1988, Nucl. Phys. A 481, 313.
- Tamii, A., Y. Fujita, H. Matsubara, T. Adachi, J. Carter,
 M. Dozono, H. Fujita, K. Fujita, H. Hashimoto,
 K. Hatanaka, T. Itahashi, M. Itoh, et al., 2009, Nuclear
 Instrum. and Methods in Phys. Research A 605, 326.
- Towner, I. S., 1987, Phys. Rep. 155, 263.
- Utsuno, Y., T. Otsuka, T. Mizusaki, and M. Honma, 1999, Phys. Rev. C 60, 054315.
- Van Isacker, P., and A. Frank, 1989, Phys. Lett. B 225, 1.
- Van Isacker, P., K. Heyde, J. Jolie, and A. Sevrin, 1986, Ann. Phys. (N.Y.) 171, 253.
- Van Isacker, P., M. A. Nagarajan, and D. D. Warner, 1992, Phys. Rev. C 45, R13.
- Vanhoy, J. R., J. M. Anthony, B. M. Haas, B. H. Benedict, B. T. Meehan, S. F. Hicks, C. M. Davoren, and C. L. Lundstedt, 1995, Phys. Rev. C 52, 2387.
- Viñas, X., R. Roth, P. Schuck, and J. Wambach, 2001, Phys. Rev. A 64, 055601.
- Warner, D. D., and P. Van Isacker, 1997, Phys. Lett. B **395**, 145.
- Weidenmüller, H. A., and G. E. Mitchell, 2009, Rev. Mod. Phys. 81, 539.
- Weller, H. R., M. W. Ahmed, H. Gao, W. Tornow, Y. K. Wu, M. Gai, and R. Miskimen, 2009, Prog. Part. Nucl. Phys. 62, 257.
- Werner, V., D. Belic, P. von Brentano, C. Fransen, A. Gade, H. von Garrel, J. Jolie, U. Kneissl, C. Kohstall, A. Linneman, A. F. Lisetsky, N. Pietralla, et al., 2002, Phys. Lett. B 550, 140.

- Werner, V., N. Benczer-Koller, G. Kumbartzki, J. D. Holt, P. Boutachkov, E. Stefanova, M. Perry, N. Pietralla, H. Ai, K. Aleksandrova, G. Anderson, R. B. Cakirli, et al., 2008, Phys. Rev. C 78, 031301(R).
- Wesselborg, C., P. von Brentano, K. O. Zell, R. D. Heil, H. H. Pitz, U. E. P. Berg, U. Kneissl, S. Lindenstruth, U. Seemann, and R. Stock, 1988, Phys. Lett. B 207, 22.
- Wesselborg, C., K. Schiffer, K. O. Zell, P. von Brentano, D. Bohle, A. Richter, G. P. A. Berg, B. Brinkmöller, J. G. M. Römer, F. Osterfeld, and M. Yabe, 1986, Z. Phys. A 323, 485.
- Wiedenhöver, I., A. Gelberg, T. Otsuka, N. Pietralla, J. Gableske, A. Dewald, and P. von Brentano, 1997, Phys. Rev. C 56, R2354.
- Williams, E., R. J. Casperson, V. Werner, H. Ai, P. Boutachkov, M. Chamberlain, G. Gürdal, A. Heinz, E. A. McCutchan, J. Qian, and R. Winkler, 2009, Phys. Rev. C 80, 054309.
- Willis, A., M. Morlet, N. Marty, C. Djalali, D. Bohle, H. Diesener, A. Richter, and H. Stein, 1989, Nuclear Physics A 499, 367.
- Wolf, A., R. F. Casten, and D. D. Warner, 1987, Phys. Lett. B **190**, 19.
- Wörtche, H. J., 1994, Doctoral thesis, Technische Universität Darmstadt.
- Yevetska, O., J. Enders, M. Fritzsche, P. von Neumann-Cosel, S. Oberstedt, A. Richter, D. Savran, and K. Sonnabend, 2010, Phys. Rev. C, (in press).
- Zamick, L., 1985, Phys. Rev. C 31, 1955.
- Zamick, L., 1986a, Phys. Lett. B 167, 1.
- Zamick, L., 1986b, Phys. Rev. C 33, 691.
- Zamick, L., and D. C. Zheng, 1992, Phys. Rev. C 46, 2106.
- Zawischa, D., 1998, J. Phys. G 24, 683.
- Zawischa, D., M. Macfarlane, and J. Speth, 1990, Phys. Rev. C 42, 1461.
- Zawischa, D., and J. Speth, 1990, Phys. Lett. B 252, 4.
- Zawischa, D., and J. Speth, 1994, Nucl. Phys. A 569, 343.
- Ziegler, W., N. Huxel, P. von Neumann-Cosel, C. Rangacharyulu, A. Richter, C. Spieler, C. De Coster, and K. Heyde, 1993, Nucl. Phys. A564, 366.
- Ziegler, W., C. Rangacharyulu, A. Richter, and C. Spieler, 1990, Phys. Rev. Lett. 65, 2515.
- Zilges, A., P. von Brentano, R.-D. Herzberg, U. Kneissl, J. Margraf, and H. H. Pitz, 1996, Nucl. Phys. A 599, 147c.
- Zilges, A., P. von Brentano, A. Richter, R. D. Heil, U. Kneissl, H. H. Pitz, and C. Wesselborg, 1990, Phys. Rev. C 42, 1945.